1	Multi-proxy record of orbital-scale changes in climate and sedimentation
2	during the Weissert Event in the Valanginian Bersek Marl Formation
3	(Gerecse Mts., Hungary)
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21 Abstract

22 The Valanginian positive carbon isotope excursion and associated environmental 23 changes, known as the Weissert Event, is the first in the series of Cretaceous Earth 24 system perturbations. Here, we develop a multiproxy cyclostratigraphy from a 31.2-m-25 thick Upper Valanginian to lowermost Hauterivian section of the Bersek Marl Formation 26 in Gerecse Mountains, Hungary, comprising alternating marlstone layers of varying clay 27 and carbonate content. The bulk carbonate δ^{13} C signal shows sustained, elevated values 28 (up to 2.7%) up to 19.2 m, followed by a decreasing trend upsection. Together with 29 biostratigraphic data, this suggests that the lower part of the section was deposited 30 during the plateau phase of the Valanginian Weissert Event. Spectral analyses of the 31 multiproxy dataset, including magnetic susceptibility measurements and gamma-ray 32 spectroscopy on the lower part of the section led to the identification of precession, 33 obliquity, long and short eccentricity signals. A mean sedimentation rate of 14 m/Myr 34 was calculated based on astronomical tuning. The cyclicity in the proxy signals reflects 35 dilution cycles by detrital inputs in the basin, which supports the idea that orbitally-36 forced humid-arid cycles controlled the pelagic alternating sedimentation during the 37 Early Cretaceous throughout the Tethyan area.

38 **1. Introduction**

The Cretaceous period is well known for oceanic anoxic events (OAEs) and carbon isotope excursions (CIEs), which have been linked to the volcanism of large igneous provinces (Erba, 2004; Jenkyns, 2010). The OAEs are episodes of increased organic carbon burial in sediments driven primarily by climate warming (Schlanger and Jenkyns, 1976). The CIEs often accompany the OAEs, but they may also occur independently, with no correlated organic-rich deposits (Westermann et al., 2010; Föllmi, 2012).

46 The Late Valanginian Weissert Event is the first positive CIE of the Cretaceous 47 (Weissert et al., 1998; Erba et al., 2004; Weissert and Erba, 2004; Föllmi, 2012). As 48 organic-rich deposits are not a characteristic of this carbon isotope excursion, it is best 49 described as a stand-alone CIE rather than an OAE (Westermann et al., 2010; Kujau et 50 al., 2012). The Weissert Event is notably associated with a widely distributed decline in 51 carbonate production in neritic regions (Weissert et al., 1998; Föllmi et al., 2006), 52 increases in atmospheric pCO_2 (Morales, 2013), more humid climate conditions and 53 intensified continental weathering (Duchamp-Alphonse et al., 2011; Kujau et al., 2013; 54 Charbonnier et al., 2016), increases in nutrient availability in marine environments 55 (Duchamp-Alphonse et al., 2007, 2014; Mattioli et al., 2014), and major turnover events 56 in marine faunas (Reboulet and Atrops, 1997; Melinte and Mutterlose, 2001; Gréselle et 57 al., 2011; Barbarin et al., 2012). Numerous studies have shown the global extent of the 58 CIE which was first observed in deep-sea sediment cores from the Southeastern Gulf of 59 Mexico (Cotillon and Rio, 1984), then subsequently observed in Italy (Lini et al., 1992; Channell et al., 1993; Sprovieri et al., 2006; Bersezio et al., 2002; Erba et al., 2004; 60 61 Amodio et al., 2008;), in the Vocontian Trough in Southern France (Hennig et al., 1999;

62 van de Schootbrugge et al., 2000; Duchamp-Alphonse et al., 2007; Gréselle et al., 2011; 63 Charbonnier et al., 2013), in the Polish Basin (Morales et al., 2015), in the Southern 64 Carpathians in Romania (Barbu and Melinte-Dobrinescu, 2008; Grădinaru et al., 2016), 65 in Greenland (Möller et al., 2015), in the Neuquén Basin in Argentina (Aguirre-Urreta el al., 2008), in Northeastern Mexico (Adatte et al., 2001), in Western Siberia (Price and 66 67 Mutterlose, 2004), and in various deep-sea sediment cores (Katz et al., 2005; Tremolada et al., 2006; Littler et al., 2011). Continental records for the Weissert Event are also 68 69 known from the Crimea and Russia (Gröcke et al., 2005; Nunn et al., 2010). In Hungary, 70 the CIE has been documented in the Hárskút section in the Bakony Mountains (Főzy et al., 2010), where the bulk rock δ^{13} C curve reported in Főzy et al. (2010) on Biancone-71 72 type carbonates shows a positive shift that is characterised by a sharp rise and a 73 possibly suppressed plateau phase due to sedimentary condensation.

74 The age of the Weissert Event and its relationship to the magmatism of a large 75 igneous province has been subject to debate (Erba, 2004; Erba et al., 2004; Sprovieri et 76 al., 2006; Jenkyns, 2010; Thiede and Vasconcelos, 2010; Martinez et al., 2013), in part 77 due to significant uncertainties in the calibrated age of the Early Cretaceous stage 78 boundaries (Hinnov and Ogg, 2007; Ogg et al., 2016). A recently obtained U-Pb age of a 79 mid-Hauterivian tuff layer in the Neuquén Basin, Argentina (Aguirre-Urreta et al., 2015), 80 anchored the astrochronology of the Valanginian-Hauterivian stages (Martinez et al., 2013, 2015) and led to a revision of the age of onset of the Weissert Event, now dated at 81 82 135.22 ± 1.0 Ma. This age is undistinguishable from the most precise ages calculated for 83 the flood-basalt activity of the Paraná-Etendeka large igneous province (Thiede and 84 Vasconcelos, 2010; Janasi et al., 2011), suggesting a link between the two events 85 (Martinez et al., 2015).

86 This study focuses on the Bersek Marl Formation in Hungary that has an almost 87 150-year-old history of research. It was first described by Hantken (1868), who 88 assigned an Early Cretaceous age to the marlstones of the Gerecse Mountains. Ninety 89 years later, Fülöp (1958) provided a thorough bio- and lithostratigraphic description of 90 these deposits. Recently, a more detailed integrated biostratigraphy for the Bersek 91 Quarry, established by Főzy (1995), Főzy and Fogarasi (2002), and Főzy and Janssen 92 (2009), determined a Valanginian–Hauterivian age for the upper part of the formation. 93 The first estimate of the average sedimentation rate of the Bersek Marl Formation is based on the assessed thickness of the formation and the estimated duration of 94 deposition based on ammonoid stratigraphy (Fogarasi, 1995b). Given an assumed 95 average sedimentation rate of 10 m/Myr, Fogarasi (1995b) posited that the deposition 96 97 of marlstone-limestone couplets in the uppermost part of the Bersek Marl Formation 98 and the lowermost part of the overlying Lábatlan Sandstone Formation were controlled 99 by precession cycles.

100 Here, the sedimentary record of the Weissert Event is investigated at the Bersek 101 Marl Formation to produce the first record of this Early Cretaceous CIE at this locality 102 and within this basin. The first objective was to compile a δ^{13} C record for the upper part 103 of the Bersek Marl Formation in order to identify whether the Weissert Event is entirely 104 or partially recorded within this section. The second objective was to accurately 105 estimate the sedimentary deposition rate by integrating cyclostratigraphic analysis on 106 bulk rock δ^{13} C, magnetic susceptibility, and gamma-ray spectroscopy measurements. 107 New calcareous nannoplankton and ammonoid biostratigraphic analyses were also 108 conducted to improve the age constraint.

109 **2. Geological setting**

110 *2.1. Tectonic and stratigraphic framework*

111 The studied outcrop is situated in one of the abandoned yards of the Bersek 112 Quarry, close to the village of Lábatlan within the Gerecse Mountains (Fig. 1A-B). The 113 GPS coordinates of the studied section are: 47.72145°N and 18.52630°E. The Gerecse 114 Mountains form part of the Transdanubian Range, which in turn belongs the ALCAPA 115 Terrane. During the Alpine orogeny, the Transdanubian Range was interlinked with the 116 geological formations that now form part of the Southern Alps and the Northern 117 Calcareous Alps (Fig. 1C). The Jurassic and Cretaceous strata of the Gerecse Mountains 118 were deposited during the evolution of a Neotethys sub-basin, starting in the Late 119 Triassic. The Late Jurassic pelagic carbonate sedimentation changed in the Berriasian, 120 when the thrust front advanced towards the foreland, and clastic input became the 121 dominant factor. This change marks the formation boundary between the Szentivánhegy 122 Limestone Formation and the overlying Bersek Marl Formation. The clastic input further 123 increased in the Barremian sediments, as inferred from a change in lithology from 124 marlstone to sandstone, marking the base of the Lábatlan Sandstone Formation (Főzy, 125 2013).

