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### Increasing the Effectiveness of Electrical Resistivity Tomography Using $\gamma_{11n}$ Configurations

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**ABSTRACT**

A new array type, the  $\gamma_{11n}$  arrays are introduced in this paper, in which the sequence of the current (C) and potential (P) electrodes is CPCP and the distance between the last two electrodes is  $n$  times the distance between the first two ones and that of the second and third one. These arrays are called quasi null arrays because they are – according to their array and behaviour – between the traditional and null arrays. It is shown by numerical modelling that in detection of small-effect inhomogeneities these configurations may be more effective than the traditional ones including the optimised Stummer configuration. Certain  $\gamma_{11n}$  configurations – especially the  $\gamma_{112}$ ,  $\gamma_{113}$  and  $\gamma_{114}$  – produced better results both in horizontal and vertical resolution investigations. On the basis of the numerical studies, the  $\gamma_{11n}$  configurations seem to be very promising in problems where the anomalies are similar to the numerically investigated ones, that is they can detect and characterise, for example, tunnels, caves, cables, tubes, abandoned riverbeds or discontinuity in a clay layer with greater efficacy than those of the traditional configurations.  $\gamma_{11n}$  measurements need less data than traditional configurations therefore also the time demand of electrical resistivity tomography (ERT) measurements can be shortened by their use.

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5 52 **Keywords:** geoelectric configuration,  $\gamma_{11n}$  configurations, depth of detectability, ERT, quasi null arrays6  
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9 54 **INTRODUCTION**10  
11 55 Geoelectric methods form a traditional group of geophysical techniques (Van Nostrand and  
12 56 Cook 1966; Alpin *et al.* 1966; Zhdanov and Keller 1993). In the early times their use was restricted to  
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14 57 mineral exploration. Today they are frequently used in numerous field problems (Butler 2005),  
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16 58 related to electrical resistivity distribution of the subsurface: hydrogeology (Kirsch 2006),  
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18 59 environmental studies (Ward 1990; Knödel *et al.* 2005), engineering (Ward 1990, Szalai *et al.*, 2009a),  
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20 60 safety purposes (Metwaly *et al.* 2008) and archaeological problems (Clark 1990), etc.21  
22 61 The number of published geoelectric arrays used for geoelectric measurements is more than  
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24 62 one hundred (Szalai and Szarka 2008a). It is widely known (mainly from Ward 1990) that each array  
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26 63 has some specific advantages and disadvantages. In studying these qualities, the arrays were  
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28 64 compared from many different aspects. One of the key parameters, the depth of the investigation  
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30 65 value was calculated by Szalai *et al.* (2009b) following the slightly different definitions given by  
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32 66 Edwards (1977) and Roy and Apparao (1971) for all arrays. Parameter sensitivity maps, which are  
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34 67 crucial in understanding the different arrays, were presented by Szalai and Szarka (2008b,c) for all  
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36 68 arrays that ever existed. Ward (1990) evaluated the geoelectric arrays from 14 various aspects.  
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38 69 Although the aforementioned investigations aimed at providing a theoretical basis for traditional  
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40 70 profiling and sounding techniques, they are also important for electrical resistivity tomography (ERT)  
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42 71 measurements because the individual arrays serve as a basis for the ERT measurements.43  
44 72 Since ERT measurements have become the dominant tool in geoelectric research in the past  
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46 73 decades, it is of crucial importance to maximize the information available when using them. There  
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48 74 are actually significant efforts to find the best possible, so-called optimized configurations (Furman *et*  
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50 75 *al.*, 2003; Stummer *et al.* 2004; Wilkinson *et al.*, 2006). The optimised configurations, e.g. the

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3 76 Stummer configuration (Stummer *et al.* 2004) – in contrast to the classical configuration approach –  
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5 77 may contain a series of very different arrays.  
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7 78 Stummer *et al.* (2004) did not however include in the optimisation procedure  $\gamma$ -type arrays,  
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9 79 and therefore not the  $\gamma_{11n}$  arrays. In the case of these arrays, the electrodes are positioned in an  
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11 80 overlapping mode that is the current, and potential electrodes follow each other alternately (see Fig.  
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13 81 1). The large value of the  $k$  geometrical factor does not, however, inevitably refer to the field  
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15 82 applicability of an array as shown by Szalai *et al.* (2002) and Szalai *et al.* (2004).  $k$  is namely the  
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17 83 function of the homogeneous half-space value which has nothing to do with the potential due to the  
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19 84 inhomogeneities which contain information important for us. If the  $\gamma_{11n}$  arrays will prove to be useful,  
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21 85 they have to be taken into account in all optimisation processes.  
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24 86 In the last few years several other motivations accumulated to study the  $\gamma_{11n}$  and  $\gamma_{m_{11n}}$   
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26 87 configurations which will be discussed in the next section. First of all however the definition of the  
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28 88 applied non-conventional arrays is given. The  $\gamma_{11n}$  arrays are presented in Figure 1. A  $\gamma_{m_{11n}}$  (mirrored  
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30 89  $\gamma_{11n}$ ) array contains a  $\gamma_{11n}$  array and its pair, a  $\gamma_{n_{11}}$  array. These arrays are the same but they are  
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32 90 orientated in opposite directions carrying out measurements. The  $\gamma_{m_{11n}}$  configuration which consists  
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34 91 of  $\gamma_{m_{11n}}$  arrays creates two times more dense data set than the  $\gamma_{11n}$  configuration in itself.  $\gamma(m)_{11n}$   
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36 92 refers to both the  $\gamma_{11n}$  and  $\gamma_{m_{11n}}$  configurations.  
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#### 43 **MOTIVATIONS TO STUDY THE $\gamma_{11N}$ AND $\gamma_{m_{11N}}$ CONFIGURATIONS**

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47 96 a.) Furman *et al.* (2003) performed a sensitivity analysis and demonstrated the supremacy of the  
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49 97 "partially overlapping arrays", which are also  $\gamma$ -type arrays, that is their electrode sequence is  
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51 98 CPCP. Szalai and Szarka (2008b) presented the normalised parameter sensitivity (nPS) maps  
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53 99 of many linear arrays. Many of them are reproduced in Figure 2. In the first row, the nPS  
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55 100 maps of the Wenner- $\alpha$  (W- $\alpha$ ), Wenner- $\beta$  (W- $\beta$ ), dipole-axial (Dp-ax) four-electrode and the  
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3 101 P-Dp three-electrode arrays are shown in a depth of one tenth of the array length (The array  
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5 102 length is the distance of the farthest electrodes which are not in the infinity). In the second  
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7 103 row the nPS maps of several characteristic  $\gamma_{n11}$  arrays are presented which are the oppositely  
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9 104 orientated versions of the  $\gamma_{11n}$  arrays therefore their nPS map is also the oppositely  
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11 105 orientated version of the those of the  $\gamma_{11n}$  arrays. The first one, the  $\gamma_{111}$  array (that is  $n=1$ ) is  
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13 106 in fact the Wenner- $\gamma$  array, a traditional array. The last one, the  $\gamma_{n11}$  array where  $n=\text{inf.}$  is a  
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15 107 null-array, the Midpoint-null or MAN array (Szalai *et al.* 2004). The  $\gamma_{811}$  array is one of the  
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17 108 series of the  $\gamma_{n11}$  arrays between the traditional and the null array. Similar arrays ( $n=1-7$ ),  
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19 109 which have similar nPS maps are investigated in this paper. Below each nPS map its maximal  
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21 110 value is shown. It is well seen that while the  $\gamma_{111}$  array's value is in the same order than those  
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23 111 of the values of the first row arrays the maximal values of the  $\gamma_{n11}$  nPS maps are drastically  
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25 112 increasing with increasing  $n$ . This high sensitivity motivated us to study the depth of  
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27 113 detectability (DD) value of these arrays.

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31 114 b.) The calculations of the depth of detectability (DD) values by Szalai *et al.* (2013b) has shown  
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33 115 that the DD values of the  $\gamma_{11n}$  configurations can be 2-2.5 times larger than that of the best  
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35 116 traditional configuration. A square resistive prism (e.g. 2 times 2m cross-section, 200  $\Omega\text{m}$   
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37 117 resistivity in a host of 100  $\Omega\text{m}$ ) proved to be detectable from 14m depth (upper side of the  
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39 118 prism) assuming 5% noise level and using  $\gamma_{113}$  configuration while it was detectable by the  
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41 119 best traditional configuration, the P-Dp one, only from 6.6m depth. In the investigation 100  
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43 120 electrodes were applied with 1m electrode distance. Applying these configurations one could  
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45 121 therefore get information from a larger depth which can be especially important in areas  
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47 122 where the space available for measurements is limited, e.g. in built-in areas.

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51 123 c.) It was shown comparing Szalai *et al.* (2011) and Szalai *et al.* (2013a) that the higher the DD  
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53 124 value of a configuration, the better are its imaging features. Among the traditional  
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55 125 configurations the DD value of the  $\beta$ -type configurations (Dp-Dp, P-Dp, Stummer, Wenner- $\beta$ ,

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3 126 electrode sequence is CCPP), was generally larger than those of the other configurations  
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5 127 (Szalai *et al.* 2011). According to the imaging capacity the sequence of the investigated  
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7 128 configurations was: Stummer, Dp-Dp, Wenner- $\beta$ , P-Dp, Wenner- $\alpha$ , P-P too (Szalai *et al.*  
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9 129 2013a). That is the ( $\beta$ -type) configurations which have larger DD values proved to have  
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11 130 better imaging capacity.

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13 131 Concluding from a., b. and c. one can say that larger  $nPS_{\max}$  value may lead to larger DD  
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15 132 values which may result in better imaging capacity. The high  $nPS_{\max}$  values of the  $\gamma_{11n}$   
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17 133 configurations could therefore result in good imaging features of these configurations.

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20 134 d.) Szalai *et al.* (2002) and Falco *et al.* (2013) have demonstrated that geometrical null arrays  
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22 135 (null arrays which provide zero signal in homogeneous half-space due to the appropriate  
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24 136 positioning of the electrodes) can be very effective in field conditions. There is only one  
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26 137 geometrical null array which can be built in 2D multielectrode systems, the MAN array (Szalai  
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28 138 *et al.* 2004 and Fig. 1). The applicability of the commercial software to invert its data is very  
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30 139 limited yet or not possible, while  $\gamma_{11n}$  configuration data may be inverted among certain  
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32 140 conditions. Due to the MAN array is a special case of this array type if  $n$  is infinite (see Fig 1)  
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34 141 the investigation of  $\gamma_{11n}$  arrays can be very useful to progress in understanding better the  
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36 142 MAN array, as well.

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40 143 e.) According to numerical calculations by Szalai *et al.* (2004) even the signal strength of the  $\gamma_{11n}$   
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42 144 arrays may be larger than that of the traditional arrays. About a dyke for example, with  
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44 145 increasing depth the size of the anomaly was 50  $\Omega\text{m}$ , 20  $\Omega\text{m}$  and 12  $\Omega\text{m}$  with the most  
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46 146 appropriate array lengths for the Wenner- $\alpha$  array, while it was about 60  $\Omega\text{m}$ , 36  $\Omega\text{m}$  and 32  
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48 147  $\Omega\text{m}$  for the MAN array, accordingly. It can be clearly seen that with the increasing depth of  
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50 148 the dyke the MAN array's signal strength became even better and better in comparison with  
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52 149 that of the traditional array. Even in field situation, in Finland, although the traditional  
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54 150 Wenner- $\alpha$  array had larger signal strength both the MAN (it was called Midpoint null array)

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3 151 and the Wenner- $\gamma$  null arrays produced larger anomalies due to horizontal resistivity changes  
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5 152 (Szalai *et al.* 2004).  
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7 153 On the basis of these experiences, the study of the  $\gamma(m)_{11n}$  configurations seems to be very  
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9 154 reasonable.  
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11 155 To understand these arrays it is very important to see that they used to produce very sharp  
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13 156 anomalies as can be expected on the basis of Figure 2 and as it was verified in both numerical  
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15 157 investigations and field measurements (Szalai *et al.*, 2004). It means that their anomalies are very  
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17 158 sensitive to the horizontal resistivity changes. Aside from one-dimensional investigations, where the  
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19 159 resistivity values supposed to change only in vertical direction this is one of the most important  
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21 160 factors of the imaging.  
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24 161 Unfortunately the commercial inversion softwares are not able to maintain such sharp  
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26 162 anomalies therefore the information which is contained in the measured data cannot be obtained.  
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28 163 This is why in this paper mostly inhomogeneities with small impact on the surface potential  
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30 164 distribution are studied. In this way the gradient of the signal will be not as large and therefore more  
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32 165 easily followed by the inversion. The other reason to investigate such inhomogeneities is that the  
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34 166 traditional arrays produce acceptable results for large impact inhomogeneities therefore for such  
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36 167 problems the application of other configurations is not required. In the future first a coarse image  
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38 168 could be obtained by a traditional configuration which could later serve as a priori model for  $\gamma(m)_{11n}$   
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40 169 configurations to refine the inverted images.  
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## 46 171 **NUMERICAL INVESTIGATIONS**

### 47 172 **The investigated configurations**

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50 173 The results obtained by the  $\gamma_{11n}$  (Fig. 1) and their mirrored version, the  $\gamma m_{11n}$  configurations  
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52 174 were compared in the paper with the results of traditionally used configurations like the dipole-  
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54 175 dipole (Dp-Dp), pole-dipole (P-Dp), Wenner- $\alpha$  and the optimised Stummer ones (Fig. 3).  
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3 176 In the investigation configurations with 60 electrodes were used with 1m electrode spacing  
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5 177 excluding the Stummer configuration. Because the Stummer configuration is available only for 30  
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7 178 electrodes (Stummer *et al* 2004) 2m electrode distance was used to get the same configuration  
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9 179 length as for the other configurations. The number of its data points is even in this case greater than  
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11 180 that of any  $\gamma(m)_{11n}$  configurations excluding only the  $\gamma_{m112}$  one (see Figs. 3 and 4). It has also to be  
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13 181 noted that in spite of the same change the Stummer configuration proved to be the best traditional  
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15 182 configuration in the investigations by Szalai *et al* (2013). It is also possible that the imaging quality of  
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17 183 the Stummer configuration could be further improved by using the same electrode distance as for  
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19 184 the other configurations, but it would lead to significant increase of the data number. Besides this  
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21 185 the imaging quality of the  $\gamma(m)_{11n}$  configurations could most likely be improved by combining them.

