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## Strong one-neutron emission from two-neutron unbound states in $\beta$ decays of the *r*-process nuclei <sup>86,87</sup>Ga

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performed. This model framework reproduces the experimental results. The shell model alone predicts  $P_{2n}$  significantly larger than  $P_{1n}$  for the <sup>87</sup>Ga decay, and it is necessary to invoke a statistical description to

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successfully explain the observation that  $P_{1n} > P_{2n}$ . Our new results demonstrate the relevance and importance of a statistical description of neutron emission for the prediction of the decay properties of multineutron emitters and that it must be included in the *r*-process modeling.

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Delayed neutron emission after  $\beta$  decay is found in neutron-rich nuclei where the energy window of the  $\beta^-$  decay  $(Q_\beta)$  is high enough to populate excited states of the daughter nucleus above the neutron separation energy  $(S_n)$ . It was first observed in 1939 [1]. The rare process of  $\beta$ -delayed twoneutron  $(\beta 2n)$  emission was first observed in 1979 [2] in <sup>11</sup>Li. In very neutron-rich nuclei where the  $Q_\beta$  is larger than the two-neutron separation energy  $(S_{2n})$ , delayed multineutron emissions may occur.

 $\beta$ -delayed neutron emission is expected to be a prevalent decay mode for thousands of neutron-rich nuclei, many of which will be accessible in new generation radioactiveion-beam facilities. Therefore, studies of this decay mode will become a major focus of their experimental program. Quantitative understanding of the neutron-emission process is required for planning future experimental activities aimed to provide data for nuclear structure or astrophysics. In particular with increasing decay energies and decreasing neutron separation energies when neutron-rich nuclei become available, more complex multineutron emission is expected to dominate their decays based on simple phase-space arguments. In this work, we show new data in the boundary region where  $\beta 2n$ becomes important, compare them with the nuclear models, and achieve good agreement between the experiment and predictions. We discovered that contrary to expectations, multineutron emission is a significant but not the main decay mode in decays of exotic isotopes of gallium. The observed effect can be explained in the framework of a statistical model [3] which assumes that particle and  $\gamma$ -ray emission after  $\beta$ decay occurs from the compound nucleus.

All of the neutron-rich nuclei on the r-process path are either one or multineutron  $\beta$ -delayed precursors. Delayed neutron emission shapes the final abundance pattern due to the changes of the isotopic population by modifying the decay path back to stability and by contributing significantly to the neutron flux after freeze-out. However, experimental data which enable the evaluation of the role of multineutron emission for the *r*-process nuclei, are almost nonexistent. The only observation of strong two-neutron emission ( $P_{2n} > 1\%$ ) in heavier nuclei, in the region relevant to the r-process nucleosynthesis, was achieved very recently for <sup>86</sup>Ga with  $P_{2n} =$ 20(10)% by Miernik *et al.* [4]. Even with the capabilities of the new generation radioactive beam facilities, the relevant multineutron emitters are very difficult to measure and the *r*-process modeling will continue to rely on predictions by nuclear theories. Therefore, new data points are of critical importance. The predictions for the <sup>86</sup>Ga [5–9] range between 21% and 56% for  $P_{1n}$  and 7% and 44% for  $P_{2n}$  and it is difficult to judge their reliability based on a single data point, particularly when  $\beta 3n$  or  $\beta 4n$  decay becomes significant.

The neutron emission probability is proportional to the integrated population of states in the available energy window  $Q_{\beta} - S_n$  and is related directly to the  $\beta$ -decay strength

function for single-neutron emitters when the competing  $\gamma$  decay is negligible. However, the neutron emission probability and decay strength decouple when two-neutron emission becomes energetically allowed; in such a case, competition between the 1*n* and 2*n* channels must be included. Very little is known about the role and sensitivity to nuclear structure for  $\beta$ -delayed neutron emitters. Often, predictions of the neutron emission probabilities are based on a simplified *cutoff* model that neglects  $\gamma$ -ray emission and assumes that only the higher multiplicity neutron emission prevails in the energy regions open to multiple neutron-emission channels [5–7]. In order to tackle the 1*n*/2*n* competition, Mumpower *et al.* [10] implemented the Hauser-Feshbach (HF) statistical model of particle and  $\gamma$ -ray emission from compound nucleus [3] with the quasiparticle random-phase (QRPA) model strength function.