The 31.2-m thick section studied lies entirely within the Bersek Marl Formation. The base of the studied section starts at the bottom of the quarry face in this yard and the underlying strata could subsequently not be sampled. This prevented the unambiguous correlation with other quarry yards that expose deeper strata. As a hiatus separates the top of the Bersek Marl Formation and the base of the overlying Lábatlan Sandstone Formation, we decided to limit the top of the studied interval at the top of the

132 Bersek Marl Formation (Fig. 1D). The section studied comprises marlstone layers with fluctuating carbonate content, limestone beds, and sporadic, thin sandstone 133 134 intercalations. The colour of the marlstone transitions from grey to purple at 27.8 m, 135 dividing the section into two informal units, which we will henceforth refer to as the 136 "grey marlstone" and the "purple marlstone". In the "grey marlstone" unit of the Bersek 137 Marl Formation, the bed thicknesses range from a few centimetres to a few decimetres, 138 but the lithological contrast between beds is low, making challenging to recognise the 139 bed boundaries in the field. The bedding of the "purple marlstone" is more apparent due 140 to the higher lithological contrast. The bed thicknesses in this unit are around 0.2–0.3 m. 141 Centimetre-thick green turbidite beds, synsedimentary faults and slumps only occur in 142 the "grey marlstone". Throughout the Bersek Quarry, the base of the overlying Lábatlan 143 Sandstone Formation is marked by a few metre-thick, green sandstone beds with an 144 erosive base, that serve as a stratigraphical marker (hereafter referred to as the "green 145 marker bed"). A detailed stratigraphical column is provided as supplementary material.

146 To date, three different stratigraphic sections have been logged and studied in 147 detail at the Bersek Quarry (Fig. 1B). The cephalopod specimens and sediment samples, 148 used for the studies by Főzy and Fogarasi (2002), Főzy and Janssen (2009), and Price et 149 al. (2011), were collected from Section III of Fig. 1B by a team led by J. Fülöp between 150 1963-65. The section studied by Főzy (1995), referred to as Section II in Fig. 1B, lies 151 ~150 m eastwards from Section III, in a different quarry yard. Fogarasi (1995b) 152 investigated sections practically identical to Section II and Section III. The section 153 studied in this paper is labelled as Section I in Fig. 1B, and is located adjacent to Section 154 II, in the same quarry yard but in a newly sampled segment of the quarry face.

155 Ammonoid and calcareous nannoplankton biostratigraphy have been previously 156 developed for the uppermost strata of the Bersek Marl Formation and the lowermost 157 part of the succeeding Lábatlan Sandstone Formation (Főzy, 1995; Főzy and Fogarasi, 158 2002; Főzy and Janssen, 2009). The rich cephalopod fauna of the "purple marlstone" and 159 the Lábatlan Sandstone Formation allows the recognition of Mediterranean ammonoid 160 zones, whilst macrofossils are scarcer in the underlying "grey marlstone". The youngest 161 strata of the Lábatlan Sandstone Formation, immediately overlying the thick "green 162 marker bed", are earliest Barremian in age (Főzy, 1995; Főzy and Fogarasi, 2002; Főzy 163 and Janssen, 2009). In Section III, the uppermost boundary of the oldest recognised 164 ammonoid zone, the Valanginian Varlheideites peregrinus Zone, is detected 165 approximately 12 m below the "green marker bed". In the same section the youngest 166 recognised ammonoid zones below the "green marker bed" are the Upper Hauterivian 167 *Pseudothurmannia ohmi* and *Balearites baelaris* Zones. The calcareous nannoplankton 168 assemblages of the same collection led to the recognition of the NK3/NC4 and possibly 169 of the NC4/NC5 Zone boundaries below the base of the Lábatlan Sandstone Formation 170 (Főzy and Fogarasi, 2002; Főzy and Janssen, 2009). In Section II, studied by Főzy (1995), 171 the characteristic Hauterivian ammonoid fauna is not present, suggesting a Late 172 Valanginian-Early Hauterivian age for the uppermost strata of this section, directly 173 under the base of the "green marker bed".

174 The oldest strata of the Bersek Marl Formation belong to the Felsővadács Breccia 175 Member. The oldest marlstone layers overlaying this unit are posited to be latest 176 Barremian/earliest Valanginian in age, based on ammonite (Vigh, 1984) and 177 nannoplankton (Fogarasi, 2001) biostratigraphy. These strata do not crop out at the 178 Bersek Quarry, but are visible in other sections.

179 *2.2. Palaeogeographic setting*

180 Interpretation of the sedimentary environment of the Bersek Marl Formation has 181 been controversial. It has been successively regarded as a shallow marine deposit 182 (Fülöp, 1958), flysch (Császár and Haas, 1984), a bathial slope deposit (Kázmér, 1987), 183 or as part of a submarine fan sequence (Sztanó, 1990). Fogarasi (1995a) noted that the 184 total thickness and the sedimentation rate of the marlstone is smaller than would be 185 expected for a submarine fan deposit and instead suggested a pelitic slope environment. 186 According to the most recent study of Fodor et al. (2013), the Bersek Marl Formation 187 was deposited on a forebulge, a slope facing the foreland basin, on the opposite side of 188 the orogenic arc. It has been suggested that the difference in lithology between the "grey 189 marlstone" and the "purple marlstone" is controlled by the basin evolution, where the 190 deposition of "purple marlstone" could reflect an interval of subdued tectonic activity in 191 an otherwise actively forming flexural basin (Fodor et al., 2013). The colour of the "grey 192 marlstone" and the occurrence of small, charred plant fragments suggests suboxic 193 bottom waters, while the reddish colour of the "purple marlstone" and the Zoophycos-194 type trace fossils could imply a transition to a more oxygenated environment (Fodor et 195 al., 2013). Petrographic studies of Császár and Árgyelán (1994) and Árgyelán (1995) 196 identified the source of distinctive heavy minerals in the turbidite beds as the suture 197 zone of the Neotethys. Coeval sandstone beds, in the corresponding units in the 198 Northern Calcareous Alps, the Schrambach and the Rossfeld Formations, have similar 199 compositions (von Eynatten and Gaupp, 1999, Krische et al., 2013).

200 **3. Material and methods**

201 Fieldwork was carried out between April 2014 and August 2015, when the 202 continuous 31.2-m thick Section I of Fig. 1B was measured and logged at the Bersek 203 Quarry. A total of 241 bulk rock samples were collected with a uniform 0.1-m spacing in 204 the lower 16.8 m (169 samples) and with a 0.2-m spacing in the upper 14.4 m (72 205 samples). The samples were cut with a diamond saw blade with low revolution speed, 206 washed with tap water to remove the weathered surfaces, and split into ~ 15 g 207 subsamples that were later processed for analyses. To avoid metallic contamination that 208 could bias the magnetic susceptibility measurements, the sample collection was done by 209 hand, and the samples were stored in plastic bags. Bulk rock samples are deposited in 210 the Department of Palaeontology and Geology of the Hungarian Natural History 211 Museum, Budapest, under inventory numbers INV 2016.215.1-241.

212 *3.1. Stable isotope analyses*

The bulk rock carbonate carbon and oxygen stable isotope analyses were performed at the University of Plymouth, on a GV Instruments IsoPrime IRMS using a Gilson 222XL autosampler. For the measurements, ~0.5 mg powder was drilled from each of the 241 bulk rock samples. Isotope ratios are reported in δ values relative to the Vienna Pee Dee belemnite (VPDB) standard. For the instrument calibration, the NBS-19, the IAEA-CO-8, and the IAEA-CO-9 standards were used. Upon replicate analyses, the standard deviation was calculated as 0.2‰ for δ^{13} C and 0.3‰ for δ^{18} O.

220 *3.2. Magnetic susceptibility measurements*

The magnetic susceptibility (MS) measurements were conducted at the Université de Pau et des Pays de l'Adour on an Agico Kappabridge MFK1-FA type instrument, with a 976 Hz 200 A/m field strength. A set of 169 samples, from the bottom 16.8 m of the section, was analysed. Each sample was measured in triplicate to assess the reproducibility of the measurements. The results are given in m³/kg (mass susceptibility). The standard deviation of the measurements is 8x10⁻³ m³/kg.

227 *3.3. Gamma-ray spectroscopy*

228 The gamma-ray spectroscopy (GRS) measurements were conducted in the field 229 using a hand-held Georadis RS-125 gamma-ray spectrometer. The instrument was fitted 230 with a 103 cm³ Na(Ti) scintillation detector with a 0.06 m diameter detector head. The 231 lower 16.8 m of the section were analysed, except for the lowermost three sampling 232 spots that were covered with soil. The measurements were done on the same spots 233 where the bulk rock samples were collected from, and each spot was analysed three times for a measurement time of 120 seconds. The total ⁴⁰K, ²³⁸U, and ²³²Th content of 234 235 the rocks are reported in uranium-equivalent ppm. The standard deviation was derived 236 from 26 test measurements on one single spot, and it was calculated as 0.9 Ue ppm.

237 *3.4. Time series analyses for cyclostratigraphy*

238 Cyclostratigraphic analyses were made on bulk carbonate δ^{13} C, magnetic 239 susceptibility, and gamma-ray spectroscopy signals. Prior to spectral analysis, the raw 240 data series were linearly interpolated every 0.1 m and long-term trends were removed. 241 The detrending procedures applied here aim at reducing the power of the lowest frequencies towards zero whilst maintaining the powers of higher frequencies. The trends were removed from the δ^{13} C, the MS, and the GRS data series by subtracting a 3rdorder polynomial regression, a 5th-order polynomial regression, and a linear regression, respectively.