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24 186 The Dp-Dp configuration was used because it proved to be the best traditional configuration  
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26 187 in the investigations by Szalai *et al.* (2013a). The P-Dp was applied because it is a three-electrode  
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28 188 array like the MAN array and the  $\gamma_{11n}$  arrays themselves are getting closer and closer to be three-  
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30 189 electrode arrays with the increasing number of  $n$ . Wenner- $\alpha$  configuration was chosen because it is  
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32 190 one of the most popular and best known configurations, while the Stummer configuration (Stummer  
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34 191 *et al.* 2004) should have to be the best conventional configuration because it was constructed using  
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36 192 an optimisation process. Comparing the  $\gamma_{11n}$  and  $\gamma_{m11n}$  results with the results of these configurations  
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38 193 one can get therefore an oversight about the abilities of them.

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41 194 In Figure 6 the results of  $\gamma_{11n}$  configurations only for  $n=1-4$  are shown which gives satisfactory  
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43 195 information. In Figures 8 and 9 at the same time the whole series of  $\gamma_{11n}$  configurations ( $n=1-7$ ) is  
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45 196 presented to have an oversight about all these configurations.  $n$  is limited to 7 because its further  
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47 197 increase leads to too less data points.  $m$  was 1-14, 1-11, 1-9, 1-7, 1-7 and 1-7 for the  $\gamma_{11n}$   
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49 198 configurations for  $n=2-7$  (Fig. 1), accordingly.

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51 199 The parameters for the traditional configurations are seen in Fig. 3. The configurations used  
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53 200 in the optimised Stummer configuration can be found in Stummer *et al.* (2004).

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3 201 The data coverage and number of data points are seen in Figure 3 for the traditional and in  
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5 202 Figure 4 for the  $\gamma_{11n}$ ,  $\gamma_{n11}$  and  $\gamma_{m_{11n}}$  configurations ( $n=2-7$ ). While the Wenner- $\alpha$  and  $\gamma_{111}$  (Wenner- $\gamma$ )  
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7 203 configurations have only 570 data points, the Stummer configuration has 669, the Dp-Dp  
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9 204 configuration 736 the P-Dp configuration 871 ones. In contrary to these configurations the  $\gamma_{11n}$   
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11 205 configurations (Fig. 3) have no more than 420 data points, that is their measuring time is significantly  
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13 206 less, than those of the traditional configurations. Increasing  $n$  the number of data points is even  
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15 207 decreasing drastically. The mirrored version of the  $\gamma_{11n}$  configurations contain two times as many  
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17 208 point as its original version, but even in this case the number of the data points are only 840, 660,  
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19 209 540, 448, 392 and 342 for  $n=2-7$ , accordingly. It means that disregarding from  $\gamma_{m_{112}}$  configuration  
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21 210 even the mirrored configurations have less data points, that is their measuring time is shorter, than  
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23 211 that of the Stummer, Dp-Dp and P-Dp configurations (which used to produce the best results among  
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25 212 the traditional configurations according to Szalai *et al.* 2013a).  
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### 31 214 Inversion parameters

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33 215 All numerical calculations presented in this article were carried out by EarthImager, Version 2.1.6  
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35 216 (EarthImager, 2006). The parameters which are different from the software's basic parameters are  
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37 217 summarized in Table 1. The basic parameters were only changed if it was necessary to get reasonable  
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39 218 results. For e.g. the *Minimum Apparent Resistivity* parameters negative resistivity values were  
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41 219 selected, because the signal may change its sign. To create Figure 6 *Pseudosection* was applied in the  
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43 220 inversion process as *Starting Model* that is the section which contains the "measured" data. In all  
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45 221 inversion process 1% Gaussian noise was added to the data (with the exclusion of the inversion  
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47 222 whose results are presented in Figure 6) and RMS and L2 norm was used to study the data misfit. L2  
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49 223 norm is defined as the sum of the squared weighted data errors (difference between  
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51 224 predicted/calculated and observed/measured resistivities). The RMS (root mean squared) error is its  
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53 225 normalised version which takes into account also the data number.  
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3 226 Finite element method (FEM) was used in the modeling which is a numerical technique for  
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5 227 finding approximate solutions to boundary value problems for differential equations. It uses  
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7 228 variational methods to minimize an error function and produce a stable solution. FEM connects many  
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9 229 simple element equations over many small subdomains, named finite elements, to approximate a  
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11 230 more complex equation over a larger domain. FEM produces more accurate forward modeling  
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13 231 solution than the finite difference method.

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16 232 In many cases it would be possible to get better results than the ones presented by taking the  
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18 233 images from other iteration steps. The selection of the most appropriate inverted section requires  
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20 234 however knowledge of the model (which was used to calculate synthetic data) because the decrease  
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22 235 of the RMS does not inevitably result in better images. In field measurements this knowledge is  
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24 236 certainly not available although the final aim of such studies (including numerical investigations) is  
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26 237 checking the applicability of different configurations in the field. Therefore field data processing  
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28 238 requires a more or less automatic inversion. This is why *Stop RMS error option* was activated (values  
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30 239 shown in Table 1) which completed the inversion in many cases after the 1st or 2nd iteration. It was  
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32 240 also important to apply the same inversion parameters to all configurations to get (automatic)  
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34 241 objective results making it possible to compare them.

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37 242 RMS and L2 may not always be adequate to estimate the image quality. Often images with  
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39 243 smaller (that is better) RMS and L2 proved to be worse because they contained more pseudo  
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41 244 anomalies and more significant (therefore more disturbing) ones and the shape of the anomaly was  
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43 245 also further from the model. The RMS and L2 are severely influenced by the resistivity values  
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45 246 themselves and they may not be as sensitive to the geometrical parameters of the anomaly which is  
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47 247 often more important. Therefore we prefer to qualify the inverted image obtained from numerical  
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49 248 investigations similarly to that presented in "The criteria to interpret the results" section.

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53 249 For many  $\gamma(m)_{11n}$  configurations and for many models RMS proved to be high. There are two  
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55 250 main reasons:

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3 251 - The denominator  $d_i^{\text{Meas}}$  in the RMS function may be very small for many data, even close to  
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5 252 zero due to that the signal may change its sign (Szalai *et al.*, 2004). It leads to very large values for  
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7 253 individual measurements and therefore their sum, the RMS may also be large.

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9 254 - The numerator may also be large for the same data point due to the rapid changes of the  
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11 255 signal close to these small values. In a theoretical case if the predicted and measured curves (for a  
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13 256 given depth) would be the same but would be slightly shifted horizontally from each other one would  
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15 257 get very large RMS. This is not the case for the traditional configurations where the horizontal  
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17 258 gradient of the signal is much smaller.

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20 259 RMS and L2 are therefore not always appropriate values to estimate the quality of  $\gamma(m)_{11n}$   
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22 260 measurements. Another value should be found to quantitatively estimate the quality of  $\gamma(m)_{11n}$ .

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24 261 In the present article several models were numerically studied which aimed to illustrate 1.  
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26 262 the effect of the resistivity contrast to the inverted image; 2. the horizontal- and 3. vertical resolution  
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28 263 capacity of the different configurations; 4. the applicability of these configurations for larger effect  
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30 264 anomalies and 5. the applicability for a realistic model.

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### 34 35 266 **The criterias to interpret the results**

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37 267 In the qualification and comparison of the results obtained by different configurations the  
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39 268 main point (1a.) is whether the model body can be seen, that is whether is there an anomaly where  
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41 269 the body should appear. Only if there is an anomaly there is a sense to continue the interpretation. In  
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43 270 this case the next point (1b.) is whether are there any other (so-called pseudo) anomalies which are  
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45 271 not awaited to be there and how much they influence the interpretation. The larger they are in their  
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47 272 extension and/or in their resistivity contrast to the background value the more they can mislead the  
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49 273 interpretation. If the body is detectable in the following the horizontal (2a.) and vertical (2b.)  
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51 274 positions of the anomaly, its size (2c.), its resistivity value (2d.) and in ideal case its shape (2e.) can be  
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53 275 compared to the model parameters and be taken into account in the interpretation.  
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3 276 In our investigations, however, where the aim was to investigate small impact model  
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5 277 bodies the principal question had to be whether the anomaly due to the model will at least appear  
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7 278 on the inverted image. This was the case e.g. in Figure 6. In case if there are more model bodies their  
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9 279 separability (3.) can be an important item, too. This is the case in the resolution investigations in  
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11 280 Figures 8. and 9 and also in Figure 10. If the bodies are separated from each other in the inverted  
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13 281 image the same questions (1.a, b and 2.a,b,c,d,e) can be regarded like for the single bodies.  
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### 18 283 **Results of the numerical investigations and their interpretation**

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20 284 Before going into the details of the numerical simulations, we explain why  $\gamma_{m_{11n}}$  configuration  
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22 285 results will be presented instead of or beside  $\gamma_{11n}$  configurations. Figure 5 shows  $\gamma_{116}$ , and  $\gamma_{m_{116}}$   
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24 286 images for a model containing three conductive prisms. Their resistivities are 10  $\Omega\text{m}$  in the 100  $\Omega\text{m}$   
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26 287 resistivity half-space. It is easy to see that the  $\gamma_{m_{116}}$  configuration is able to separate the prisms from  
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28 288 each other better than the  $\gamma_{116}$  configuration by itself. Especially the prisms in the right side separate  
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30 289 from each other very well. The effect of the prism in the middle of the section is also more  
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32 290 remarkable in the  $\gamma_{m_{116}}$  image.  
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36 291 Investigations of further models not presented here motivated us to use the  $\gamma_{m_{11n}}$   
37  
38 292 configurations. They produced better images than the  $\gamma_{116}$  configurations, especially if the  
39  
40 293 inhomogeneities are in deeper parts of the model. This is the reason we prefer to use the mirrored  
41  
42 294 version of these configurations. In the present stage of the investigations, it seems, however, to be  
43  
44 295 reasonable to present also the  $\gamma_{116}$  results, at least for a few models.  
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46  
47 296 The results of DD investigations by Szalai *et al.* (2013b) and MAN configuration studies (Szalai  
48  
49 297 *et al.*, 2004) referred to the usefulness of the  $\gamma_{11n}$  configurations especially if the effect of the  
50  
51 298 inhomogeneity is small, that is if its size/depth ratio and/or its resistivity contrast to the host is small.  
52  
53 299 At first the image of a small size prism will be compared with those of the often applied dipole-dipole  
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55 300 (Dp-Dp), pole-dipole (P-Dp), optimised Stummer (St) and  $\gamma_{11n}$  ( $n=1-4$ ) configuration's images (Fig. 6).  
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3 301 While the resistivity of the host in Figure 6 was 100  $\Omega\text{m}$ , the resistivities of the prism were 500,  
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5 302 160 and 140  $\Omega\text{m}$  in the columns in Figure 6. The depth of the upper side of the prism is 3.9 m, its  
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7 303 thickness is 2.0 m, and its horizontal extension is 2.5 m, between 26.5 and 29 m. For the applied  
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9 304 inversion parameters see Table 1.

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11 305 If the resistivity of the prism was 500  $\Omega\text{m}$  (Fig. 6, left column), that is the resistivity contrast to  
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13 306 the background was significant the prism proved to be detectable by each configurations, however  
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15 307 the application of the  $\gamma_{111}$  configuration is not suggested due to the significant artefacts which are in  
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17 308 the same resistivity range than the “real” anomaly itself. The other configurations detect the model  
18  
19 309 clearly, they position it correctly and the pseudoanomalies were not comparable to the real anomaly.  
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21 310 The size of the anomaly is, however, larger than expected.