This model (QRPA-HF) predicts  $P_{1n}$  or  $P_{2n}$  by following statistical decays of both the delayed- $\gamma$  and neutron emission one by one until all the excitation energy is exhausted. The particle and  $\gamma$  emission process in the statistical model is not sensitive to the details of the nuclear structure of the involved nuclei.

Substantial  $P_{1n}$  values are reported for neutron-rich N >50 Ga (Z = 31) isotopes, 21.2(9)% in <sup>82</sup>Ga (weighted average of [11–13]), 62.8(25)% in <sup>83</sup>Ga [14], 74(14)% in <sup>84</sup>Ga [15] [superseded by 40(7)% [16] and 53(20)% [17]], 70(5)% for <sup>85</sup>Ga [18], and 60(10)% for <sup>86</sup>Ga [4]. The strong delayedneutron emission branching ratios are due to their large  $\beta$ decay energy window. A detailed neutron emission study was done by Madurga et al. [16] for <sup>83,84</sup>Ga. Observation of high-energy neutrons emitted after  $\beta$  decay was interpreted as a signature of the shell structure effects dominating the  $\beta$ -decay process [16]. Madurga *et al.* [16] compared existing data and calculations for half-lives and branching ratios of  $^{82-87}$ Ga decays, based on the details of the  $\beta$ -strength distribution, but no statistical model treatment was included to make predictions for  $P_{xn}$ . Good agreement between the prediction by the shell model and experimental data was achieved for  $P_{1n}$  and half-lives by choosing a 50% quenching factor for the Gamow-Teller strength B(GT). This quenching was deduced from the experimental neutron spectrum and by adding a contribution from forbidden transitions based on experiments. The fact that the nuclear half-lives for Ga isotopes are relatively long despite the large  $Q_{\beta}$  values reflect the concentration of the B(GT) in highly excited neutronemitting states in Ge isotopes. The model by Möller et al. [5,9] uses QRPA calculations to make predictions of  $P_{1n,2n}$  values and the model in principle reflects very similar shell-structure effects. The details of the model will result in a different strength distribution and delayed neutron emission probabilities. Most notably, the effects of deformation are included.

The focus of the present work is to study the delayed neutron emission for nuclei expected to be 2n precursors such

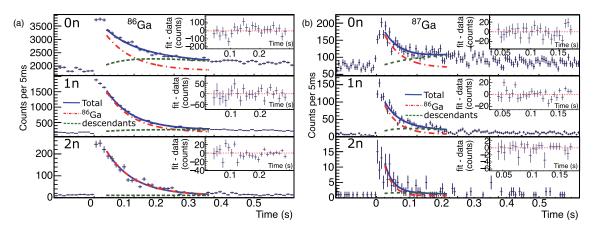


FIG. 1. Decay curves and residuals gated by neutron multiplicity 0, 1, and 2 for (a) <sup>86</sup>Ga and (b) <sup>87</sup>Ga obtained in this work.

as <sup>84–87</sup>Ga. In the cases of <sup>86,87</sup>Ga, both the shell model and QRPA predict that the majority of the B(GT) strength and resulting  $\beta$ -decay feeding within the  $Q_{\beta}$  window is concentrated above the two-neutron separation energy. The predicted decay mode of <sup>87</sup>Ga based on the shell model plus cutoff model results in two-neutron emission ( $P_{2n} = 69\%$ ) dominating over single-neutron emission. For the <sup>84,85</sup>Ga, smaller but significant  $P_{2n} \approx 10\%$  values are predicted by the cutoff models.