246 The spectrum of the whole series was calculated using the multitaper method 247 applied with three 2π -tapers (2π -MTM; Thomson, 1982, 1990). Confidence levels were 248 calculated using robust red-noise modelling, modified according to Tukey's end-point 249 rule (Tukey, 1977; Mann and Lees 1996; Meyers, 2014). In addition to the power 250 spectra, two evolutive spectrograms were generated per dataset using Time-Frequency 251 Fast Fourier Transform (T-F FFT). Firstly, for longer periodicities, an 8-m sliding 252 window and a 0.1-m window step were applied to the complete detrended series. 253 Secondly, for shorter periodicities, a 4-m sliding window and a 0.1-m window step were 254 applied after a second low-pass filter was subtracted from the detrended series to 255 exclude low-frequency, high-power cyclicities. The power spectra and the evolutive 256 spectra are then interpreted together. Variations in the sedimentation rate can result in 257 one Milankovitch-cycle being expressed over several frequencies on the 2π -MTM 258 spectra of the sedimentary series (e.g., Weedon, 2003; Martinez et al., 2016). Changes in 259 the sedimentation rate can be recognised in the evolutive spectra by the deviation of the 260 spectral bands, assigned to the cyclicity in question (Martinez et al., 2015). To isolate 261 certain frequencies from the rest of the series, Taner band-pass filters were applied 262 (Taner, 2003; Meyers, 2014).

263 *3.5. Calcareous nannoplankton*

264 Study of the calcareous nannoplankton assemblages were not intended to be 265 comprehensive and were therefore restricted to 12 samples. Successive highly 266 calcareous and highly clay-rich marlstones from both the "grey marlstone" (samples 267 BQ53-BQ56) and the "purple marlstone" (samples BB48-BB51) were targeted, which 268 corresponded with alternating cycles of peak and minimum carbonate content. Another 269 four samples were studied in order to help define the suspected Valanginian-270 Hauterivian stage boundary (samples BB 1-4). Standard smear slide preparation 271 followed the conventional technique described by Bown (1998). The slides were studied 272 at x1000 magnification under cross-polarized light using an Olympus BX51 microscope 273 with an oil immersion objective. For each sample, 40 fields of view were scanned and 274 the number of complete, non-fragmented coccoliths counted to quantify total 275 nannoplankton abundance (Table 1).

276 *3.6. Ammonoids*

277 During the fieldwork in 2014 and 2015, no systematic fossil collection was 278 performed, but several ammonoid specimens were collected both ex situ and in situ 279 from Section I. The biostratigraphic subdivision of Section II is based on a small-scale 280 but systematic ammonoid collection, which yielded about two dozen ammonoids below 281 the "green marker bed" (Főzy, 1995). The cephalopod fauna of Sections I and II is 282 comparable with the fauna of the lower part of Section III, more specifically with Section 283 C of Főzy and Janssen (2009). This section comprises beds labelled from 200 to 258, and 284 yielded 1660 fossil cephalopod specimens, collected bed-by-bed, which served as a solid 285 foundation for biostratigraphy. Undoubtedly, this fauna originates from below the

"green marker bed" (Fig. 2 in Főzy and Janssen, 2009); however, due to the lack of
detailed documentation of the collection, its precise position remains uncertain.

288 *3.7. X-ray fluorescence measurements*

289 X-ray fluorescence (XRF) measurements were made on a suite of 8 samples to 290 estimate the carbonate content of beds with different lithologies. A total of 20 elements 291 were analysed with a Thermo Scientific Niton XL3t 900 GOLDD+ portable XRF analyser, 292 equipped with a 50 kV X-ray tube and a silver target anode, at the Research Centre for 293 Natural Sciences, Hungarian Academy of Sciences, Budapest. For quantitative analysis, 294 the standardless fundamental parameters method was used with Compton-295 normalization. For data evaluation, Excel and Statistica 12 software was used. The CaCO₃ 296 content was calculated from the XRF data by multiplying the Ca_{XRF} content, given in 297 weight percent, by 2.5. The standard error of the Ca_{XRF} measurements is lower than 298 0.2 wt%.

299 **4. Results**

300 *4.1. Stable isotope analyses*

At the Bersek Quarry section, the bulk carbonate $\delta^{13}C_{VPDB}$ values range from 0.2‰ to 3.6‰, with an average of 2.6‰ (Fig. 2). The isotope curve can be divided into two parts on the basis of long-term trends. From 0 m to 19.2 m above the base of the section, the curve is stable with no long-term trend, with values oscillating between 1.6‰ and 3.6‰ and an average value of 2.7‰. The second part of the curve, starting from 19.4 m above the base, displays a decreasing trend, with values dropping to arange between 2.8% to 0.2%.

The bulk carbonate $\delta^{18}O_{VPDB}$ values vary between -5.6‰ and -1.3‰, with an average of -2.7‰ (Fig. 2). The curve shows a slight trend towards increasing values. These negative values are depleted compared to unaltered marine calcite (van de Schootbrugge et al., 2000). The slight covariance of $\delta^{13}C$ and $\delta^{18}O$ values suggests some degree of diagenetic overprint. As the oxygen isotopes are generally more prone to alteration than the carbon isotope values (Sprovieri et al., 2006), we interpret only the latter in the following discussion.

315 *4.2. Magnetic susceptibility measurements*

The magnetic susceptibility values fall between 6.266x10⁻⁸ m³/kg and 125.01x10⁻⁶ m³/kg, with an average of 44.781x10⁻⁸ m³/kg (Fig. 2). Three samples (BQ4, BQ17, and BQ37) show values over 160x10⁻⁸ m³/kg and are considered as outliers as they contain sand grains originating from nearby turbidite beds, disturbing the information from the pelagic chronic sedimentation. Three minima can be found at 3 m, 10 m, and 14.5 m. No distinctive long-term drift can be observed in the magnetic susceptibility values.

323 *4.3. Gamma-ray spectroscopy*

The gamma-ray spectroscopy values fall between 27.5 Ue ppm and 46.8 Ue ppm (Fig. 2). As the organic matter content does not change significantly in the studied section and the turbidite beds are too rare and thin to considerably change the measured values, the gamma-ray values produced here are considered to reflect the changes in the detrital vs. carbonate contribution in the sediment. The minima of the gamma-ray spectroscopy curve are located at the same stratigraphic levels as those in the magnetic susceptibility curve. The gamma-ray spectroscopy curve shows a longterm trend toward slightly decreasing values.

4.4. Time series analyses for cyclostratigraphy

333 The power spectrum of the magnetic susceptibility signal shows spectral peaks 334 over the 99% confidence level (CL) with periods of 5.12 m and 0.23 m, and over the 95% 335 CL with periods of 1.60 m and 0.28 m (Fig. 3A). Other peaks exceed the 90% CL with 336 periods of 0.56 and 0.40 m, but do not reach the 95% CL. As they are observed in the 337 other proxies and in the evolutive spectral analyses, these peaks will be discussed in the 338 following sections of the manuscript. In the spectrogram focusing on the low frequencies 339 (Fig. 4H), the peak centred on 5.12 m is detected throughout the series with the highest spectral power around 12 m above the base of the section. This band first increases from 340 341 3.8 m to 7.9 m, from the base of the section to 6 m above the base, and then decreases to 342 \sim 5 m, towards 15 m above the base. In the spectrogram focusing on the high frequencies 343 (Fig. 4I), the peak centred on 1.60 m first shows a bifurcation to two periods at 1.6 m 344 and 0.9 m, from the base of the section to 5 m above the base. After an interval from 5 m 345 to 7 m above the base of the section, the band appears again from 7 m to 13 m above the 346 base, with a period evolving from 1.3 m to 1.9 m. In the upper part of the section, the 347 expression of this band is observed with a period of 0.9 m. The peaks of periods 0.23 m 348 and 0.28 m appear from 4 m to 13 m above the base of the section. Another band of periods is observed with periods ranging from 0.4 m to 0.6 m at 5 m, 13 m and 16 mabove the base of the series.

351 The power spectrum of the gamma-ray spectroscopy signal shows peaks over the 352 99% CL with periods of 5.12 m and 0.81 m (Fig. 3B). A peak with a period of 0.57 m 353 reaches the 95% CL. Another peak is observed with a period of 1.60 m over the 90% CL. 354 Although this peak does not reach the 95% CL, we will discuss the expression of this 355 peak on the spectrograms. No significant peak with periods shorter than 0.3 m can be 356 recognised because, irrespective of the 0.1 m sample step, the measurements cannot be 357 regarded pointwise. Due to its size, the detector collects 95% of the signal from a 0.3 m 358 radius which smoothens the short-term signal and increases the effective Nyquist-359 frequency of the GRS signal. In the spectrogram focusing on the low frequencies (Fig. 360 4E), the cyclicity band centred on 5.12 m is observed throughout the series with high 361 spectral power. Its period increases from 4.8 m to 7.0 m from the base of the section to 362 7 m above the base, and then decreases to 3.0 m towards 16 m above the base. In the 363 spectrogram focusing on the high frequencies (Fig. 4F), the cyclicity band centred on 364 1.60 m is apparent throughout the section, with periods ranging from 2.0 m to 0.9 m. 365 The band centred on 0.57 m appears from 2.5 m to 6.5 m above the base of the section.