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23 311 The conventional configurations proved to be better for this model because the effect of the  
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25 312 anomaly is large due to the large resistivity contrast between the model and the background. For  
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27 313 such models they work properly, while the inversion of the  $\gamma(m)_{11n}$  configuration data for large effect  
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29 314 inhomogeneities is not well resolved yet.

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31 315 For the 160  $\Omega\text{m}$  prism (Fig. 6, middle column) only the  $\gamma_{113^-}$ , and  $\gamma_{114}$  configuration images are  
32  
33 316 somewhat convincing, while the traditional configurations proved to be rather ineffective. The  $\gamma_{113}$ ,  
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35 317 and  $\gamma_{114}$  images present a resistive anomaly at the right position which arises quite characteristically  
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37 318 from the background and the artefacts are smaller in their size than the anomaly. The  $\gamma_{114}$   
38  
39 319 configuration results are the most convincing from the whole series, although the anomaly is not at  
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41 320 all sharp in its case, neither.

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43 321 Finally the 140  $\Omega\text{m}$  prism (Fig. 6, right column) was detectable by all configurations excluding  
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45 322 the Stummer one, but the anomalies were more or less mispositioned. The  $\gamma_{113}$  and the  $\gamma_{114}$   
46  
47 323 configurations seem to be the closest to the real model, although they, too, produced significant  
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49 324 artefacts. **The most important is however the presence of an anomaly at the position of the**  
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51 325 **inhomogeneity which is inevitable for a correct interpretation.**  
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3 326 Artefacts could nevertheless easily mislead the interpretation of field data. To avoid  
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5 327 misinterpretation, however, there are several possibilities. If the location of the target is more or  
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7 328 less known even images with artefacts enable its more precise localization and description. In this  
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9 329 case the artefacts should not have taken into account. If there may be more prism-like objects and  
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11 330 there is not any information regarding their position it is possible: 1. to compare data of different  
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13 331 geoelectric configurations. If images of many configurations display an anomaly at the same position,  
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15 332 it is highly probable that there is an inhomogeneity. If an anomaly appears on only 1-2 images, it is  
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17 333 most likely an artefact. An anomaly with a large value and extension (e.g., on the  $\gamma_{114}$  image in the  
18  
19 334 right column between about 34 and 44 m) does not have any pair on the  $\gamma_{112}$  or  $\gamma_{113}$  images.  
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21 335 Therefore its validity is strongly questioned. In contrast, the anomaly in the middle of the section  
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23 336 appears on all of these images (although not exactly at the same position) increasing the probability  
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25 337 of the existence of an inhomogeneity there. The MOST algorithm (Leontarakis and Apostolopoulos  
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27 338 (2012, 2013) basically uses the same principle. In this process, artefacts due to random noises  
28  
29 339 eliminate each other while the anomalies due to real objects strengthen each other (see Fig. 7,  
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31 340 discussed later more detailed). 2. Similar procedures can be applied carrying out measurements  
32  
33 341 several times with the same configuration (stacking). 3. The comparison of geoelectric results with  
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35 342 results of other geophysical measurements, or the joint inversion of different data sets could also  
36  
37 343 decrease the uncertainty. 4. A direct investigation at the problematic places is also possible through  
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39 344 excavations or boreholes. Which of these procedures is applied it is a question of money.

345 While for the models with high resistivity contrast (500  $\Omega$ m) the traditional configurations  
346 proved to be better their quality decreases faster with the decreasing resistivity contrast than that of  
347 the  $\gamma_{11n}$  configurations. In the small resistivity contrast range the application of the  $\gamma_{11n}$  configurations  
348 seem to be more worthwhile and the quality of their image can be further improved by the Model  
349 Stacking (MOST) procedure introduced by Leontarakis and Apostolopoulos (2012, 2013). Stacking the  
350 models of different configurations the model of the combined configuration leads to a final model

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3 351 almost free of artefacts with extremely high resolution in shape and positioning, and an intense  
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5 352 representation of the targets. The Model Stacking procedure is based on a simple statistical  
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7 353 approach, calculating the geometric mean of the different values, which are given by each model for  
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9 354 the same point of the half-space (Leontarakis and Apostolopoulos, 2012, 2013).

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11 355 The RMS error values for the images in Figure 6 were between 4.2 and 5.2% for all  
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13 356 configurations, that is in this sense there was not any significant difference among the investigated  
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15 357 configurations.

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18 358 **Figure 7 presents the effect of the MOST procedure.** In the first row ~~in Figure 7~~ the MOST  
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20 359 results made of the combination of the P-Dp and Dp-Dp configuration results are shown for both the  
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22 360 160  $\Omega\text{m}$  and 140  $\Omega\text{m}$  prism models. This configuration combination contains 1540 data points. For  
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24 361 the 160  $\Omega\text{m}$  prism the MOST result is more convincing than the results of the simultaneous  
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26 362 configurations (Fig. 6), but it still contains a lot of artefacts. The MOST result made of the  $\gamma_{112}$ ,  $\gamma_{113}$  and  
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28 363  $\gamma_{114}$  configurations (Fig. 7, second row) is much more convincing in spite of that this combination  
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30 364 contains only about 30% less, 1020 data points. The resistivity values of the artefacts are in this case  
31  
32 365 not comparable with that of the real anomaly and it became sharper and more characteristic than in  
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34 366 the individual images (Fig. 6). The quality of the image could be even further improved by stacking all  
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36 367 of these configurations (Fig. 7, third row), but thus the measurement becomes less economic.

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39 368 The situation is about the same for the 140  $\Omega\text{m}$  prism model (Fig. 7, right column). The MOST  
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41 369 procedure led to reasonable image for both (traditional-, and  $\gamma_{11n}$ -) configuration combinations. In  
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43 370 this case even the combination of all configurations (Fig. 7, third row) seem to be reasonable if the  
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45 371 aim is to get high quality image even among very wrong conditions. Disregarding from the artefact at  
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47 372 the end of the profile the prism was clearly detected in this way and its all geometrical parameters  
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49 373 that are its horizontal and vertical positions and even its size are satisfactory.

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3 374 If artefacts do not disappear even after carrying out the MOST procedure it is still possible to  
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5 375 use other geophysical techniques or apply direct procedures to decrease the uncertainty of the  
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7 376 interpretation.

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9 377 Summarising the results of Figures 6 and 7 it can be stated that while the traditional Dp-Dp-,  
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11 378 and Stummer configurations proved to be well usable if the resistivity contrast was larger (500  $\Omega$ m),  
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13 379 the  $\gamma_{11n}$  configurations proved to be more and more fruitful in comparison with the traditional  
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15 380 configurations if the contrast was smaller (160, 140  $\Omega$ m). The advantageous features of the  $\gamma_{11n}$   
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17 381 configurations became especially spectacular by combining them using the MOST procedure. The  
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19 382 application of the  $\gamma_{11n}$  configurations and applying the MOST procedure is therefore highly  
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21 383 recommended in case if a small impact anomaly should be found in a noisy environment.  
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25 384 In the proceeding Figures only Wenner- $\alpha$  and Stummer configuration results will be shown  
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27 385 from the traditional ones, because we wanted to compare our results with the results of a popular  
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29 386 traditional configuration (W- $\alpha$ ) and with that of the best traditional configuration (St). At this stage of  
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31 387 the studies we found important to present also the  $\gamma_{11n}$  configuration results beside of the  $\gamma_{m_{11n}}$   
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33 388 configuration ones.  
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36 389 Figure 8 presents the horizontal resolution capacity of the Wenner- $\alpha$ , Stummer-,  $\gamma_{11n}$ -, and  
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38 390  $\gamma_{m_{11n}}$  ( $n=1-7$ ) configurations. For the applied inversion parameters see Table 1. The model  
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40 391 parameters are presented on the top of Figure 8. The Wenner- $\alpha$  configuration was unable to  
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42 392 separate the conductive prisms from each other. The Stummer configuration clearly separates the  
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44 393 right hand prism from the others and the separateness of the second prism from the right side may  
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46 394 be supposed, too. The  $\gamma_{11n}$ -, and  $\gamma_{m_{11n}}$  configurations (from  $n=2$ ) separate the prism farthest on the  
47  
48 395 right from the others even more convincingly creating a high resistivity region (29-38m) between the  
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50 396 prisms. With increasing  $n$  the second prism from the right side separates itself more convincingly  
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52 397 from the other prisms (there is again a high resistivity zone between 18m and 22m). The first two  
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54 398 bodies in the left side - whose distance is comparable to their depth (4m versus 4.9m) - could not  
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3 399 have been separated from each other by neither of the studied configurations. From the point of  
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5 400 view of their horizontal resolution capacity, both the  $\gamma_{m_{11n}}$ , and the  $\gamma_{11n}$  configurations proved to be  
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7 401 definitely better than even the optimised Stummer configuration, not speaking about the Wenner-  
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9 402  $\alpha$  configuration. The RMS value was below 2% for each configurations.

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11 403 Figure 9 demonstrates the results of vertical resolution investigations for the same  
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13 404 configurations. For the applied inversion parameters see Table 1. The model parameters are given on  
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15 405 the top of Figure 9. All prisms closer to the surface were detected by each configuration. The  
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17 406 Stummer and the  $\gamma(m)_{11n}$  ( $n=1,2$ ) configurations proved to be almost perfect regarding all quality  
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19 407 parameters. The near-surface anomalies of the W- $\alpha$  and the other  $\gamma(m)_{11n}$  configurations are not as  
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21 408 sharply delineated, but they are satisfactory, too.

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24 409 For the prisms on deeper levels, only the one on the right side was observed by all  
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26 410 configurations (excluding only the  $\gamma_{117}$  configuration), but it merged into the one above it because  
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28 411 they are too close to each other.

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31 412 From the prism pair on the left side, the deeper one proved to be almost undetectable even by  
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33 413 the Stummer configuration. In contrary, most  $\gamma(m)_{11n}$  results refer to the existence of the deeper  
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35 414 prism. They show a long, narrow anomaly downwards (e.g.  $\gamma_{115}$ ,  $\gamma_{116}$  and  $\gamma_{m_{117}}$  configurations) or even  
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37 415 an anomaly which delineates well the prism pair below each other (e.g.  $\gamma_{117}$ ,  $\gamma_{m_{113}}$  and  $\gamma_{m_{114}}$   
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39 416 configurations). The W- $\alpha$  configuration indicates the deeper anomaly, as well, but with a very wide  
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41 417 and uncertain anomaly.

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44 418 The Stummer configuration indicates weakly the existence of the deeper prism in the middle of  
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46 419 the section by, but the anomalies produced by the  $\gamma_{113}$ ,  $\gamma_{116}$ ,  $\gamma_{m_{116}}$ ,  $\gamma_{m_{117}}$  configurations and especially  
47  
48 420 that of the  $\gamma_{m_{113}}$  configuration are more convincing. Their anomalies are narrower and/or get closer  
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50 421 to the depth of the deeper prism and/or the values of the anomalies differs more from the  
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52 422 background value.

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3 423 For this model the RMS values of the  $\gamma_{11n}$ - $\gamma_{m_{11n}}$  configurations from  $n=3$  were systematically  
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5 424 much larger (20-32%) than those of the traditional configurations (below 2%). As it has however  
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7 425 already been mentioned the principal question is the similarity of the inverted image to the reality  
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9 426 which is the model in numerical investigations. RMS value seems increase with increasing  $n$  in this  
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11 427 model. It may happen because with increasing  $n$  the arrays approach the null array situation and  
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13 428 produce sharper and sharper anomalies.

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16 429 The separation of the prisms below each other was impossible for all the configurations we  
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18 430 studied, but certain  $\gamma_{11n}$ - and  $\gamma_{m_{11n}}$  configurations and especially the  $\gamma_{m_{113}}$  configuration proved to be  
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20 431 better in detection of the deeper bodies.

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22 432 Next we investigated a model which did not seem to be favourable for the  $\gamma_{m_{11n}}$   
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24 433 configurations (Fig. 10) because of the large size of the inhomogeneities (inversion parameters are in  
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26 434 Table 1.).

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29 435 The model in the left column of Figure 10 is very similar to the one studied by Wilkinson *et al.*  
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31 436 (2006). Here the anomalous bodies with large resistivity contrast (100  $\Omega\text{m}$  in comparison to the 10  
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33 437  $\Omega\text{m}$  half-space value) and the large size can be seen better in the Stummer image than in most  $\gamma_{m_{11n}}$   
34  
35 438 configuration images. Although the deepest (from the detectability point of view) most critical body  
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37 439 is displayed more convincingly by the  $\gamma_{m_{11n}}$  ( $n=2-6$ ) configurations these images contain several  
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39 440 pseudoanomalies, as well. In such a case again a solution similar to the one applied by Leontarakis  
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41 441 and Apostolopoulos (2012, 2013) could be suggested to suppress the pseudoanomalies and highlight  
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43 442 the real anomalies. We would like to call your attention also to the  $\gamma_{m_{112}}$  configuration which gives -  
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45 443 in spite of its higher RMS value - the most characteristic image of the prisms disregarding from the  
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47 444 pseudoanomaly in the left bottom part of the section.