We studied neutron-rich Ga isotopes by means of  $\beta$ neutron- $\gamma$  spectroscopy at the Radioactive-Isotope Beam Factory (RIBF) at the RIKEN Nishina Center. Neutron-rich nuclei were produced by in-flight fission of a 345 MeV/nucleon <sup>238</sup>U<sup>86+</sup> beam induced on a 4-mm-thick<sup>9</sup>Be production target. Fission fragments were separated and identified in the BigRIPS in-flight separator [19] on an event-by-event basis [20]. A total of  $7 \times 10^4$  and  $6 \times 10^3$  ions of <sup>86</sup>Ga and <sup>87</sup>Ga, respectively, were transported to the final focal plane for decay measurement.

The secondary ions of interest were implanted into active stoppers made of double-sided silicon-strip detectors (DSSSDs) which were capable of performing ion and  $\beta$ correlation measurements. Advanced Implantation Detector Array (AIDA) [21] was used in the first run while the DSSSD stack, wide-range active silicon-strip stopper array for  $\beta$  and ion detection (WAS3ABi) [22], and an yttrium orthosilicate (YSO) scintillator [23,24] were employed for the second run. The typical total rate of the ion implantation in AIDA during the first run was  $\approx$ 150 cps, while that in WAS3ABi during the second run was  $\approx$ 60 cps.

The active stopper array was placed in the center of the high-density polyethylene moderator of the <sup>3</sup>He neutron counter array, BRIKEN [25]. The BRIKEN system is composed of 140 <sup>3</sup>He counters and two clover-type HPGe detectors [26] for high-resolution  $\gamma$ -ray detection. In this configuration, BRIKEN has 62(2)% neutron effective efficiency ( $\epsilon_n$ ) at  $\approx$ 1 MeV of neutron energy [25].

The neutron-gated ion- $\beta$  time spectra obtained in the second run are presented in Fig. 1. Neutron events are correlated with a  $\beta$ -decay event within a 200- $\mu$ s time window after the  $\beta$ -ray detection. The half-lives and initial decay rates at  $T_{\beta} - T_{\text{ion}} = 0$  for each neutron multiplicity ( $A_{0n}$ ,  $A_{1n}$ , and  $A_{2n}$ ) are obtained by binned maximum likelihood fitting to a convolution of contributions from the decays of parent, daughter,  $\beta 1n$ , and  $\beta 2n$ -daughter as well as a linear background, neglecting the small contribution of other descendants. The half-lives from 1n spectra are adopted since the 1n spectra have a smaller component from the decays of descendant nuclei than On spectra and the statistical error is smaller than in the 2n spectra, see Refs. [27,28]. The  $T_{1/2}$  and  $P_{xn}$  values obtained in this work are summarized in Table I. Because the neutron energy distribution is not known, the neutron detection efficiency makes the dominant contribution to the systematic error associated with our  $P_{xn}$  values. The  $P_{1n}$  and  $P_{2n}$  values obtained in this work for <sup>86</sup>Ga are consistent with the data by Miernik *et al.* [4]. The  $T_{1/2}$  and  $P_{1n}$ ,  $P_{2n}$  values of <sup>87</sup>Ga are obtained for the first time in this work. Our  $P_{2n}$  value for  ${}^{85}$ Ga, 1.3(2)% is not consistent with the reported upper limit, <0.1% by Miernik *et al.* [18]. Our  $P_{2n}$  value is more reliable due to the fact that we have observed coincidence between two neutrons and the 247-keV  $\gamma$ -line from the  $1/2^+_1 \rightarrow$  $5/2_{g.s.}^+$  transition in <sup>83</sup>Ge following the decay of <sup>85</sup>Ga. A more complete discussion of the coincidence spectroscopy results will be reported in a future publication. The  $P_{2n}$  discrepancy may be attributed to differences in detection thresholds for  $\beta$ particles.

Figure 2(a) shows the comparison of the experimental neutron branching ratio with the shell model calculations with the cutoff model by Madurga *et al.* [16] and the same shell-model calculations but with the Hauser-Feshbach statistical model [30]. When comparing the new experimental results with the predictions by the shell-model calculations, we notice

TABLE I. Half-lives,  $P_{1n}$ , and  $P_{2n}$  obtained in this study. Q values are adopted from [29]; asterisks show statistical errors; daggers, systematic errors.