The power spectrum of the δ^{13} C signal shows cyclicities over the 99% CL with periods of 0.59 m, 0.32 m, and 0.21 m, and peaks over the 95% CL with periods of 5.23 m, 1.96 m, 0.74 m and 0.29 m (Fig. 3C). In the spectrogram focused on low frequencies (Fig. 4B), the cyclicity band centred on 5.23 m is observed with high amplitudes from the base of the section to 9 m above the base, and from 18 m above the base to the top of the section. From 10 m to 18 m above the base of the section, this band has much lower amplitudes. Its period varies between 5.7 m and 5.1 m. On the spectrogram focused on

373 the high frequencies (Fig 4C), the band centred on 1.96 m is observed from the base of 374 the section to 3 m with a period of 1.8 m. It appears transiently around 14 m and 19 m 375 above the base of the section, with periods of 1.7 m and 2.2 m, respectively. It then 376 appears as a continuous band from 23 m to the top of the series. The band with periods 377 ranging from 0.74 m to 0.58 m first appears around 6 m above the base of the section 378 with a period of 0.8 m. It is then observed from 8 m to 12 m above the base of the section 379 with two periods of 0.8 m and 0.6 m. It reaches the highest amplitudes from 15 m to 380 27 m above the base of the section. In this interval, this band has periods decreasing 381 from 0.8 m to 0.6 m. It is then expressed from 28 m above the base to the top of the 382 section with a period of 0.9 m. The band of periods ranging from 0.32 m to 0.29 m 383 appears between 8 m and 12 m above the base of the section. The peak centred on 384 0.43 m is observed with low amplitudes around 3 m and 18 m above the base of the 385 section.

386 Comparing the power spectra with the spectrograms of the three signals 387 distinguishes three periodicity bands. The most prominent periodicity in all three 388 signals appears in the spectrograms focusing on the high frequencies with a mean 389 period of \sim 5 m. The peaks at 1.60 m and 1.96 m form the next band of periods, with a 390 mean period of \sim 1.7 m. The peaks ranging from 0.95 m to 0.43 m are only apparent in 391 the GRS and the δ^{13} C signals and have a mean period of ~0.6 m. The peaks ranging from 392 0.32 m to 0.21 m are only apparent in the magnetic susceptibility and the δ^{13} C signals 393 and has a mean period of ~ 0.3 m.

394 *4.5. Calcareous nannoplankton*

Nannoplankton in the studied samples have low abundance and medium to poor
preservation. Reworking appears negligible despite the presence of fragmentary
specimens. Diagenetic overprint, not uncommon for Cretaceous assemblages of a similar
old age, is manifest in partial dissolution of specimens. The predominant taxa belong to
the cosmopolitan genera *Watznaueria* and *Nannoconus* which are also abundant in
other Tethyan localities (e.g., Duchamp-Alphonse et al., 2007).

401 For nannofossil-based biostratigraphic assignment and recognition of marker 402 taxa, the zonal scheme of Bown (1998) was used. The presence of marker taxa and 403 recognized first and last appearance datum (FAD and LAD) events allow recognition of 404 nannoplankton zones NK3-NC4 throughout the studied section, whereas diagnostic species of the younger NC4-NC5 Zones are absent. Key marker species whose presence 405 406 proves the NK3 and NC4 Zones are *Calcicalathina oblongata* (occurring in all samples) 407 and *Cruciellipsis cuvillieri*. Subzone-level subdivision within the zones NK3 and NC4 is 408 not possible as it would require recognition of marker taxa *Eiffellithus windii*, 409 Rucinolithus wisei and Eiffellithus striatus, none of which have been identified within 410 Section I. Another notable absence is that of *Tubodiscus verenae*, the marker species 411 used elsewhere to determine the boundary between nannoplankton zones NK3 and 412 NC4. Thus, the combined use of NK3-NC4 Zones are recommended, suggesting that the 413 most likely chronostratigraphic assignment of Section I is the upper Valanginian, 414 possibly ranging into the lower Hauterivian.

415 *4.6. Ammonoids*

416 As Section I, the focus of this study, and Section II (Főzy, 1995) are situated close 417 to each other in the same quarry yard, it is not surprising that they represent very

similar stratigraphic ranges for the "grey marlstone" and "purple marlstone" strata 418 419 below the "green marker bed". These ranges include the Upper Valanginian, and possibly extend into the lowermost Hauterivian. Some of the most diagnostic 420 421 ammonoids of these sections are illustrated in Fig. 5. However, Section III (Főzy and 422 Janssen, 2009) is more complete below the "green marker bed" and represents, at least 423 partially, most of the higher Hauterivian ammonoid zones, including the *Crioceratites* 424 Subsaynella sayni, Plesiospitidiscus ligatus, Balearites lorvi. balearis. and 425 Pseudothurmannia ohmi Zones.

426 *4.7. X-ray fluorescence measurements*

427 Elemental geochemistry measurements were conducted primarily to determine 428 the carbonate content of the marlstone and its relationship to the lithology and the 429 nannoplankton abundance (Table 1). The carbonate content of the beds in the "grey 430 marlstone" is quite consistent, fluctuating around 65%. On the contrary, in the "purple 431 marlstone" the contrast between the strata is more pronounced, changing from around 432 40% in the clay-rich beds to as high as 90% in the calcareous beds. In both units, the 433 carbonate content and the total nannoplankton abundance show a strong positive 434 correlation ($R^2=0.99$).

435 **5. Discussion**

436 *5.1. Biostratigraphy*

437 The ammonoid fauna at the Bersek Quarry has a clear Mediterranean affinity that 438 can be compared with the well-documented cephalopod assemblages of sections in the 439 Vocontian Basin and the Provence Platform in France (Reboulet, 1995), the Betic 440 Cordillera in Spain (Company et al., 2003), and the Northern Calcareous Alps in Austria 441 (Lukeneder, 2005). In Sections I, II and III, long-ranging phylloceratids and lytoceratids 442 are the most common forms, but these do not allow a precise biostratigraphic 443 subdivision. However, in Sections I and II, the occurrence of *Oosterella* and large olcostephanids (e.g., Olcostephanus densicostatus) is diagnostic and indicates a Late 444 Valanginian to earliest Hauterivian age. The neocomitids in these sections (e.g., 445 446 Neocomites neocominesis, Teschenites subflucticulus and Teschenites callidiscus) are also assigned to the Valanginian. *T. callidiscus*, found only in an *ex situ* block, is the index 447 448 form of the uppermost Valanginian T. callidiscus Zone of Reboulet (1995), the 449 equivalent of the *T. callidiscus* Subzone of the *Criosarasinella furcillata* Zone in Reboulet 450 et al. (2014). Section III, described in detail in Főzy and Janssen (2009), is more 451 biostratigraphically complete. Most of the Hauterivian ammonoid zones were, at least 452 partially, documented from the sequence of marlstone below the "green marker bed" 453 (Főzy and Janssen, 2009). In the lower part of Section III, the common appearance of the 454 large- and small-sized olcostephanids (Olcostephanus densicostatus and Olcostephanus 455 *nicklesi*) is diagnostic. The middle and upper part of this section is characterized by the 456 abundance of crioceratids (including Crioceratites nolani and Crioceratites duvali). The 457 genera *Abrytusites* and *Plesiospitidiscus* appear in the upper part of the Section III. Also 458 important is the occurrence of some rare, but stratigraphically important taxa, such as 459 Olcostephanus jeannoti, Subsaynella sayni, and Subsaynella mimica, Euptychoceras 460 meyrati and Pseudothurmannia ohmi.

The bloom of the representatives of the family Holcodiscidae, characterised by a smooth band on the ventrolateral region (e.g., *Jeanthieuloyites* spp.), is a unique feature of the Bersek fauna and was observed in all three sections. Such an abundance of the *Spitidiscus*-related species is unknown from the classical Upper Valanginian sections in France (Reboulet, 1995), but were reported by Avram (1995) from the Carpathians.

The comparison of the ammonoid faunas of the three Bersek Quarry sections reveals that Section I and II are Late Valanginian or Late Valanginian to earliest Hauterivian in age, whereas Section III also includes younger strata up to the Upper Hauterivian strata. This difference can be explained by the submarine erosion during deposition of the "green marker bed", that removed the unconsolidated Middle to Upper Hauterivian layers of sediment from Sections I and II.

472 A crucial issue is the lack of Acanthodiscus radiatus, the zonal index of the 473 lowermost Hauterivian Acanthodiscus radiatus Zone, which appears mainly in the 474 successions of platform deposits in southeast France (e.g., Reboulet, 1995). This species 475 is missing from the deeper water fauna recovered at the Bersek Quarry. Nevertheless, 476 the Valanginian/Hauterivian boundary can be drawn in Section III. The coeval first 477 appearance of the genus Saynella (represented by a big, smooth specimen) and 478 *Olcostephanus hispanicus,* together with the bloom of the genus *Crioceratites,* suggests 479 that the boundary can be placed between Bed 236 and Bed 237 (Főzy and Janssen, 480 2009). These ammonoids were absent in Section I and II, where fewer total fossils were 481 collected. Therefore, the Valanginian/Hauterivian boundary cannot be confidently drawn in these sections, but the possibility of its presence cannot be excluded. 482

483 Our nannoplankton biostratigraphic results are in good agreement with those484 drawn from the ammonoids and concur with the conclusions of previous studies by

485 Fogarasi (1995b, 2001) and Főzy and Fogarasi (2002), which were based on 486 significantly larger sampling (130 smear slides) and documented the presence of 63 487 nannofossil taxa from Section III. These earlier semi-quantitative analyses also observed 488 the dominance of common Tethyan forms such as *C. oblongata, Nannoconus* sp. and *W.* 489 barnesiae. Although many Tethyan marker species were reported and permit 490 biostratigraphic assignment from the NK3 Zone upwards, two key taxa for defining zone 491 and subzone boundaries, *T. verenae* and *Lithraphidites bollii*, have not been reliably 492 recorded (Főzy and Fogarasi, 2002). However, in Section III, Fogarasi (2001) 493 documented the successive LADs of two marker species that range though Section I in 494 our samples, C. cuvillieri and C. oblongata, thus establishing that deposition of the 495 Bersek Marl Formation continued into nannoplankton zone NC5, and the "green marker 496 bed" is best assigned to the NC5b-c Subzones. Thus, the nannoplankton biostratigraphic 497 evidence also suggest that the base of the "green marker bed" is erosive and the topmost 498 "purple marlstone" layers are highly diachronous between sections I and III, being not 499 younger than earliest Hauterivian in the former, and as young as Late Hauterivian in the 500 latter.