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50 445 If however the near-surface bodies are not present (Fig. 10 middle column), all  $\gamma_{m_{11n}}$   
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52 446 configurations give better results than the Stummer configuration. They can separate the two deeper  
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54 447 bodies from each other more impressively. It is especially true if  $n$  is at least 3. The  $\gamma_{m_{11n}}$  images

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3 448 remain as impressive even in the presence of near-surface bodies if they do not influence to the  
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5 449 surface potential too much (Fig. 10 right column). It is remarkable that also the near-surface bodies  
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7 450 are presented more convincingly by the  $\gamma_{11n-7}$ , than by the traditional configurations. The  $\gamma_{m116}$   
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9 451 configuration produced the best image which resembles the best to the model. It separated all  
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11 452 prisms unambiguously from each other, their horizontal positioning was perfect, their vertical  
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13 453 positioning was reasonable, like that of their shape. Of course due to their limited effect to the  
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15 454 surface potential the resistivity values of the anomalies may not be very good.

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18 455 We found that  $\gamma_{11n}$  and  $\gamma_{m11n}$  configurations may be more productive even in the  
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20 456 investigation of bodies which have larger effect to the potential distribution.

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22 457 At last Figure 11 demonstrates a realistic example (inversion parameters are in Table 1): a  
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24 458 hole in the liner on the bottom of a waste deposit. On the basis of the former results we did not find  
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26 459 important to show the  $\gamma_{111}$  result in this case. The liner's resistivity was supposed to be 10000  $\Omega\text{m}$ ,  
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28 460 while the background's was supposed to be 100  $\Omega\text{m}$ . The liner on the bottom of the waste deposit  
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30 461 used to be namely a kind of plastic which has a very high resistivity value. The hole was supposed -  
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32 462 for simplicity reasons - to have the same resistivity like that of the "waste" itself and the rock below  
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34 463 the waste deposit. It is a simple model for the given situation, but it is able to handle the main point  
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36 464 of the problem, the detectability of the hole and its positioning.

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38 465 In this case the fundamental question is whether there is a hole in the liner. Regarding this  
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40 466 question all configurations with the exception of the  $W-\alpha$  proved to be satisfactory, because all of  
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42 467 them presented a conductive anomaly close to the expected position which refers to the existence of  
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44 468 a hole in the resistive liner. Regarding the second most important question, the horizontal position of  
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46 469 the hole, in case of the  $\gamma_{m11n}$  configurations ( $n=5-7$ ) the hole is although horizontally in the middle of  
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48 470 the conductive anomaly, but its horizontal extension is much larger than that of the body's. The  
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50 471 Stummer configuration indicates a discontinuity in the resistive layer, but it is strongly mispositioned  
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52 472 and much wider than the hole that is the Stummer configuration was not able to localize it precisely.  
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54 473 In contrary the  $\gamma_{m11n}$  configurations ( $n=2-4$ ) and especially the  $\gamma_{m112}$  and  $\gamma_{m113}$  ones produced narrow

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3 474 anomalies at the right location. These configurations seem therefore to be convenient to detect a hole  
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5 475 and to localize it to fulfill the most principal tasks.

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7 476 Although it is not important from the point of view of the given problem, we note that the  
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9 477 segment below the liner is resistive in the inverted section due to the fact that the current is not able  
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11 478 to penetrate below the resistive liner. Regarding it differently the whole bottom part of the section  
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13 479 below the liner would be most likely below the DOI (depth of investigation) level introduced by  
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15 480 Oldenburg and Li (1999) because the DOI level used to be closer to the surface where there are  
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17 481 bodies with large resistivity contrast to the average values. It refers in turn to the fact that below this  
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19 482 level surface data are insensitive to the value of the physical property of the earth.

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## 22 484 CONCLUSIONS

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24 485 A new configuration type, the  $\gamma_{11n}$  configurations are introduced, which have not been  
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26 486 investigated yet. Our numerous former studies let us assume that such so-called quasi null  
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28 487 configurations can be very useful complements to the traditional configurations.

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31 488 Numerical investigations showed that although models which have large impact to the surface  
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33 489 potential were presented better by the traditional configurations the quality of their image decreases  
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35 490 faster with decreasing model impact than that of the  $\gamma_{11n}$  configurations. For small impact models the  
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37 491 application of the  $\gamma_{11n}$  configurations is worthwhile at least together with a traditional configuration.  
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39 492 It was shown that the quality of the image of  $\gamma(m)_{11n}$  configurations can even be further improved by  
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41 493 the Model Stacking procedure resulting a good image even for small-effect models. **If the application  
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43 494 of the MOST procedure does not help avoid the uncertainties, then a combination of the geoelectric  
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45 495 results with results of other geophysical investigations, or direct investigation of the possible  
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47 496 inhomogeneities is recommended.**

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51 497 Most of the  $\gamma_{11n}$  configurations proved to be definitely better than those of the traditional  
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53 498 configurations in horizontal resolution investigations (especially with larger  $n$  values.) The mirrored

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3 499 version of the  $\gamma_{11n}$  configurations (the  $\gamma_{m_{11n}}$  configurations) were even better than the original  
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5 500 configurations.

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7 501 Although in the vertical resolution studies the separation of the anomalies directly below each  
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9 502 other proved to be impossible for all studied configurations, certain  $\gamma(m)_{11n}$  configurations and  
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11 503 especially the  $\gamma_{m_{113}}$  configuration proved to be good in detection of the deeper bodies.

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13 504 For certain models  $\gamma_{m_{11n}}$  configurations may be better even in case of large-impact  
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15 505 inhomogeneities, as it was illustrated, too. The  $\gamma_{m_{11n}}$  configurations proved to be better than even  
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17 506 the Stummer configuration also in detection and positioning of a hole in the liner in a realistic field  
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19 507 example.

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21 508 Summarising the numerical results it can be stated that the  $\gamma(m)_{11n}$  configurations are more  
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23 509 sensible to small impact models, than the traditional configurations, including the optimised  
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25 510 Stummer configuration, giving better image about them. They proved to have better horizontal  
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27 511 resolution, as well. In case of model bodies below each other many of them were able to indicate  
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29 512 the existence of the lower body and even its vertical position in contrary to the traditional  
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31 513 configurations. It is in accordance with the larger depth of detectability values of these  
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33 514 configurations which were calculated by Szalai *et al.* (2014). These statements are right in spite of the  
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35 515 smaller data coverage of these configurations which could however be increased by the  
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37 516 simultaneous use of different  $\gamma(m)_{11n}$  configurations. Applying the Model Stacking procedure by  
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39 517 combining the images of several  $\gamma(m)_{11n}$  configurations the results could even be further improved.

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41 518 In the present paper we mostly concentrated on models which seem to be most promising for  
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43 519 the  $\gamma_{11n}$  and  $\gamma_{m_{11n}}$  configurations, according to our theoretical considerations. On the basis of these  
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45 520 investigations we propose that problems like detection and characterisation of tunnels, caves, cables,  
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47 521 tubes, abandoned riverbeds, lack of continuity in clay layers could be effectively solved by these  
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49 522 configurations. Their use is also recommended in problems where false alarms are less important  
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51 523 than high resolution, e.g. in dam investigations or waste deposit monitoring. They can be useful in  
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524 any problems where small changes are expected with time, e.g. in any monitoring problems. Due to  
525 the reduced effect of the inhomogeneity below conductive or resistive layers,  
526 the  $\gamma(m)_{11n}$  configurations should be effective in such problems.

527 They can be especially productive in comparison to other configurations in areas where the  
528 space available for measurements is limited.

529 The time required for measurements with the  $\gamma(m)_{11n}$  configuration is moreover less than that  
530 of the traditional configurations, because disregarding from the  $\gamma m_{112}$  configuration they contain less  
531 data points, than the traditional configurations. The combined application of  
532 different  $\gamma(m)_{11n}$  configurations is rather economic and it can highly improve the efficiency of the  
533 measurements. A combination with traditional array results or with results of other geophysical  
534 measurements can also make it very straightforward to get the best possible interpretation.

535 We think on the basis of the presented investigations that the  $\gamma(m)_{11n}$  configurations might give  
536 significant contribution to the geoelectric method. Their further study is therefore highly  
537 recommended.

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22 608 **FIGURE CAPTIONS**

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27 610 **Figure 1:** The  $\gamma_{11n}$  arrays. Stars denote current, circles denote potential electrodes;  $a$  is the electrode  
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29 611 distance.

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31 612 **Figure 2:** The normalised parameter sensitivity (nPS) maps of several traditional and typical  $\gamma_{n11}$   
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33 613 arrays in a depth of one tenth of the array length. Stars denote current, circles denote  
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35 614 potential electrodes. Below the maps their maximal values can be seen.

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37 615 **Figure 3:** Left side: the applied traditional configurations with their parameters. Stars denote current  
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39 616 electrodes, full circles potential electrodes. Right side: Data coverage and number of data  
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41 617 points for the same configurations.

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44 618 **Figure 4:** Data coverage and number of data points for the  $\gamma_{11n}$ ,  $\gamma_{n11}$  and  $\gamma_{m_{11n}}$  configurations ( $n=2-7$ ).

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46 619 **Figure 5:** Example to show the advantage of the mirrored configurations against the single ones. The  
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48 620 model is given on the top of the figure. The resistivity of the prisms is 10  $\Omega\text{m}$  while the half-  
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50 621 space resistivity is 100  $\Omega\text{m}$ .

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3 622 **Figure 6:** Inverted sections for several traditional-, and  $\gamma_{11n}$  configurations. Resistivities of the prisms  
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5 623 are 500, 160 and 140  $\Omega\text{m}$ . Background resistivity is 100  $\Omega\text{m}$ . The model is given on the top of  
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7 624 the figure.

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10 625 **Figure 7:** MOST images of different configuration combinations

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12 626 **Figure 8:** Horizontal resolution investigation for different traditional-, and  $\gamma_{11n}$  configurations. The  
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14 627 model is given in the first row.

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16 628 **Figure 9:** Vertical resolution investigation for different traditional-, and  $\gamma_{11n}$  configurations. The model  
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18 629 is given in the first row.

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20 630 **Figure 10:** Left column: Inversion results from the Wenner- $\gamma$ , Stummer and  $\gamma(m)_{11n}$  ( $n=1-7$ )  
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22 631 configurations for the model similar to that in Wilkinson *et al.* (2006). Middle column: results  
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24 632 for the same model without the near-surface anomalous bodies. Right column: the first model  
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26 633 with smaller near-surface inhomogeneities. The model is given in the first row.

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29 634 **Figure 11:** The effect of a hole in the liner at the bottom of a waste deposit from different  
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31 635 configurations. The uppermost Figure is the model we investigated.

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### 34 35 637 **TABLE CAPTIONS**

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38 638 **Table 1** The parameters applied in the numerical investigations which are different from the basic  
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40 639 (Surface) parameters of the EarthImager v2.1.6. software. The parameters different from the basic  
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42 640 ones are written in bold.