	$T_{1/2}$	Branching ratio (%)		Q values (MeV)		
Nucl.	(ms)	$P_{1n}$	$P_{2n}$	$Q_{eta}$	$Q_{\beta 1n}$	$Q_{\beta 2n}$
<sup>84</sup> Ga	97.6(12)	$44(4)^{\dagger}$	$1.6(2)^{\dagger}$	14.1(2)	8.3(2)	5.2(2)
<sup>85</sup> Ga	95.3(10)	$90(7)^{\dagger}$	$1.3(2)^{\dagger}$	13.3(3)	10.2(3)	5.0(3)
<sup>86</sup> Ga	49(2)	74(2)*(8) <sup>†</sup>	16.2(9)*(6) <sup>†</sup>	15.3(6)	11.0(4)	7.9(4)
<sup>87</sup> Ga	29(4)	81(9)*(8) <sup>†</sup>	10.2(28)*(5) <sup>†</sup>	14.8(6)	12.1(7)	7.7(5)

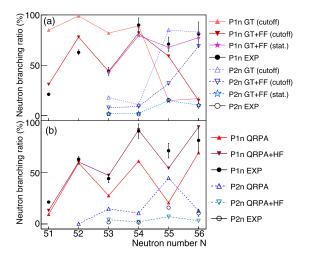


FIG. 2. Comparison of experimental  $P_{1n}$  and  $P_{2n}$  values reported for Ga precursors with calculations. (a) Comparison with shell-model calculations by Madurga *et al.* [16]. "GT" and "GT+FF" in the legend show the shell-model calculations with pure Gamow-Teller and GT + first forbidden transitions, respectively. Each calculation is coupled with the cutoff model (cutoff) and the statistical model (stat.). (b) Comparison with QRPA calculations [5] and those with the Hauser-Feshbach statistical model (QRPA+HF) [9]. Experimental values at N = 53-56 are obtained in this work. N = 51 and 52 are from the references shown in the text.

a discrepancy between the cutoff model and experimental data for all the investigated  $\beta 2n$  gallium precursors, most dramatically manifested in <sup>87</sup>Ga. The  $P_{2n}$  values measured here are much smaller than the cutoff model predictions.

The model we adopt to estimate the GT decay strength distribution for Ga isotopes [16] was based on a shell-model calculation using the NUSHELLX code [31] with hybrid interactions and the truncation described previously in [16,32,33]. In this model, the  $\beta$ -decay properties are dominated by the Gamow-Teller decay of the <sup>78</sup>Ni-core states, leaving the nucleus in the highly excited state because of the N = 50 shell gap. The coupling of valence neutrons and protons to the excitations of the <sup>78</sup>Ni core produces a high density of  $\beta$ -populated states. For <sup>86,87</sup>Ga the B(GT) threshold is close to the two-neutron separation energy in the daughter. In both cases, the majority of the  $\beta$  decay feeds states with  $E > S_{2n}$ . The calculation in Madurga *et al.* [16] shows a strong odd-even effect in the P<sub>1n</sub> systematics. This apparent regularity at <sup>82–85</sup>Ga may break down when the two-neutron emission channel opens up.

An important element of the decay process description implemented in this framework is the contribution from the first forbidden transitions to the low excited states in Ge daughters. Despite small matrix elements, their intensities are amplified by the large phase-space factor and result in a significant population of the neutron-bound states decreasing the  $P_{1n}$  as can be observed in Fig. 2(a) up to <sup>85</sup>Ga. The  $Q_{\beta}$ dependence further enhances the observed odd-even effect.