501 *5.2. Recognition of the Weissert Event*

502 The shape of the bulk carbonate δ^{13} C curves during the Weissert Event are 503 similar across most of the previously reported sections (Price et al., 2016). At the Early– 504 Late Valanginian transition, the δ^{13} C values rise steeply to values that are approximately 505 1–2‰ higher than before the onset. After this increase, a plateau phase of elevated δ^{13} C 506 values is observed, that is followed by a slow and gradual decline. During the 507 Valanginian and Hauterivian, the plateau of the Weissert Event is the only interval 508 where the δ^{13} C values stabilise at high values (Price et al., 2016).

The δ^{13} C curve of the Bersek Quarry shows elevated values around 2.7‰ until an 509 510 inflexion point at 19.2 m above the base of the section. From there on, the values are 511 gently decreasing. The good preservation of the orbital cyclicities supports the primary 512 origin of the high values and the overall trends (see also discussion in Pellenard et al., 513 2014). Therefore, based on the inferred latest Valanginian/earliest Hauterivian age of 514 the Bersek Quarry section, it is reasonable to interpret the inflexion in the bulk 515 carbonate δ^{13} C curve as marking the termination of the plateau phase of the Weissert 516 Event.

517 *5.3. Astronomically forced cyclic changes in the record*

The ratio of the observed three key periodicities to each other (5.1:1.7:0.6:0.2-0.3) is broadly similar to the ratio of the mean long eccentricity, short eccentricity, obliquity, and precession periods (405:100:36.6:21), according to the La2004 model (Laskar et al., 2004). Therefore, the recognised ~5 m, ~1.7 m, ~0.6 m, and ~0.2-0.3 m periods are interpreted as the 405-kyr ("long") eccentricity, the 100-kyr ("short") eccentricity, the obliquity, and the precession cycles, respectively.

To test this interpretation, the magnetic susceptibility, the gamma-ray spectroscopy signals, and the lower part of the δ^{13} C signal were calibrated using the proposed short eccentricity signal filtered from the magnetic susceptibility signal, in which 12 complete short eccentricity cycles can be counted (Fig. 4J-L). The magnetic susceptibility is used here as a reference because the high resolution of this signal allows the detection of short periods, unlike the GRS signal, so that the evolution of the short eccentricity can be monitored throughout the series on spectrograms (Fig. 4I). In addition, the eccentricity cycles appear continuous on spectrograms of the magnetic susceptibility signal, unlike the δ^{13} C signal (Fig. 4C). However, at levels 1.6 m, 5.7 m and 10.5 m the evolutive spectrum of the MS series shows bifurcation in the short eccentricity band (Fig. 4I), so that its evolution is unclear.

535 The power spectrum of the tuned magnetic susceptibility signal shows 536 prominent peaks at 360 kyr, 105 kyr, and 41 kyr (Fig. 3D). These are nearly identical to 537 the expected periods of the long eccentricity, short eccentricity, and the obliquity cycles. 538 The other peak at 29 kyr can be a consequence of short-term variations of the 539 sedimentation rate or a consequence of aliasing in intervals showing lower 540 sedimentation rates. The power spectrum of the tuned gamma-ray spectroscopy signal 541 shows prominent peaks at 360 kyr, 109 kyr, and 41 kyr, that corresponds to the long 542 eccentricity, short eccentricity, and obliquity periods, respectively (Fig. 3E). The power spectrum of the tuned δ^{13} C signal shows prominent peaks at 500 kyr, 106 kyr, and 543 544 42 kyr, reasonably representing the long eccentricity, the short eccentricity, and the 545 obliquity periods, respectively (Fig. 3F).

546 The amplitude modulation of the 100-kyr band filtered in the magnetic 547 susceptibility signal was further investigated. To define the amplitude modulation of the 548 proposed short eccentricity signal, Hilbert transform was applied to the tuned band-549 passed signal (Fig. 6; Meyers, 2014). The enveloping amplitude modulation signal shows 550 maxima roughly coinciding with the maxima of the proposed long eccentricity signal 551 filtered from the magnetic susceptibility (Fig. 6A). The main peak on the corresponding 2π -MTM power spectrum has mean period of 418 kyr, close to the period of the 405-kyr 552 eccentricity cycle (Fig. 6B). The main modulator of the short eccentricity signal in 553

554 geological time series is the long eccentricity. The close relationship between the 555 amplitude modulation signal of the presumed short eccentricity signal and the 556 presumed long eccentricity signal reinforces our cyclostratigraphical interpretation.

557 *5.4. Sedimentation rate*

558 The average sedimentation rate at the lower part of the Bersek Quarry section is 559 14 m/Myr. It varies between 9–19 m/Myr, with maxima around 1 m, 6 m, and 13 m 560 above the base of the section. The band-pass filter output signals corresponding to the 561 long eccentricity signal are identified in the same phase in the gamma-ray spectroscopy 562 and the magnetic susceptibility datasets, with maxima around 1 m, 6 m, and 14 m above 563 the base of the section (Fig. 8). These maxima almost coincide with the maxima of the 564 sedimentation rate. The first two long eccentricity cycles in the δ^{13} C signal appear to be 565 in antiphase compared to the magnetic susceptibility and gamma-ray spectroscopy 566 signals, but then become in-phase with them around 15 m above the base. In this part of 567 the section, the amplitude of the long eccentricity filter output signal is low, therefore it 568 is likely that the shift in the phase is only apparent and does not result from 569 palaeoenvironmental changes.

The plateau phase of the Weissert Event in the Bersek Quarry section spans until 19.2 m above the base. It was shown, that the deposition of lowermost 16.8 m of the section took approximately 1.25 Myr. It is regarded here as a minimum estimate for the plateau phase of the Weissert Event, since the entire plateau phase is likely not preserved here and not all the preserved interval is included in this interval. Our minimum duration estimate nonetheless agrees with the durations calculated in contemporaneous sections (e.g. Martinez et al., 2015).

577 The average 10 m/Myr sedimentation rate estimated by Fogarasi (1995b) falls 578 within the range of our results. He suggested that the 0.2-m thick marlstone-limestone 579 couplets could indicate precession cycles. According to our calculations, a marlstone-580 limestone couplet driven by precession forcing should be approximately 0.1–0.3 m thick. 581 This is also the variation that is observable in the field in the "purple marlstone" unit. 582 The average sedimentation rate during the Late Valanginian-Early Hauterivian in the 583 Gerecse Mountains is significantly lower than in the Vocontian Trough. The average 584 sedimentation rate during this time at the Orpierre and the La Charce/Vergol/Morenas sections is around 48 m/Myr (Charbonnier et al., 2013; Martinez et al., 2013), and at the 585 586 Angles/Reynier section is around 40 m/Myr (Martinez et al., 2013). In the Capriolo 587 section, located in the Umbria-Marche Basin, a sedimentation rate of 17 m/Myr was 588 calculated for the Late Valanginian-Early Hauterivian, which is comparable to our 589 results (Sprovieri et al., 2006).

590 *5.5. Impact of the sample distance on the cyclostratigraphic interpretations*

591 The precession cycles have shown to have periods close to the detection limit of 592 our measured series (Fig. 3). Low-density sampling in the studied time series can cause 593 distortion of the spectrum in the precession band, making the record of the orbital 594 cycles unclear (Weedon, 2003). Furthermore, a highly fluctuating sedimentation rate 595 can smooth the power spectrum at high frequencies and decrease the power and 596 significance levels of the spectral peaks in an important proportion of the spectrum (see 597 Martinez et al., 2016). The effect of the low-density sample distance, and the variations 598 in the sedimentation rate on the spectral analyses were tested on four ETP (Eccentricity 599 - Tilt - Precession) series calculated for the Late Valanginian using the La2004 (Laskar et

al., 2004) solution (Fig. 7). Three modelled ETP series represent a hypothetical section
with a sedimentation rate of 14 m/Myr, analogous to the Bersek Quarry section,
sampled every 0.05 m, 0.1 m and 0.2 m. A fourth ETP series of the same hypothetical
section, sampled every 0.01 m, is regarded to be the ideal representation of the
astronomical cycles.

605 On the power spectrum corresponding to the 5-kyr and the 10-kyr sample 606 distances, the peaks associated with the long and the short eccentricity, the obliquity, 607 and the precession cycles are present (Fig. 7A-B). However, on the power spectrum 608 corresponding to the 20-kyr sample distance, the peaks associated with the precession 609 cycles are absent (Fig. 7C). Compared to the power spectrum of the ideal ETP series, the 610 obliquity and eccentricity signals are not impacted by aliasing when analysing the ETP 611 series sampled at a resolution of 5 kyr and 10 kyr. When analysing the ETP series 612 sampled with a resolution of 20 kyr, significant changes in power occur at all 613 Milankovitch-band (Fig. 7C), so that the amplitude of the eccentricity band is affected. 614 The series sampled at 10 kyr thus still allows a correct reconstruct of the amplitude of 615 the eccentricity cycles.