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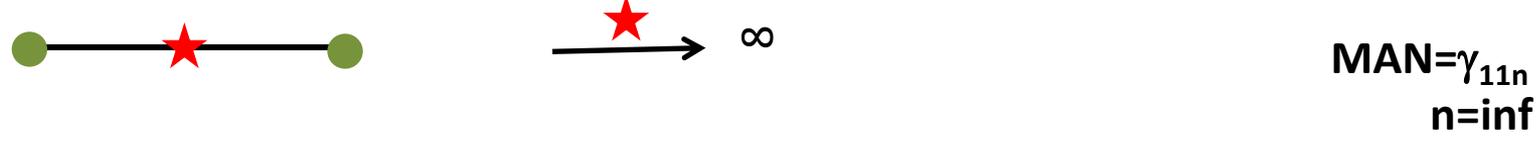
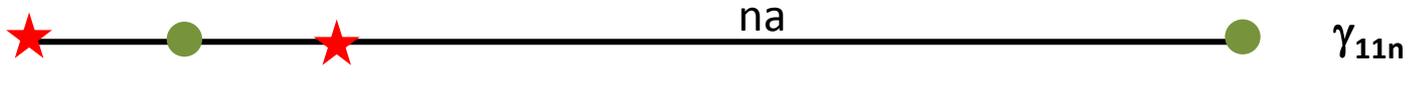
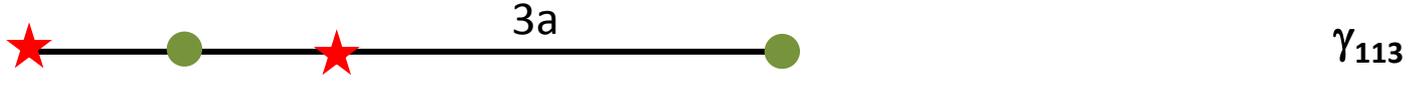
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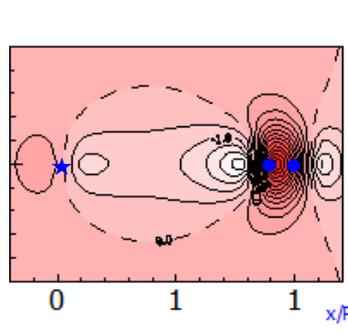
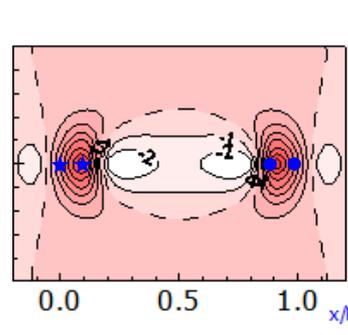
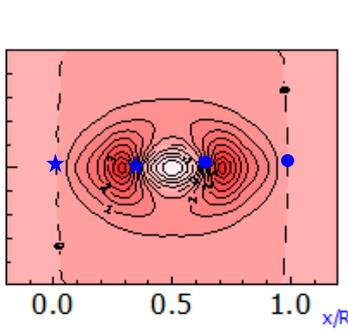
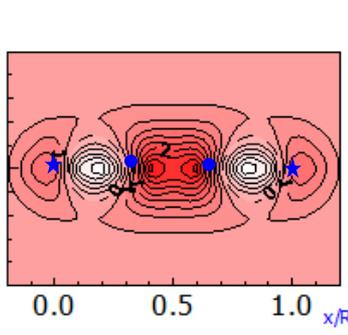
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W- $\alpha$

W- $\beta$

DP-ax

P-DP



maximal value

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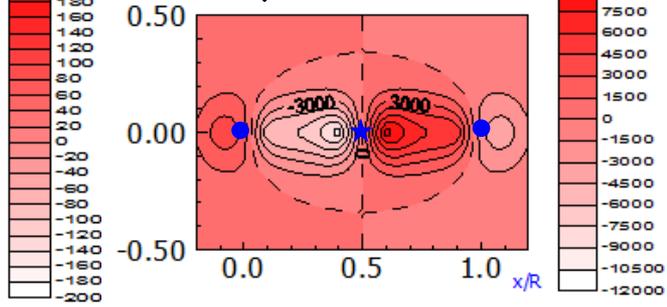
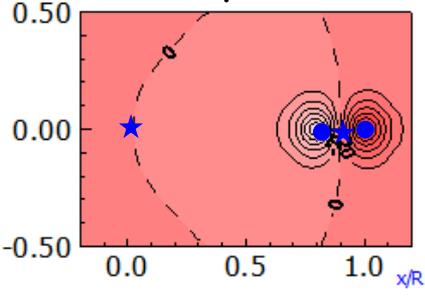
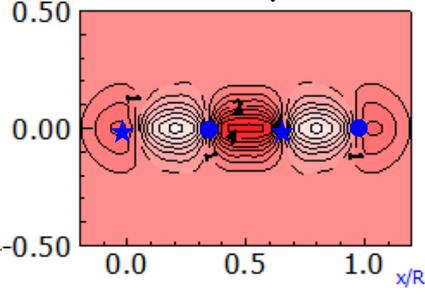
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W- $\gamma$

$\gamma_{811}$

$\gamma_{n11}$ =MAN



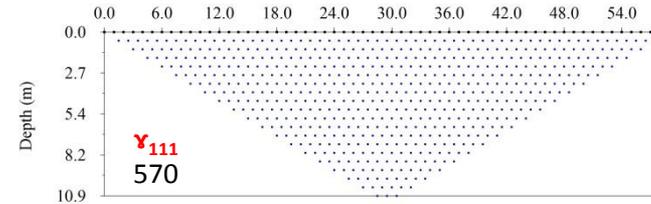
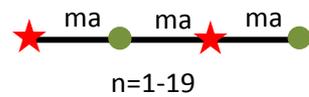
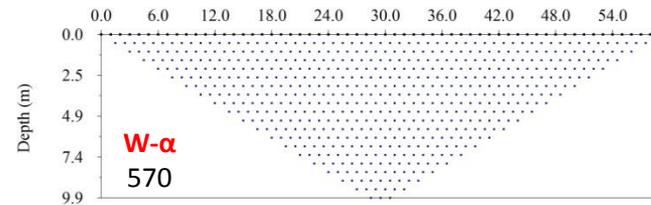
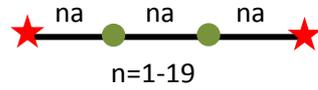
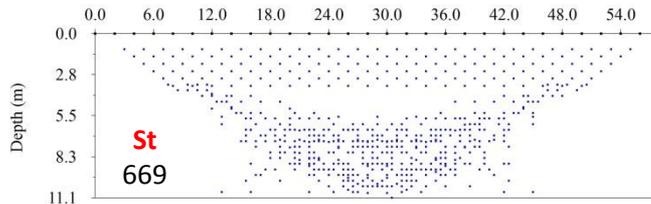
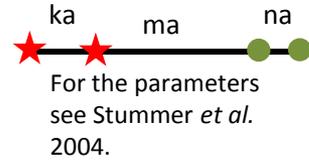
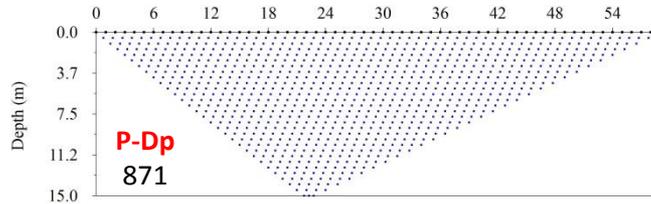
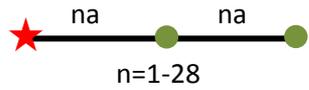
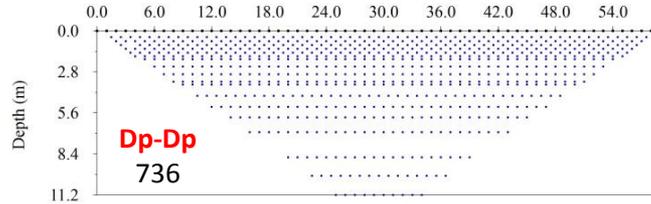
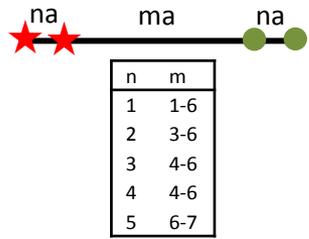
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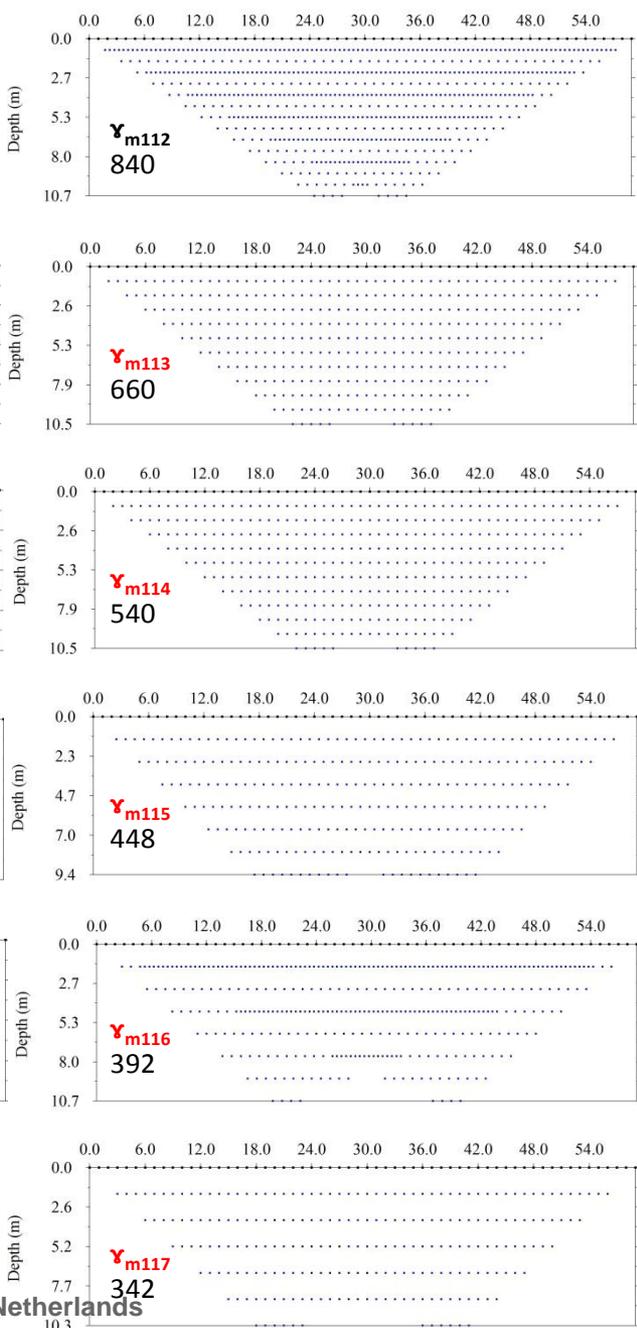
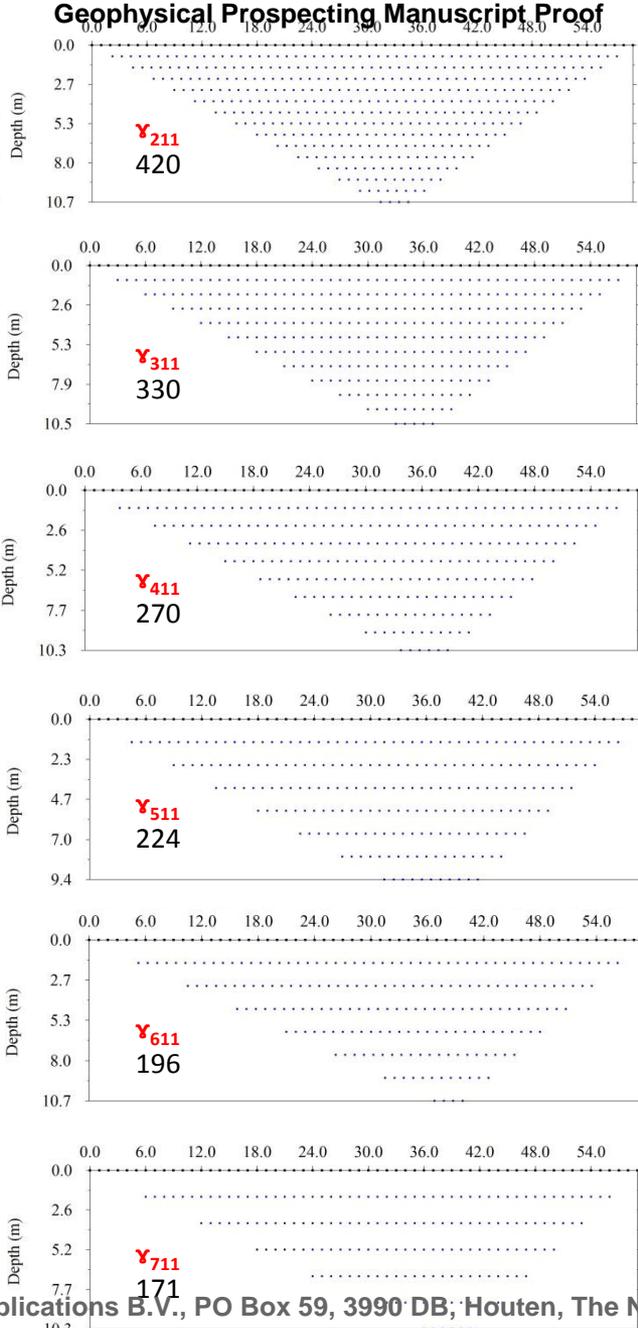
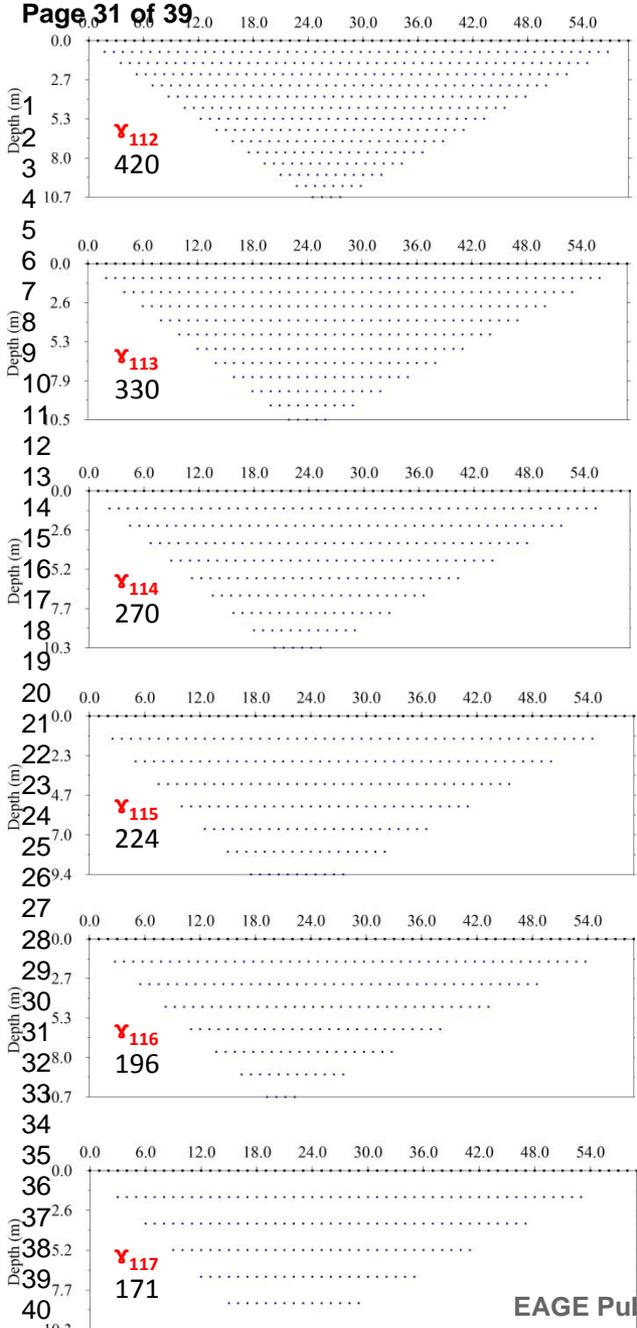
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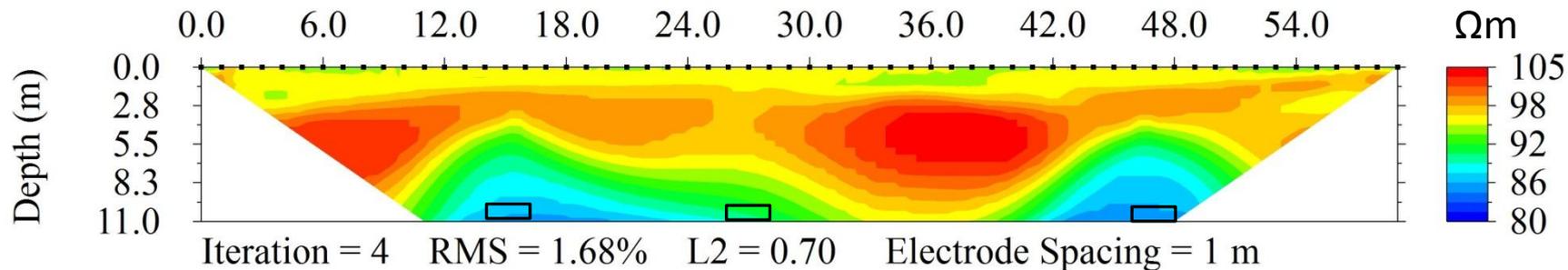
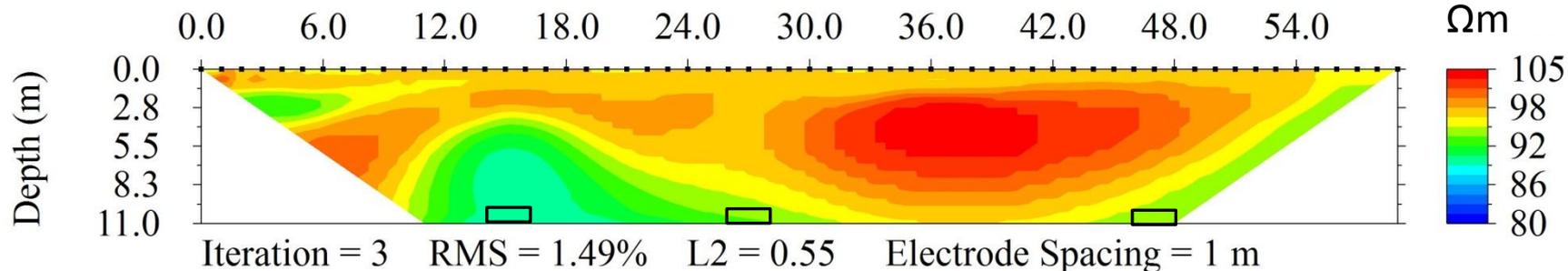
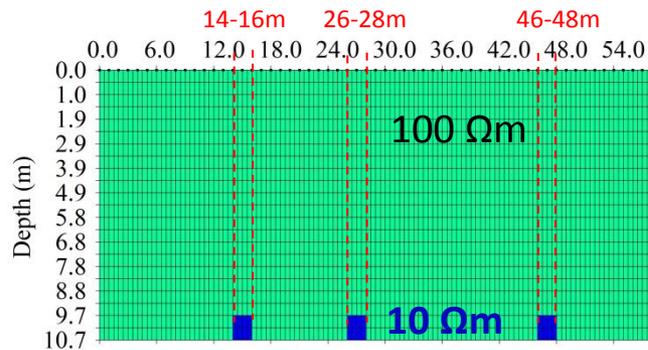
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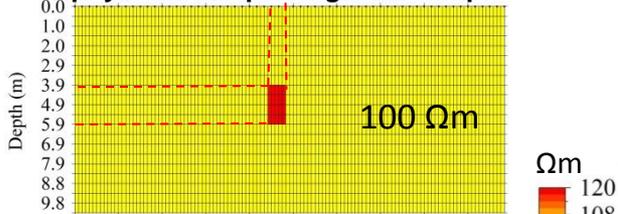
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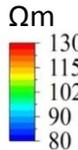