The inherent uncertainties of the B(GT) strength calculations as well as decay energies and neutron separation energies are expected to be strongly coupled with half-lives and  $P_{xn}$ . In order to investigate the consequences of B(GT)

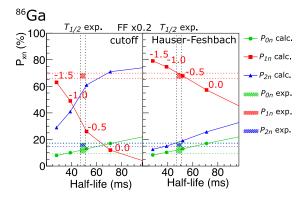


FIG. 3. Shell-model predictions for  $P_{xn}$  vs  $T_{1/2}$  for cutoff and statistical model for the decay of <sup>86</sup>Ga.  $P_{xn}$  values are plotted as a function of shifting the B(GT) distribution from -1.5 to 0.0 MeV and fixed FF contribution. The horizontal and vertical dashed lines show the range of the experimental  $T_{1/2}$  and  $P_{xn}$  values with errors. For the experimental values of <sup>86</sup>Ga, overlaps between Miernik *et al.* [4] are plotted. The calculated  $P_{xn}$  with B(GT) shift up to +1.5 MeV are even farther away from the experimental values both for the cutoff and the statistical model. The plot demonstrates that even if there were uncertainties in the shell model as large as  $\pm 1.5$  MeV, it is impossible to reproduce the experimental data within the cutoff model. In contrast, when using a statistical model, calculated  $P_{xn}$  values stay around the experimental values and are relatively insensitive to the amount of B(GT) shift.

strength uncertainty, we varied the relative position of the B(GT) distribution to match the experimental data on  $T_{1/2}$  and  $P_{xn}$ . We assumed 50% quenching of the strength as in Ref. [16]. The FF contribution is constrained by the  $P_{0n}$  and half-lives.

 $Q_{\beta}$  and neutron separation energies are taken from the recent mass evaluations [29]. This procedure allows us to determine the best parameters for each isotope to describe  $P_{0n}$  and  $T_{1/2}$ , but as shown in Fig. 3, the cutoff model is not able to reproduce the experimental  $P_{1n}$  and  $P_{2n}$ . The same scheme to vary the strength distribution was repeated including statistical particle evaporation and  $\gamma$  emission [10,30]. The results are plotted in the right panel of Fig. 3. In this case for the A = 84-87 isotopes of gallium, we find very good agreement with experiment without major modification of the B(GT) distribution positions (by only -0.5 MeV) and the adjusted value of the FF contribution: a factor 0.2 for <sup>86</sup>Ga and 1.0 for <sup>87</sup>Ga compared to that of <sup>83</sup>Ga and <sup>84</sup>Ga. We consider these empirical adjustments to be within the model uncertainties.

In contrast to the cutoff model, the inclusion of the statistical model correctly reproduces the dominating role of one-neutron emission from two-neutron unbound states. The same conclusion, on the necessity of adding a statistical model can be drawn from QRPA model, see Fig. 2(b), which compares the predictions for  $P_{1n}$  and  $P_{2n}$  with [5] and without [9] HF. Here, however, a very strong odd-even effect is due to the combined effects of deformation [34] or forbidden transitions, which persists in <sup>86</sup>Ga and <sup>87</sup>Ga and results in worse agreement between data and the prediction.

This result is the first demonstrated case in medium and heavy nuclei where the effects of statistical emission must be

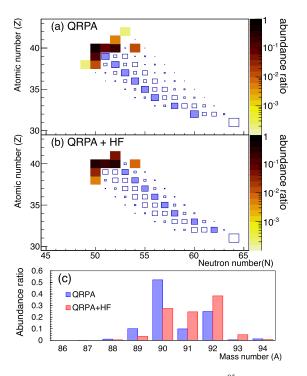


FIG. 4. Decay paths and final abundances of  ${}^{95}$ Ga nuclei simulated from  $P_{xn}$  and  $T_{1/2}$  values by (a) QRPA and (b) QRPA+HF calculation. The area of the blue boxes is proportional to its maximum population during the decay process. The filled boxes show the most abundant nuclei in each isotope. (c) Mass (*A*) projection of the final abundances, plotted as ratios relative to the initial number of  ${}^{95}$ Ga nuclei.