616 The effect of a fluctuating sedimentation rate was tested by setting a 617 sedimentation rate cyclically fluctuating from 9 m/Myr to 19 m/Myr with maxima 618 coinciding with the maxima of the long eccentricity to the previously described ETP 619 series. On the power spectra of the modelled signals, the dispersion of the frequencies is 620 apparent, but the Milankovitch-cyclicities remain detectable (Fig. 7D-F). In particular, at 621 0.05 m and 0.1 m, the power spectra do not significantly differ from the power spectrum of the ETP series sampled 0.01 m (Fig. 7D-E). Our test results imply that in a 622 623 hypothetical section that is analogous with the Bersek Quarry section, a proxy signal

624 sampled with a sample step of 0.1 m is capable of recording the long and the short 625 eccentricity, the obliquity, and the precession. Even with a high variation in the 626 sedimentation rate, our sensitivity analysis demonstrates that the long and the short 627 eccentricity signals remain sufficiently well preserved to support our interpretation that 628 the high-power peaks at the high-frequency end of the power spectrum of the δ^{13} C and 629 the magnetic susceptibility signals are feasibly associated with the precession cycles in 630 the lower part of the studied section. A proxy signal sampled with an average sample 631 step of 0.2 m, representative of the upper 14.4 m section sampling protocol, is capable to 632 record the long and the short eccentricity and the obliquity, with however large 633 disturbances in the amplitude of these cycles compared to the series sampled at 0.01 m 634 (Fig. 7F). These tests show that the interval of the series sampled at 0.1 m is suitable for 635 the recognition of the actual amplitude of the eccentricity cycles and thus viable for 636 astronomical tuning.

637 *5.6. Origin of the cyclicities in the record*

638 The origin of the observed Milankovitch cyclicities in the measured proxy signals 639 is most likely related to the variation in terrestrially-derived detrital input (e.g., Cotillon, 640 1987; Mutterlose & Ruffell, 1999; Reboulet et al., 2003; Martinez et al., 2015; Lukeneder 641 et al., 2016). The intensification of the hydrological cycle typical during more humid 642 periods is likely to correspond to an increase in detrital material. This contributes to the 643 elevated gamma-ray spectroscopy and magnetic susceptibility values, with coinciding 644 maxima and minima, respectively. The increased nutrient availability, resulting from 645 enhanced terrestrial input, supports elevated primary production and subsequently 646 decreases the bulk carbonate δ^{13} C values. Hence, it explains the inverse correlation between the δ^{13} C and the MS-GRS signals. The maxima of the sedimentation rate at the studied section concur with the maxima of the long eccentricity in the MS-GRS signals, implying that the sedimentation rate is linked to orbital forcing (Fig. 8). Furthermore, in the "purple marlstone" the δ^{13} C values and the carbonate content show a positive correlation. Since low carbonate content is characteristic of a humid climate with high runoff rate, a decrease in δ^{13} C values agrees with the dilution model (e.g., Fogarasi, 1995).

654 The response of the sedimentary record to the climatic variations is similar across the Tethyan realm, and has been observed in the Vocontian Basin (SE France; 655 656 Cotillon, 1987), in the Lower Saxony Basin (NW Germany; Mutterlose & Ruffell, 1999), 657 and in the Subbetic Domain (SE Spain; Moiroud et al., 2012). In these geological settings, 658 the more argillaceous beds were deposited under a humid climate with high continental 659 runoff, whereas the beds with a higher carbonate content indicate a semi-arid climate 660 with a lower continental detrital influx (Mutterlose & Ruffell, 1999; Moiroud et al., 661 2012). These variations have been observed at both the small scale (i.e., precession or 662 obliquity controlled bed couplets) and larger scale (i.e., eccentricity controlled bed 663 bundles). Migration of carbonates from the originally more clay-rich to the more 664 calcareous beds during the early diagenesis processes has been notably invoked to 665 explain the onset of the marl-limestone alternations (Munnecke et al., 2001). However, 666 the bundling of the climatic cycles into small and larger scales in the above-mentioned 667 geological settings and the fact that clay minerals (insoluble and related to changes in 668 humidity levels) follow the bundling between the precession and the eccentricity cycles 669 strongly supports an orbital control on the Tethyan marlstone-limestone alternations 670 via humid-arid cycles (Cotillon, 1987; Mutterlose & Ruffell, 1999; Martinez et al., 2015). 671 At the Bersek Quarry, the carbonate content and the total nannoplankton abundance in 672 both the "grey marlstone" and the "purple marlstone" show a strong positive 673 correlation. This implies that the carbonate content is highly dependent on 674 nannoplankton production. Even if carbonate migration might have happened, it did not 675 have a significant effect on the lithology. Conversely, the lithological cycles follow the 676 pattern of the orbital forcing. Even the sedimentation rate appears to follow the filter of 677 the long eccentricity cycle, which support the link between the marl-limestone 678 alternations of the Gerecse Mountains and humid-arid cycles orbitally forced.

679 6. Conclusions

680 In this study, a multi-proxy approach was used to assess a local sedimentary 681 record of the globally important Weissert Event. Bulk carbonate carbon and oxygen 682 stable isotope measurements, gamma-ray spectroscopy and magnetic susceptibility 683 analyses were carried out in a 31.2 m thick section at the Bersek Quarry, type locality of 684 the Bersek Marl Formation, in the Gerecse Mountains, Transdanubian Range, Hungary. A 685 Late Valanginian to possibly earliest Hauterivian age of the section is confirmed by 686 ammonoid and calcareous nannoplankton biostratigraphy. Cyclostratigraphic analyses 687 performed on all the studied proxy signals suggest an average sedimentation rate of 688 14 m/Myr. Based on the preservation of Milankovitch-cyclicity and comparison with other Tethyan sections, the values and trend of the δ^{13} C signal are considered to reflect a 689 primary signal. The inflexion point of the δ^{13} C curve, where the sustained plateau of high 690 691 values transitions to a decreasing trend, is interpreted as the termination of the plateau phase of the Late Valanginian Weissert Event. The plateau phase of the Weissert Event 692 693 as recorded in the Bersek Quarry section is at least 1.4 Myr in duration, in agreement

with the range of estimates obtained throughout the Tethyan realm. The proxy signals and the sedimentary pattern of the Bersek Marl Formation suggest the presence of dilution cycles which fits the model established for Tethyan realm and argues that orbitally forced humid-arid cycles are a major determining factor in the generation of the (hemi-)pelagic marlstone-limestone alternations during the Valanginian– Hauterivian ages.

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985 **Figure captions**

986 **Fig. 1**.

987 (A) Location of the Bersek Quarry section on a generalized map of structural units in 988 Hungary (after Haas, 2012). (B) Satellite aerial view of the Bersek Quarry (Google Earth 989 image captured in 2014). The section studied in this paper is marked with "I". Section 990 "II" was studied by Főzy (1995) and Fogarasi (1995b). Section "III" was studied by Főzy 991 and Fogarasi (2002), Főzy and Janssen (2009), and Price et al. (2011). (C) 992 Paleogeographic reconstruction of the Western Tethys area during the 993 Jurassic/Cretaceous transition (after Csontos and Vörös, 2004). SA – Southern Alps, LAA - Lower Austroalpine Nappes, MAA - Middle Austroalpine Nappes, TI - Tirolic Nappes 994 995 (part of the Northern Calcareous Alps). (D) Photograph of the studied section. The 996 boundary between the Lábatlan Sandstone Formation and the Bersek Marl Formation is 997 marked by the erosive "green marker bed". The Bersek Marl Formation is divided into 998 two informal units based on the colour of the marlstone: the "grey marlstone" and the 999 overlying "purple marlstone".

1000 **Fig. 2.**

Bulk carbonate δ^{13} C, δ^{18} O, gamma-ray spectroscopy (GRS), and magnetic susceptibility (MS) signals measured at the Bersek Quarry. The thick lines below each signal denote the long-term trend. The inset shows the correlation between the δ^{13} C and δ^{18} O values, divided into two sets based on the trend in the δ^{13} C signal: plateau phase (0–18.2 m, circles) and slow decrease (18.3–31.2 m, squares). The asterisk marks the end of the plateau phase of the Weissert Event.

1007 **Fig. 3.**

 2π -MTM power spectra of the magnetic susceptibility (A), gamma-ray spectroscopy (B), 1008 1009 and the bulk carbonate δ^{13} C (C) signals. The peaks are interpreted as the 405 kyr "long" 1010 eccentricity, the 100-kyr "short" eccentricity, the obliquity, and the precession cycles. 1011 The indicated values denote the periods of the main significant peaks in metres. The 1012 width of the filters used to separate the long eccentricity are as follows: 0–0.3516 for the 1013 magnetic susceptibility, 0-0.3906 for the gamma-ray spectroscopy, and 0-0.2930 1014 cycles/m for the δ^{13} C signal. The width of the filters used to separate the short eccentricity are as follows: 0.5078-1.3672 cycles/m for the magnetic susceptibility and 1015 for the gamma-ray spectroscopy signals, and 0.3529–1.5294 cycles/m for the δ^{13} C 1016 1017 signal. 2π -MTM power spectra of the tuned magnetic susceptibility (D), gamma-ray 1018 spectroscopy (E) and the bulk carbonate δ^{13} C (F) signals.

1019 **Fig. 4**.