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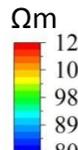


Resistivity of the prism:

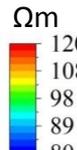
500 Ωm



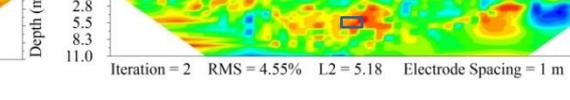
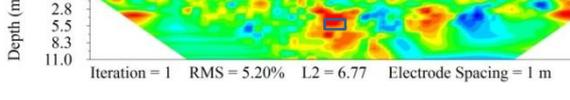
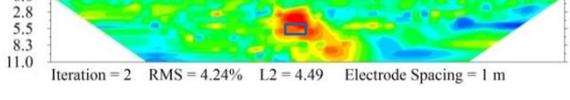
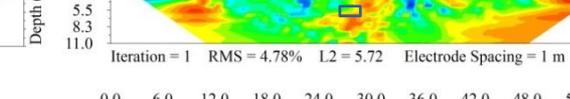
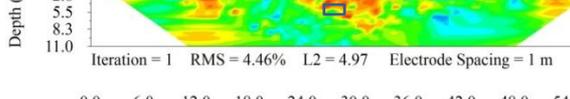
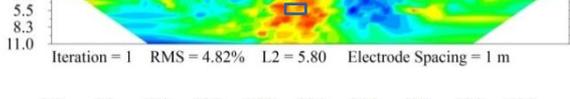
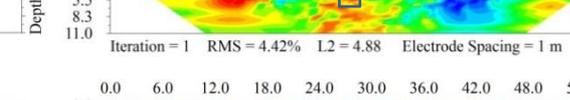
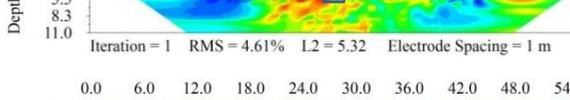
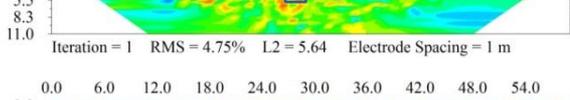
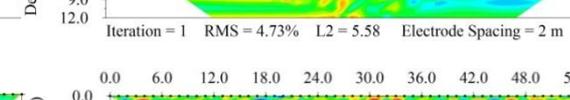
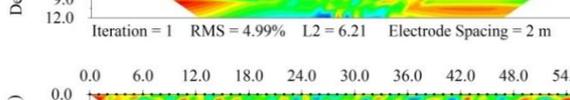
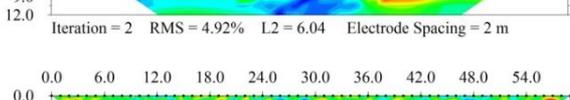
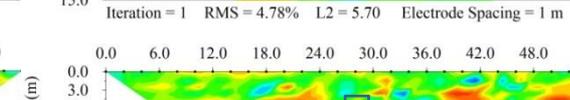
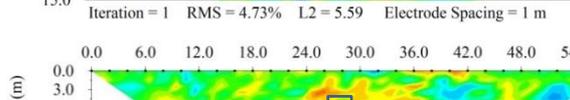
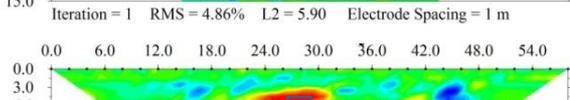
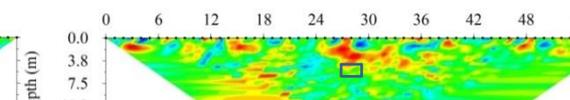
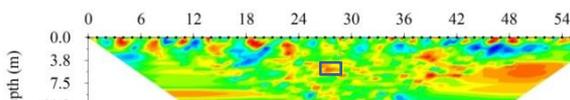
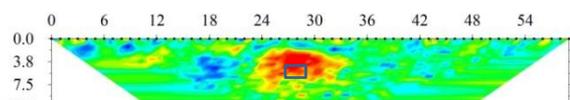
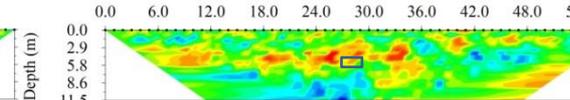
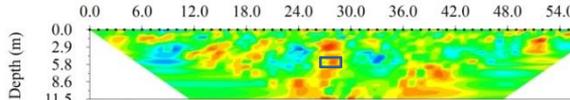
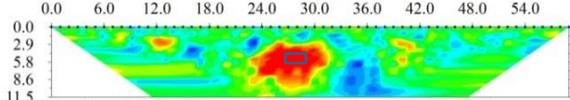
160 Ωm



140 Ωm



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Resistivity of the prism

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140  $\Omega\text{m}$

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9 Dp-Dp

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14  $\gamma_{113}$

15  $\gamma_{114}$

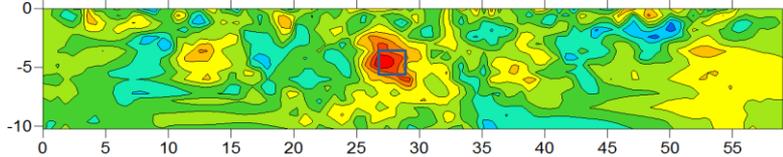
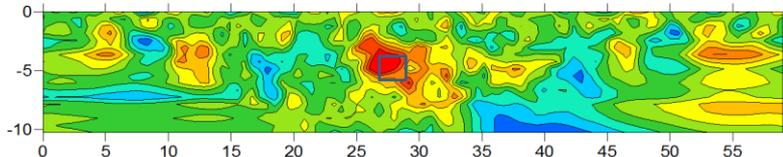
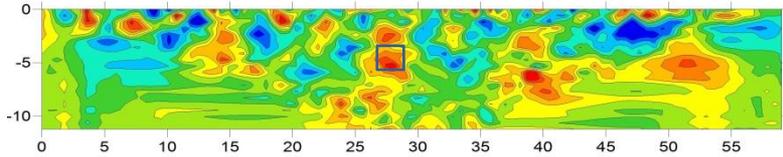
19 P-Dp

20 Dp-Dp

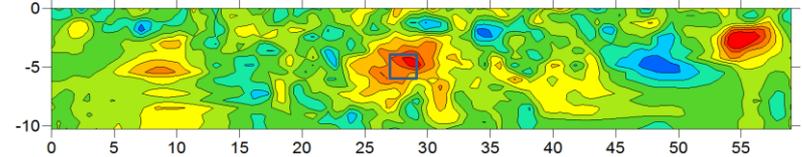
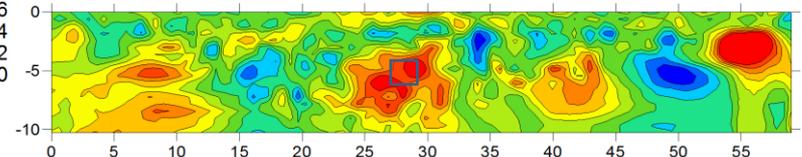
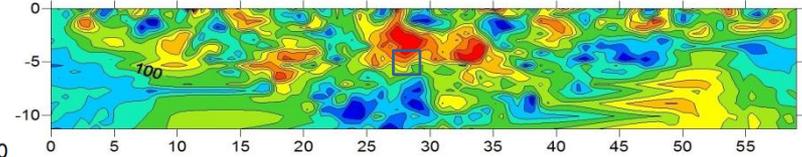
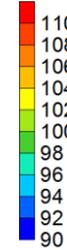
21  $\gamma_{112}$