considered in order to model  $\beta$ -delayed multineutron emission. We have also examined the influence of the inclusion of the statistical model on the isotopic distribution of the r-process abundances. As pointed out by Mumpower et al. in their theoretical evaluation [10], this is particularly important for the *r*-process modeling in scenarios where the majority of the nuclei are  $\beta$ -delayed multineutron emitters, such as in the recently discovered neutron star merger [35]. We have modeled the final isotopic distribution resulting from the decay of a single r-process isotope following up every possible decay path on the way to stability. As an example, we have chosen <sup>95</sup>Ga, which was identified to be one of the most exotic abundantly populated gallium isotopes at the freeze-out [36] in the low-entropy scenario. The results are shown in Fig. 4, where all the isotopes populated in the decay of <sup>95</sup>Ga are drawn, the lifetimes and branching ratios from both Möller QRPA models were used [5,9]. Among N > 50 isotopes populated in the chain of <sup>95</sup>Ga decay, more than half of them are populated in the first second after formation of <sup>95</sup>Ga and all of them are 1n, 2n, or 3n emitters. Their respective  $P_{xn}$  determine the final isotopic distribution. The inclusion of HF increases the population of heavier zirconiums significantly.

In summary, we discovered new  $\beta$ -delayed two-neutron emitters <sup>84,85,87</sup>Ga, and measured their two-neutron branching ratio for the first time. For <sup>87</sup>Ga, its  $P_{1n}$  and  $T_{1/2}$  values are measured also for the first time. The  $P_{1n}$  and  $T_{1/2}$  of <sup>84,85,86</sup>Ga are measured with better precision than previous studies. In

all of the nuclei, the shell-model and QRPA calculations could reproduce the experimental neutron branching ratios and the half-lives only if the statistical model is incorporated. The conventional cutoff model cannot describe the experimental data using previously established model parameters. The results show that the measurements of  $\beta$ -delayed neutron emission branching ratios cannot be used in a straightforward way to deduce the strength distribution, but a model of competing particle and  $\gamma$ -ray emission must be included. These results suggest that decays via one-neutron emission dominate even from states which are two-neutron unbound and that it is critical to consider the competition between one- and two-neutron emission in  $\beta$ -delayed neutron emission models, which is of particular importance for r-process modeling. We have demonstrated the sensitivity of the final isotopic distribution to the inclusion of the statistical model. The statistical model approach, which is insensitive to the details of the nuclear structure, provides a simple prescription of  $\beta$ -neutron modeling. Further studies are needed to prove if it is universal.