1020 Detrended signals of the $\delta^{13}C$ (A), gamma ray spectroscopy (D), and magnetic 1021 susceptibility (G) measurements, and the corresponding filtered long eccentricity filter 1022 output signals. The first spectrograms of each signal (B, E, H) focuses on the low 1023 frequencies, while the second spectrograms (C, F, I) represent the whole spectrum. The 1024 second spectrogram is filtered, thus the frequencies of the respective first spectrograms 1025 are not represented. The inset shows the short eccentricity signals filtered form the 1026 magnetic susceptibility (I), gamma-ray spectroscopy (K), and the δ^{13} C signals (L). The 12 1027 dotted lines indicate the age model used for the tuning. The width of the Taner-filters 1028 are shown on Fig. 3.

1029 **Fig. 5**.

1030 Representative ammonoids collected from Section I (this study) and Section II (Főzy, 1031 2005). Figured specimens are deposited in the Department of Palaeontology and 1032 Geology of the Hungarian Natural History Museum, Budapest, with inventory numbers 1033 prefixed by "INV". (A) *Neocomites neocomiensis*; INV. 2016.192; Section I, 14.9 m. (B) 1034 Lytoceras sp.; INV. 2016.193; Section I, ex situ. (C) Olcostephanus sp.; INV. 2016.194; 1035 Section I, 1.7 m. (D, E) Jeanthieulovites sp.; INV. 2016.195); Section I, ex situ. (F) 1036 Phyllopachyceras winkleri; INV. 2016.196; Section I, 5.0 m. (G) Oosterella sp.; INV. 1037 2016.197; Section I, ex situ. (H, I) Neocomites sp.; INV. 2016.198); Section II, Bed 26. (J) 1038 Neolissoceras grasianum; INV. 2016.199; Section II, Bed 26. (K) Olcostephanus 1039 densicostatus; (INV. 2016.200), Section II, Bed 27. (L) Teschenites callidiscus; INV. 2016.201; Section II, ex situ. (M) Teschenites subflucticulus; INV. 2016.202; Section II, 1040 1041 Bed 26. (N) Jeanthieuloyites cf. quinquestriatus; INV. 2016.203; Section II, Bed 16. (O) 1042 *Jeanthieuloyites* cf. *quinquestriatus*; INV. 2016.204; Section II, Bed 28.

1043 **Fig. 6.**

1044 (A) Amplitude modulation of the short eccentricity band filtered in the tuned magnetic 1045 susceptibility signal. (B) 2π -MTM power spectrum of the amplitude modulation signal. 1046 The main peak has mean period of 418 kyr, close to the period of the 405-kyr 1047 eccentricity cycle

1048 **Fig. 7.**

1049 Sensitivity analysis of the impact of the sample step and the fluctuating sedimentation 1050 rate on cyclostratigraphic analyses: (A–C) 2π -MTM power spectra of modelled ETP 1051 series representing a Valanginian section with an average sedimentation rate of 1052 14 m/Myr, analogous to the Bersek Quarry section. The three power spectra show the

aliasing effect of an increasing (0.05-0.2 m) sample step. The shaded peaks represent the ideal spectrum showing no effect of aliasing. (D–F) 2π -MTM power spectra of the same modelled series with fluctuating sedimentation rates. The sedimentation rate was set to vary between 9–19 m/Myr with maxima coinciding with the long eccentricity maxima.

1058 **Fig. 8**.

1059 Conceptual model of cyclic oscillations measured in carbon isotope ratio and magnetic 1060 susceptibility and inferred changes in climate and sedimentation. The inverse correlation of the filtered long eccentricity in the δ^{13} C and the GRS-MS signals imply an 1061 1062 orbitally-forced dilution model for the Bersek Marl Formation. (A) During humid 1063 periods, the increase in detrital influx contributes to the elevated sedimentation rate 1064 and MS-GRS values, and through the enhanced production rate to the low δ^{13} C values. 1065 (B) A decrease in detrital influx and hence the MS-GRS values during more arid periods 1066 induces an increase in the bulk carbonate $\delta^{13}C$ values and a decrease in the 1067 sedimentation rate.

























Frequency (cycles/m)



1 **Table 1.**

Sample	Level (m)	Unit	$CaCO_3$ (wt%)	Nannoplankton abundance
BQ53	7.3	"grey marlstone"	67	47
BQ54	7.4	"grey marlstone"	62	43
BQ55	7.5	"grey marlstone"	65	45
BQ56	7.6	"grey marlstone"	74	50
BB48	30.2	"purple marlstone"	42	29
BB49	30.4	"purple marlstone"	78	55
BB50	30.6	"purple marlstone"	45	30
BB51	30.8	"purple marlstone"	90	60

2 Calcium carbonate content (calculated from XRF data) and total nannoplankton abundance

3 (counts of entire specimens in 40 fields of view) of samples collected from the "grey marlstone"

4 and the "purple marlstone" units.

1 Highlights

- Magnetic susceptibility, gamma ray, δ^{13} C, and δ^{18} O data from the Bersek Formation.
- 3 A new western Tethyan record of the Late Valanginian Weissert Event is described.
- Spectral analysis reveals presence of Milankovitch periodicities in proxy signals.
- 5 Dilution cycles control the sedimentation and the proxy signals.
- 6 Minimum duration of the δ^{13} C plateau of the Weissert Event is 1.25 Myr long.

Sample	Leve1	(m)	δ ¹³ C	(‰	VPDB)	δ ¹⁸ 0	(‰	VPDB)
BQm20		0.00			2.36			-4.11
BQm19		0.10			3.07			-2.48
BQm18		0.20			3.08			-3.02
BQm17		0.30			2.22			-3.27
BQm16		0.40			2.43			-3.76
BQm15		0.50			2.19			-3.75
BQm14		0.60			1.95			-3.75
BQm13		0.70			2.29			-3.46
BQm12		0.80			2.31			-2.91
BQm11		0.90			3.01			-3.06
BQm10		1.00			2.85			-2.90
BQm9		1.10			2.98			-2.33
BQm8		1.20			2.98			-2.78
BQm7		1.30			2.93			-2.57
BQm6		1.40			2.60			-3.41
BQm5		1.50			3.04			-3.03
BQm4		1.60			2.80			-2.80
BQm3		1.70			2.85			-2.19
BQm2		1.80			3.31			-2.19
BQm1		1.90			2.78			-2.47
BQO		2.00			2.60			-3.10
BQ1		2.10			2.26			-3.65
BQ2		2.20			3.24			-2.16
BQ3		2.30			2.57			-3.63
BQ4		2.40			2.30			-4.60
BQ5		2.50			3.08			-2.81
BQ6		2.60			2.81			-2.50
BQ7		2.70			2.52			-3.30
BQ8		2.80			2.80			-2.38
BQ9		2.90			3.50			-1.90
BQ10		3.00			3.57			-1.98
BQ11		3.10			3.29			-2.55
BQ12		3.20			2.81			-3.05
BQ13		3.30			2.62			-2.81
BQ14		3.40			3.22			-2.55
BQ15		3.50			3.13			-1.43
BQ16		3.60			2.92			-2.50
BQ17		3.70			2.20			-5.60
BQ18		3.80			3.03			-2.74
BQ19		3.90			3.11			-2.94
BQ20		4.00			2.71			-2.77
BQ21		4.10			2.60			-2.70
BQ22		4.20			2.80			-2.90

BQ23	4.30	2.71	-3.21
BQ24	4.40	3.25	-2.63
BQ25	4.50	3.31	-2.52
BQ26	4.60	2.97	-2.71
BQ27	4.70	2.30	-3.06
BQ28	4.80	2.33	-3.63
BQ29	4.90	2.11	-3.85
BQ30	5.00	2.03	-3.53
BQ31	5.10	3.06	-2.29
BQ32	5.20	2.84	-1.94
BQ33	5.30	2.74	-2.21
BQ34	5.40	2.54	-2.74
BQ35	5.50	2.04	-4.30
BQ36	5.60	3.01	-2.34
BQ37	5.70	2.10	-3.90
BQ38	5.80	2.74	-2.90
BQ39	5.90	2.41	-3.17
BQ40	6.00	2.60	-2.37
BQ41	6.10	2.99	-2.58
BQ42	6.20	2.38	-3.32
BQ43	6.30	2.16	-2.89
BQ44	6.40	2.50	-2.98
BQ45	6.50	2.84	-3.07
BQ46	6.60	2.16	-4.39
BQ47	6.70	3.13	-2.52
BQ48	6.80	3.25	-2.07
BQ49	6.90	2.79	-2.47
BQ50	7.00	2.42	-2.76
BQ51	7.10	2.71	-3.60
BQ52	7.20	2.46	-3.25
BQ53	7.30	2.70	-2.90
BQ54	7.40	2.90	-2.92
BQ55	7.50	2.53	-2.79
BQ56	7.60	2.74	-2.46
BQ57	7.70	3.18	-2.16
BQ58	7.80	2.80	-3.30
BQ59	7.90	2.64	-2.57
BQ60	8.00	3.03	-2.64
BQ61	8.10	2.95	-2.53
BQ62	8.20	3.10	-1.99
BQ63	8.30	2.50	-3.00
BQ64	8.40	2.26	-3.78
BQ65	8.50	2.98	-1.97
BQ66	8.60	3.23	-1.89