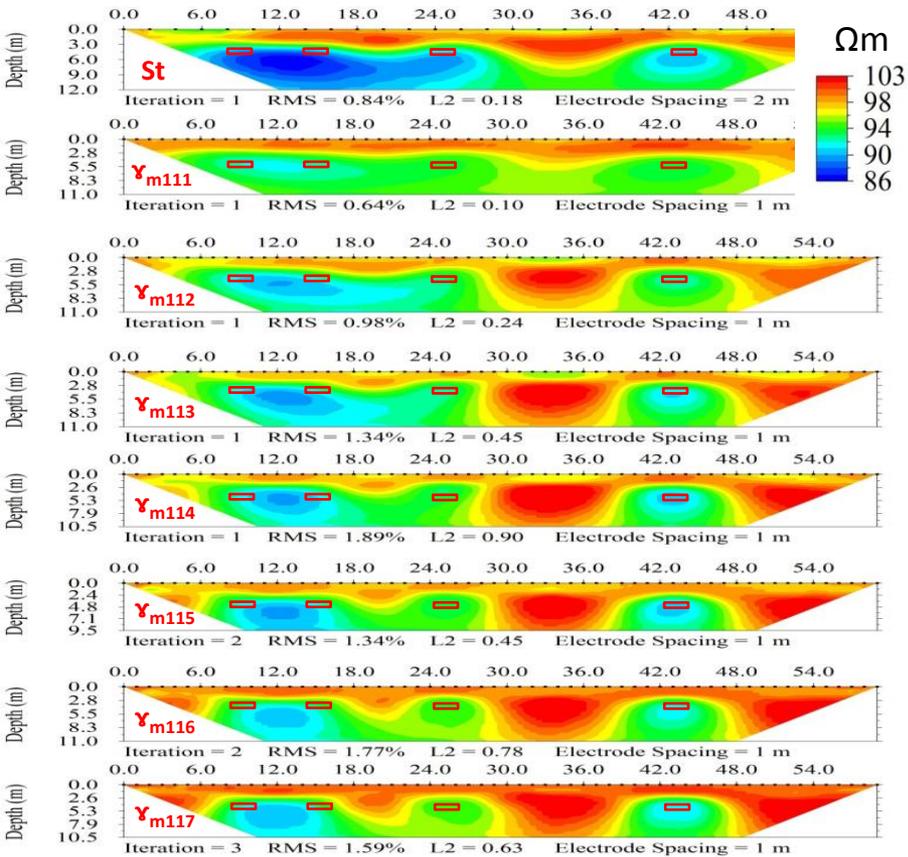
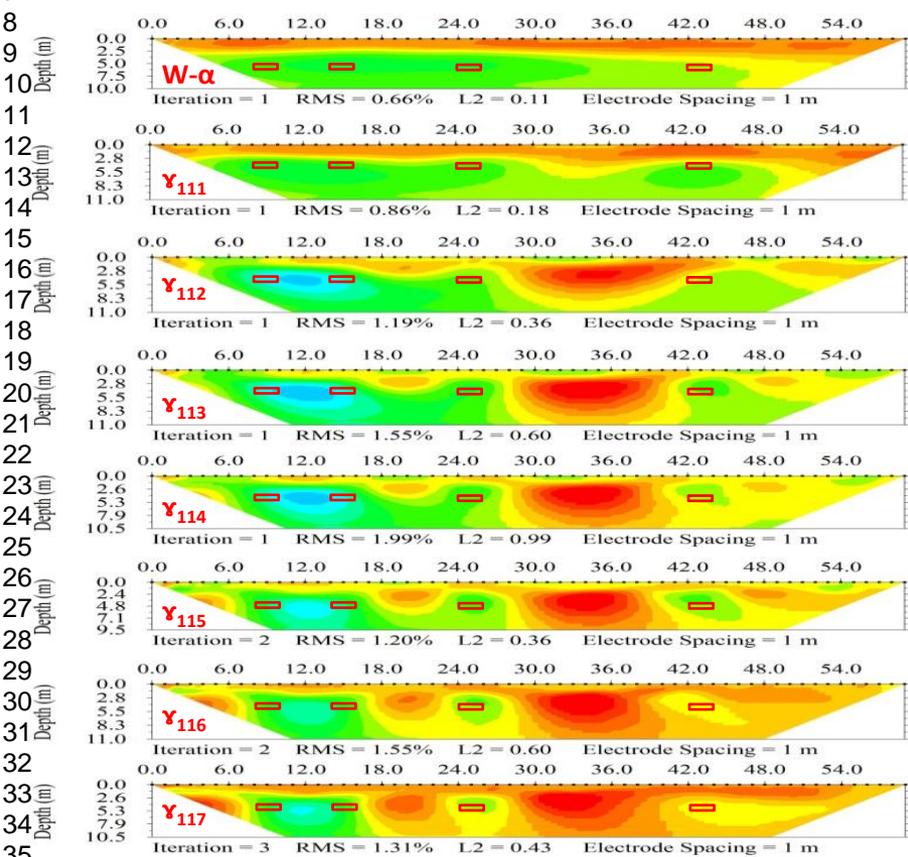
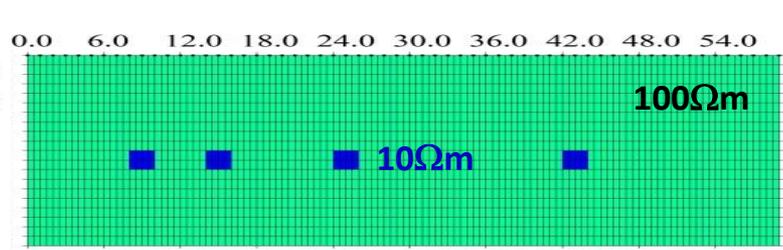
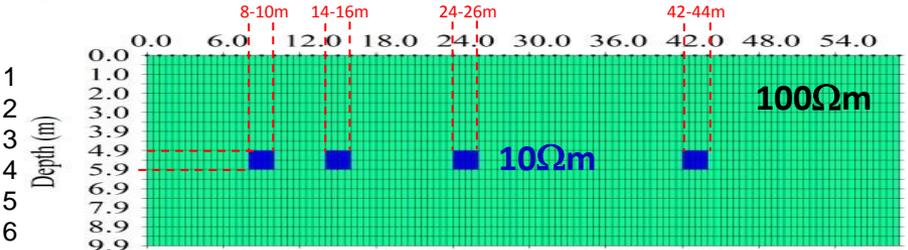
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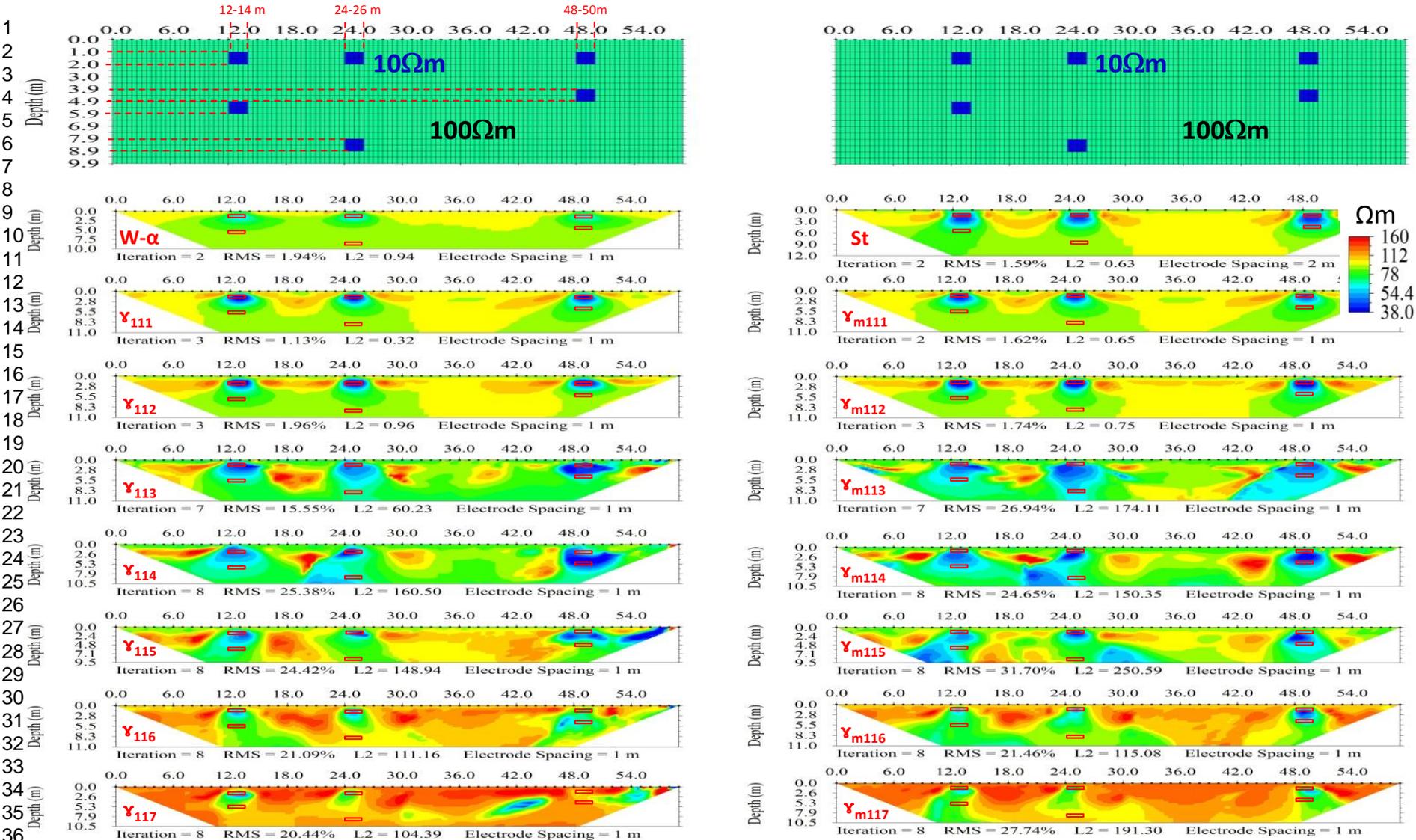
23  $\gamma_{114}$