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- [1] R. B. Roberts, R. C. Meyer, and P. Wang, Phys. Rev. 55, 510 (1939).
- [2] R. E. Azuma, L. C. Carraz, P. G. Hansen, B. Jonson, K. L. Kratz, S. Mattsson, G. Nyman, H. Ohm, H. L. Ravn, A. Schroder, and W. Ziegert, Phys. Rev. Lett. 43, 1652 (1979).
- [3] T. Kawano, P. Möller, and W. B. Wilson, Phys. Rev. C 78, 054601 (2008).
- [4] K. Miernik, K. P. Rykaczewski, C. J. Gross, R. Grzywacz, M. Madurga, D. Miller, J. C. Batchelder, I. N. Borzov, N. T. Brewer, C. Jost, A. Korgul, C. Mazzocchi, A. J. Mendez, Y. Liu, S. V. Paulauskas, D. W. Stracener, J. A. Winger, M. Wolińska-Cichocka, and E. F. Zganjar, Phys. Rev. Lett. **111**, 132502 (2013).
- [5] P. Möller, B. Pfeiffer, and K.-L. Kratz, Phys. Rev. C 67, 055802 (2003).
- [6] T. Marketin, L. Huther, and G. Martínez-Pinedo, Phys. Rev. C 93, 025805 (2016).
- [7] I. N. Borzov, Phys. At. Nucl. 74, 1435 (2011).
- [8] K. Miernik, Acta Phys. Pol. B 47, 739 (2016).
- [9] P. Möller, M. R. Mumpower, T. Kawano, and W. D. Myers, At. Data Nucl. Data Tables 125, 1 (2019).
- [10] M. R. Mumpower, T. Kawano, and P. Möller, Phys. Rev. C 94, 064317 (2016).
- [11] R. A. Warner and P. L. Reeder, Radiat. Eff. 94, 27 (1986).
- [12] F. M. Mann, M. Schreiber, R. E. Schenter, and T. R. England, Nucl. Sci. Eng. 87, 418 (1984).
- [13] G. Rudstam, K. Aleklett, and L. Sihver, At. Data Nucl. Data Tables 53, 1 (1993).
- [14] J. A. Winger, S. V. Ilyushkin, K. P. Rykaczewski, C. J. Gross, J. C. Batchelder, C. Goodin, R. Grzywacz, J. H. Hamilton, A. Korgul, W. Królas, S. N. Liddick, C. Mazzocchi, S. Padgett, A. Piechaczek, M. M. Rajabali, D. Shapira, E. F. Zganjar, and I. N. Borzov, Phys. Rev. Lett. **102**, 142502 (2009).
- [15] J. A. Winger, K. P. Rykaczewski, C. J. Gross, R. Grzywacz, J. C. Batchelder, C. Goodin, J. H. Hamilton, S. V. Ilyushkin, A. Korgul, W. Królas, S. N. Liddick, C. Mazzocchi, S. Padgett, A. Piechaczek, M. M. Rajabali, D. Shapira, E. F. Zganjar, and J. Dobaczewski, Phys. Rev. C 81, 044303 (2010).
- [16] M. Madurga, S. V. Paulauskas, R. Grzywacz, D. Miller, D. W. Bardayan, J. C. Batchelder, N. T. Brewer, J. A. Cizewski, A. Fijałkowska, C. J. Gross, M. E. Howard, S. V. Ilyushkin, B. Manning, M. Matoš, A. J. Mendez, K. Miernik, S. W. Padgett, W. A. Peters, B. C. Rasco, A. Ratkiewicz *et al.*, Phys. Rev. Lett. **117**, 092502 (2016).
- [17] D. Verney, D. Testov, F. Ibrahim, Y. Penionzhkevich, B. Roussière, V. Smirnov, F. Didierjean, K. Flanagan, S. Franchoo, E. Kuznetsova, R. Li, B. Marsh, I. Matea, H. Pai, E. Sokol, I. Stefan, and D. Suzuki, Phys. Rev. C 95, 054320 (2017).
- [18] K. Miernik, K. P. Rykaczewski, R. Grzywacz, C. J. Gross, M. Madurga, D. Miller, D. W. Stracener, J. C. Batchelder, N. T. Brewer, A. Korgul, C. Mazzocchi, A. J. Mendez, Y. Liu, S. V. Paulauskas, J. A. Winger, M. Wolińska-Cichocka, and E. F. Zganjar, Phys. Rev. C 97, 054317 (2018).
- [19] T. Kubo, Nucl. Instrum. Methods Phys. Res., Sect. B 204, 97 (2003).
- [20] T. Ohnishi, T. Kubo, K. Kusaka, A. Yoshida, K. Yoshida, M. Ohtake, N. Fukuda, H. Takeda, D. Kameda, K. Tanaka, N. Inabe, Y. Yanagisawa, Y. Gono, H. Watanabe, H. Otsu, H. Baba, T. Ichihara, Y. Yamaguchi, M. Takechi, S. Nishimura *et al.*, J. Phys. Soc. Jpn. **79**, 073201 (2010).
- [21] C. J. Griffin, T. Davinson, A. Estrade, D. Braga, I. Burrows, P. Coleman-Smith, T. Grahn, A. Grant, L. J. Harkness-Brennan,

M. Kogimtzis, I. Lazarus, S. Letts, Z. Liu, G. Lorusso, K. Matsui, S. Nishimura, R. D. Page, M. Prydderch, V. Pucknell, S. Rinta-Antila, O. Roberts, D. A. Seddon, J. Simpson, J. Strachan, S. L. Thomas, and P. J. Woods, PoS (NIC XIII), 097 (2014).