BQ67	8.70	3.13	-1.94
BQ68	8.80	3.45	-1.80
BQ69	8.90	2.85	-2.73
BQ70	9.00	3.19	-2.05
BQ71	9.10	2.69	-2.23
BQ72	9.20	3.61	-1.79
BQ73	9.30	2.60	-2.96
BQ74	9.40	3. 32	-2.50
BQ75	9.50	1.60	-4.70
BQ76	9.60	3.30	-2.45
BQ77	9.70	2.80	-2.40
BQ78	9.80	3.00	-2.45
BQ79	9.90	3.22	-1.96
BQ80	10.00	2.94	-2.23
BQ81	10.10	2.49	-4.17
BQ82	10.20	2.40	-2.90
BQ83	10.30	2.81	-2.92
BQ84	10.40	2.90	-2.50
BQ85	10.50	2.91	-1.28
BQ86	10.60	3.12	-2.23
BQ87	10.70	2.71	-2.82
BQ88	10.80	2.71	-2.65
BQ89	10.90	3.25	-2.11
BQ90	11.00	3.08	-2.33
BQ91	11.10	2.82	-2.27
BQ92	11.20	2.20	-3.10
BQ93	11.30	2.66	-3.02
BQ94	11.40	1.77	-3.20
BQ95	11.50	2.60	-3.60
BQ96	11.60	2.86	-2.29
BQ97	11.70	3.25	-1.94
BQ98	11.80	2.59	-3.10
BQ99	11.90	2.93	-2.52
BQ100	12.00	2.30	-3.20
BQ101	12.10	2.71	-3.07
BQ102	12.20	2.90	-2.50
BQ103	12.30	2.70	-3.10
BQ104	12.40	1.99	-3.04
BQ105	12.50	3.16	-2.05
BQ106	12.60	3.09	-2.17
BQ107	12.70	2.94	-2.19
BQ108	12.80	3.07	-2.63
BQ109	12.90	3.20	-1.88
BQ110	13.00	2.06	-3.30

BQ111	13.10	2.77	-2.36
BQ112	13.20	2.90	-2.20
BQ113	13.30	2.53	-2.79
BQ114	13.40	2.63	-2.86
BQ115	13.50	2.60	-2.80
BQ116	13.60	2.26	-2.76
BQ117	13.70	2.53	-2.58
BQ118	13.80	2.80	-2.41
BQ119	13.90	2.16	-2.81
BQ120	14.00	2.24	-3.52
BQ121	14.10	2.73	-2.66
BQ122	14.20	2.87	-2.07
BQ123	14.30	2.65	-2.61
BQ124	14.40	2.60	-2.41
BQ125	14.50	2.83	-1.85
BQ126	14.60	3.04	-2.15
BQ127	14.70	2.74	-2.19
BQ128	14.80	2.78	-2.69
BQ129	14.90	2.93	-2.19
BQ130	15.00	2.68	-2.58
BQ131	15.10	2.73	-2.28
BQ132	15.20	2.61	-2.41
BQ133	15.30	1.75	-3.75
BQ134	15.40	2.09	-2.74
BQ135	15.50	2.90	-2.20
BQ136	15.60	2.73	-3.17
BQ137	15.70	3.13	-2.17
BQ138	15.80	2.77	-2.46
BQ139	15.90	2.50	-2.80
BQ140	16.00	2.73	-2.83
BQ141	16.10	2.28	-2.95
BQ142	16.20	2.40	-2.80
BQ143	16.30	3.15	-2.04
BQ144	16.40	2.95	-2.01
BQ145	16.50	3.00	-2.20
BQ146	16.60	2.86	-3.38
BQ147	16.70	2.74	-2.58
BQ148	16.80	2.30	-2.50
BC1	17.00	2.86	-2.01
BC2	17.20	2.63	-2.25
BC3	17.40	3.35	-1.71
BC4	17.60	2.19	-3.99
BC5	17.80	2.76	-2.00
BC6	18.00	2.83	-2.02

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18.40	2.63	-2.45
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19.00	2.99	-2.04
19.20	2.88	-2.12
19.40	1.91	-3.39
19.60	2.75	-2.89
19.80	2.38	-2.48
20.00	2.11	-2.64
20. 20	2.31	-2.33
20.40	2.83	-2.01
20.60	2.35	-2.51
20.80	2.60	-2.20
21.00	2.90	-1.70
21.20	2.90	-2.30
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22.00	3.10	-2.00
22.20	1.64	-3.39
22.40	2.85	-1.84
22.60	2.41	-1.99
22.80	1.95	-2.60
23.00	2.59	-2.32
23.20	2.71	-2.05
23.40	2.37	-2.66
23.60	2.22	-2.67
23.80	2.46	-2.09
24.00	1.97	-2.67
24.20	2.00	-3.26
24.40	1.76	-3.08
24.60	1.72	-3.56
24.80	2.18	-2.27
25.00	2.60	-2.11
25.20	2.31	-1.90
25.40	2.30	-2.10
25.60	2.50	-2.10
25.80	1.50	-3.20
26.00	2.20	-2.40
26.20	2.30	-2.40
26.40	1.90	-3.00
26.60	2.20	-2.40
26.80	2.40	-2.10
	18. 20 $18. 40$ $18. 60$ $18. 80$ $19. 00$ $19. 20$ $19. 40$ $19. 60$ $19. 80$ $20. 00$ $20. 20$ $20. 40$ $20. 60$ $20. 80$ $21. 00$ $21. 20$ $21. 40$ $21. 60$ $22. 20$ $22. 40$ $22. 60$ $23. 80$ $23. 00$ $23. 20$ $23. 40$ $23. 60$ $23. 80$ $24. 00$ $24. 20$ $24. 40$ $24. 60$ $25. 00$ $25. 20$ $25. 40$ $25. 60$ $25. 80$ $26. 00$ $26. 60$ $26. 80$	18.20 2.91 18.40 2.63 18.60 2.58 18.80 3.10 19.00 2.99 19.20 2.88 19.40 1.91 19.60 2.75 19.80 2.38 20.00 2.11 20.20 2.31 20.40 2.83 20.60 2.35 20.80 2.60 21.00 2.90 21.20 2.90 21.40 2.60 21.40 2.60 21.40 2.85 22.00 3.10 22.20 1.64 22.40 2.85 22.60 2.41 22.80 1.95 23.00 2.59 23.40 2.37 23.60 2.46 24.00 1.97 24.20 2.00 24.40 1.76 24.40 1.76 24.60 1.72 23.80 2.46 24.00 1.97 24.20 2.00 24.40 2.75 25.00 2.31 25.40 2.30 25.60 2.50 25.80 1.50 26.00 2.20 26.40 1.90 26.40 1.90 26.60 2.20 26.80 2.40

BB32	27.00	1.90	-3.50
BB33	27.20	1.90	-2.40
BB34	27.40	2.20	-2.10
BB35	27.60	2.10	-2.20
BB36	27.80	2.00	-2.00
BB37	28.00	1.08	-3.10
BB38	28.20	1.43	-2.72
BB39	28.40	1.28	-4.20
BB40	28.60	2.04	-1.71
BB41	28.80	1.19	-2.64
BB42	29.00	1.37	-2.49
BB43	29.20	1.85	-1.68
BB44	29.40	1.85	-1.76
BB45	29.60	1.20	-3.43
BB46	29.80	1.20	-2.74
BB47	30.00	0.79	-3.84
BB48	30.20	1.26	-2.94
BB49	30.40	1.88	-1.48
BB50	30.60	1.03	-3.09
BB51	30.80	1.91	-1.59
BB52	31.00	1.20	-2.03
BB53	31.20	1.67	-1.30

	4.025E-07
	5.254E-07
	3.631E-07
46.50	5.009E-07
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45.23	8.902E-07
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46.00	7.695E-07
42.50	5.186E-07
38.60	6.682E-07
38.03	4.202E-07
37.93	3.657E-07
35.03	2.080E-07
34.30	2.147E-07
34.73	3.765E-07
35.13	6.633E-07
36.93	5.951E-07
36.67	4.286E-07
35.97	5.479E-07
33.27	3.394E-07
33.80	2.895E-07
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38.83	1.685E-06
37.07	4.121E-07
37.20	4.478E-07
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32.83	1.131E-07
27.83	7.355E-08
27.47	6.266E-08
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38.10	2.949E-07
34.93	3.272E-07
33.60	2.733E-07
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39.73	6.468E-07
37.50	6.354E-07
36.43	3.181E-07
37.53	6.560E-07
36.43	5.149E-07
37.75	4.691E-07
36.87	6.317E-07
37.23	1.250E-06
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36.17	8.726E-07
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41.23	1.040E-06
40.40	4.219E-07
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37.60	4.814E-07
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32.80	5.227E-07
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35.23	1.171E-06
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32.77	1.659E-07
32.40	1.743E-07
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31.90	2.252E-07
34.00	2.357E-07
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37.43	2.987E-07
33.97	8.815E-07
33.27	4.217E-07
33.33	1.212E-07
33.30	2.574E-07
37.50	4.091E-07
38.97	5.970E-07
40.53	6.587E-07
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37.57	4.259E-07
37.93	4.043E-07
37.93	5.181E-07
37.93	5.260E-07
34.87	6.062E-07
35.27	3.980E-07
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33.33	3.512E-07
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29.27	1.423E-07
33.93	1.703E-07
32.97	2.520E-07
34.17	2.128E-07
31.53	2.524E-07
31.77	1.277E-07
34.83	2.954E-07
33.07	1.985E-07
33.07	1.772E-07
32.80	1.097E-07
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33.13	2.977E-07
29.87	4.749E-07
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40.83	3.351E-07
39.83	2.418E-07
35.60	3.027E-07
36.20	3.384E-07
35.70	3.167E-07
32.80	2.433E-07
34.70	6.668E-07
33.00	8.825E-07
33.37	3.087E-07

42	29
78	55
45	30
90	60

ton abundance



INV. 2016.192 Neocomites neocomiensis