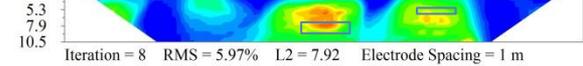
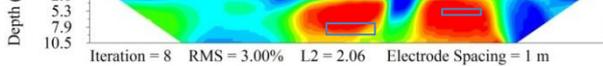
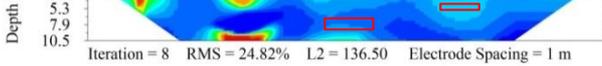
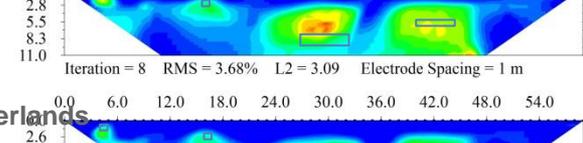
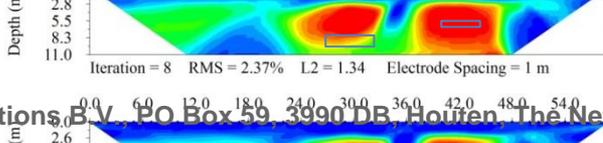
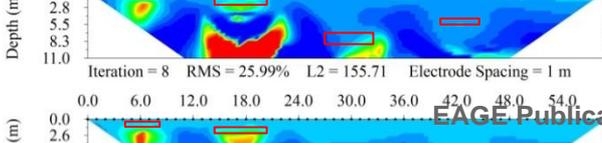
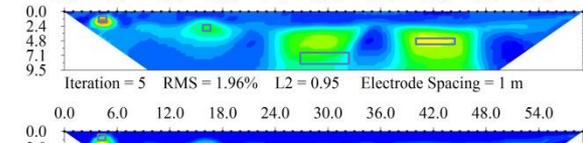
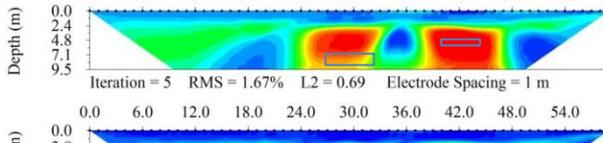
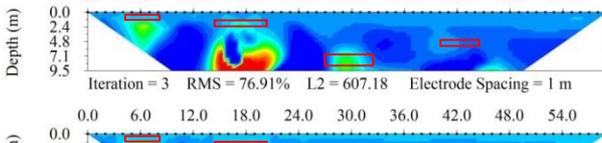
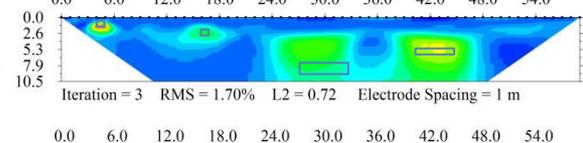
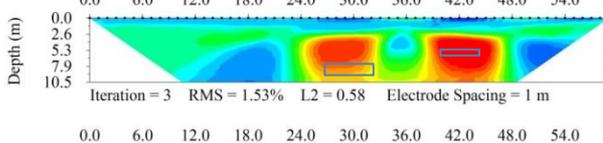
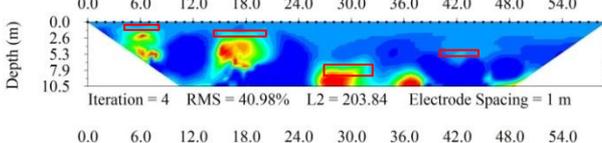
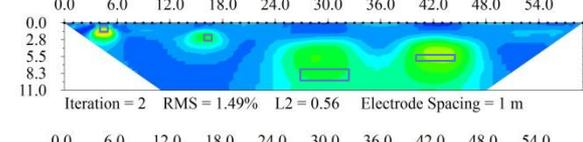
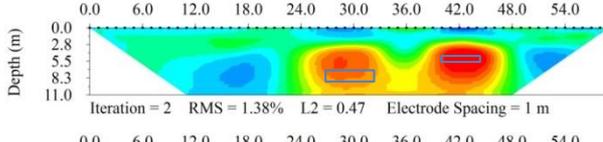
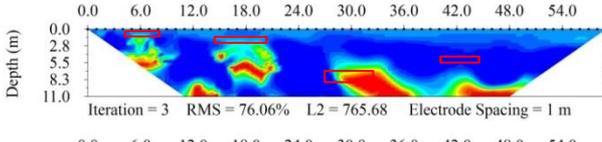
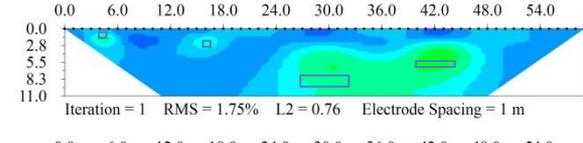
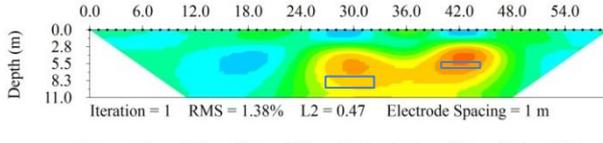
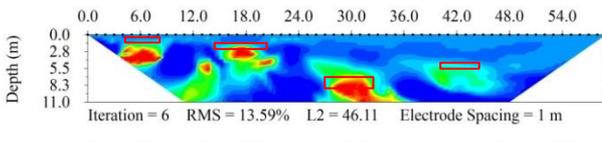
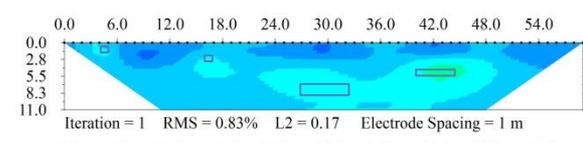
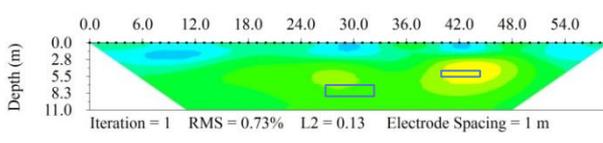
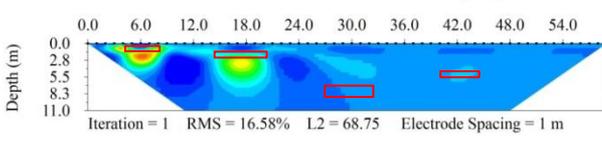
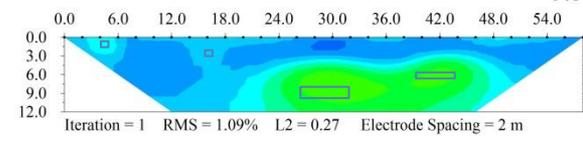
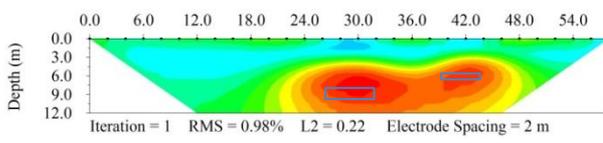
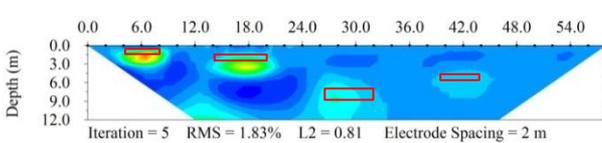
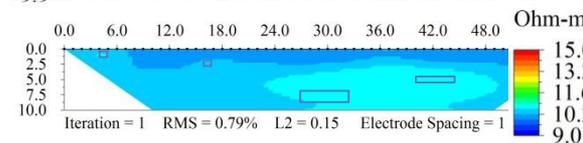
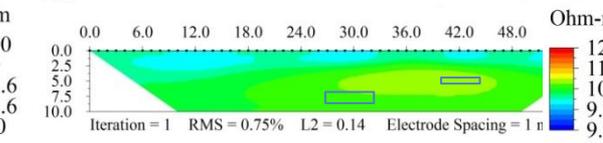
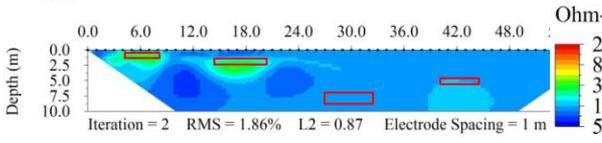
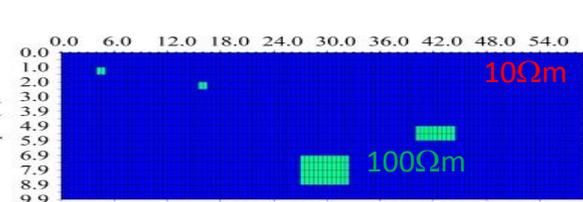
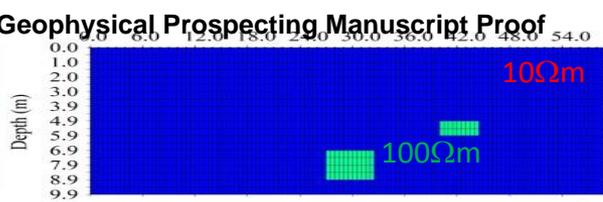
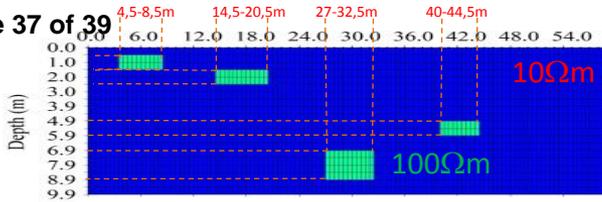


$\Omega\text{m}$

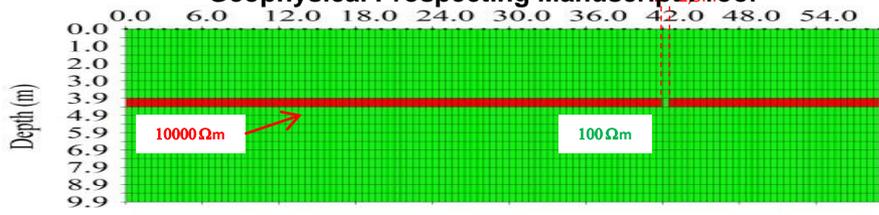




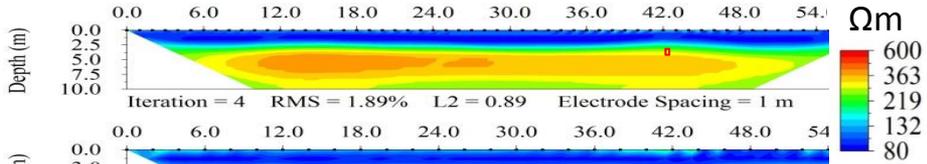




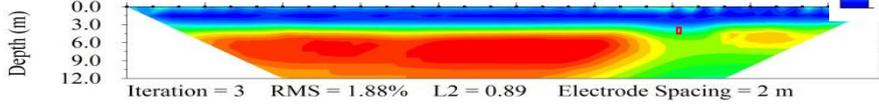
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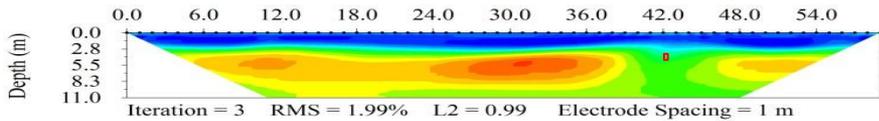
W-α



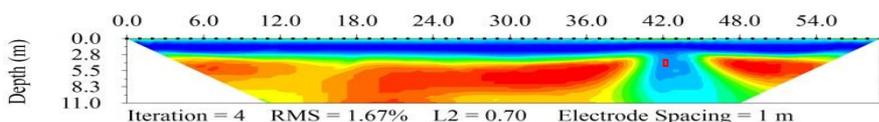
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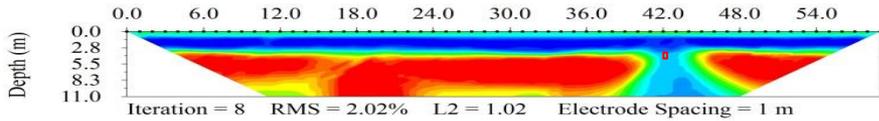
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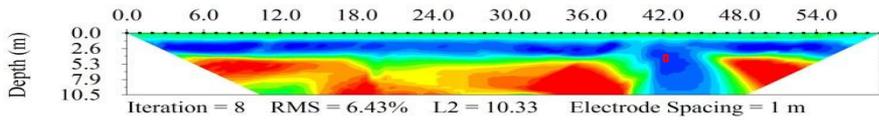
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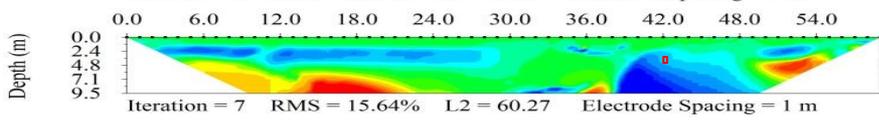
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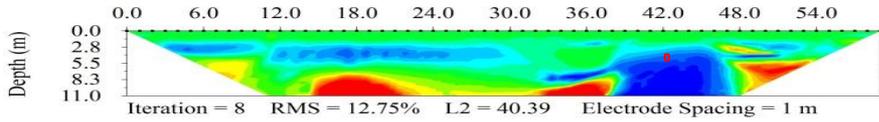
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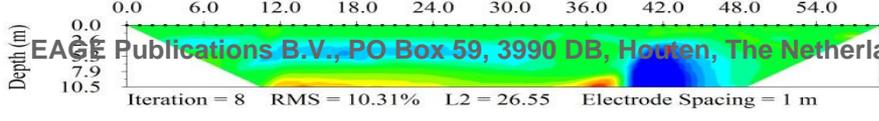
Y<sub>m115</sub>



Y<sub>m116</sub>



Y<sub>m117</sub>



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	Surface	Fig. 6.	Fig. 8-11.
Minimum apparent resistivity ( $\Omega m$ ) =	1	<b>-100000000</b>	<b>-10000</b>
Maximum apparent resistivity ( $\Omega m$ ) =	10000	<b>1000000</b>	10000
Keep All Data :	No	<b>Yes</b>	No
Lower-layer-thickness / Upper-layer-thickness =	1.1	<b>1</b>	<b>1</b>
Depth of Inverted Model / Depth of Pseudosection =	1.1	<b>1</b>	<b>1</b>
Max number of iteration of nonlinear inversion =	8	<b>20</b>	8
Stop RMS error =	3%	<b>5%</b>	<b>2%</b>
Stop when L2 norm is small enough:	No	<b>Yes</b>	No
Initial Lagrange multiplier or roughness factor =	10	<b>1</b>	<b>100</b>
Starting model:	Avg AppRes	<b>Pseudosection</b>	Avg AppRes
Estimated noise of resistivity data =	3%	<b>2%</b>	<b>2%</b>
Initial damping factor of resistivity =	10	<b>1</b>	<b>100</b>