- [22] S. Nishimura, G. Lorusso, Z. Xu, J. Wu, R. Gernh, H. S. Jung, Y. K. Kwon, Z. Li, K. Steiger, and H. Sakurai, RIKEN Accelerator Prog. Rep. 46, 182 (2013).
- [23] R. Grzywacz *et al.*, RIKEN Accelerator Prog. Rep. **51**, 150 (2018).
- [24] R. Yokoyama, M. Singh, R. Grzywacz, A. Keeler, T. T. King, J. Agramunt, N. T. Brewer, S. Go, J. Heideman, J. Liu, S. Nishimura, P. Parkhurst, V. H. Phong, M. M. Rajabali, B. C. Rasco, K. P. Rykaczewski, D. W. Stracener, J. L. Tain, A. Tolosa-Delgado, K. Vaigneur *et al.*, Nucl. Instrum. Methods **937**, 93 (2019).
- [25] A. Tarifeño-Saldivia, J. L. Tain, C. Domingo-Pardo, F. Calviño, G. Cortés, V. H. Phong, A. Riego, J. Agramunt, A. Algora, N. Brewer, R. Caballero-Folch, P. J. Coleman-Smith, T. Davinson, I. Dillmann, A. Estradé, C. J. Griffin, R. Grzywacz, L. J. Harkness-Brennan, G. G. Kiss, M. Kogimtzis *et al.*, J. Instrum. 12, P04006 (2017).
- [26] C. J. Gross, T. N. Ginter, D. Shapira, W. T. Milner, J. W. McConnell, A. N. James, J. W. Johnson, J. Mas, P. F. Mantica, R. L. Auble, J. J. Das, J. L. Blankenship, J. H. Hamilton, R. L. Robinson, Y. A. Akovali, C. Baktash, J. C. Batchelder, C. R. Bingham, M. J. Brinkman, H. K. Carter *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **450**, 12 (2000).
- [27] B. C. Rasco, N. T. Brewer, R. Yokoyama, R. Grzywacz, K. P. Rykaczewski, A. Tolosa-Delgado, J. Agramunt, J. L. Tain, A. Algora, O. Hall, C. Griffin, T. Davinson, V. H. Phong, J. Liu, S. Nishimura, G. G. Kiss, N. Nepal, and A. Estrade, Nucl. Instrum. Methods Phys. Res., Sect. A **911**, 79 (2018).
- [28] A. Tolosa-Delgado, J. Agramunt, J. L. Tain, A. Algora, C. Domingo-Pardo, A. I. Morales, B. Rubio, A. Tarifeño-Saldivia, F. Calviño, G. Cortes, N. T. Brewer, B. C. Rasco, K. P. Rykaczewski, D. W. Stracener, J. M. Allmond, R. Grzywacz, R. Yokoyama, M. Singh, T. King, M. Madurga *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **925**, 133 (2019).
- [29] M. Wang, G. Audi, F. G. Kondev, W. J. Huang, S. Naimi, and X. Xu, Chin. Phys. C 41, 030003 (2017).
- [30] T. Kawano, P. Talou, I. Stetcu, and M. B. Chadwick, Nucl. Phys. A 913, 51 (2013).
- [31] B. A. Brown and W. D. M. Rae, Nucl. Data Sheets 120, 115 (2014).
- [32] C. Mazzocchi, A. Korgul, K. P. Rykaczewski, R. Grzywacz, P. Bączyk, C. R. Bingham, N. T. Brewer, C. J. Gross, C. Jost, M. Karny, M. Madurga, A. J. Mendez, K. Miernik, D. Miller, S. Padgett, S. V. Paulauskas, D. W. Stracener, and M. Wolińska-Cichocka, Acta Phys. Pol. B 46, 713 (2015).
- [33] M. F. Alshudifat, R. Grzywacz, M. Madurga, C. J. Gross, K. P. Rykaczewski, J. C. Batchelder, C. Bingham, I. N. Borzov, N. T. Brewer, L. Cartegni, A. Fijałkowska, J. H. Hamilton, J. K. Hwang, S. V. Ilyushkin, C. Jost, M. Karny, A. Korgul, W. Królas, S. H. Liu, C. Mazzocchi *et al.*, Phys. Rev. C **93**, 044325 (2016).
- [34] P. Sarriguren, Phys. Rev. C 91, 044304 (2015).
- [35] B. P. Abbott *et al.* (LIGO Scientific Collaboration and Virgo Collaboration), Phys. Rev. Lett. **119**, 161101 (2017).
- [36] J. Lippuner and L. F. Roberts, Astrophys. J. Suppl. Ser. 233, 18 (2017).