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Amphibole perspective to unravel pre-eruptive processes and conditions in volcanic plumbing systems beneath intermediate arc volcanoes: a case study from Ciomadul volcano (SE Carpathians)

Balázs Kiss · Szabolcs Harangi · Theodoros Ntaflos ·
Paul R. D. Mason · Elemér Pál-Molnár

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10 Abstract Ciomadul is the youngest volcano in the Car-11 pathian-Pannonian region produced crystal-rich high-K 12 dacites that contain abundant amphibole phenocrysts. The 13 amphiboles in the studied dacites are characterized by large 14 variety of zoning patterns, textures, and a wide range of 15 compositions (e.g., 6.4-15 wt% Al₂O₃, 79-821 ppm Sr) 16 often in thin-section scale and even in single crystals. Two 17 amphibole populations were observed in the dacite: low-Al 18 hornblendes represent a cold (<800 °C) silicic crystal 19 mush, whereas the high-Al pargasites crystallized in a hot 20 (>900 °C) mafic magma. Amphibole thermobarometry 21 suggests that the silicic crystal mush was stored in an upper 22 crustal storage ($\sim 8-12$ km). This was also the place where 23 the erupted dacitic magma was formed during the remo-24 bilization of upper crustal silicic crystal mush body by hot

- A1 Communicated by T L. Grove.
- A2 B. Kiss · S. Harangi · E. Pál-Molnár
- A3 MTA-ELTE Volcanology Research Group, Pázmány Péter
- A4 sétány 1/C, Budapest 1117, Hungary
- A5 B. Kiss · S. Harangi
- A6 Department of Petrology and Geochemistry, Eötvös Loránd
- A7 University, Pázmány Péter sétány 1/C, Budapest 1117, Hungary
- A8 B. Kiss (🖂) · E. Pál-Molnár
- A9 Vulcano Research Group, Department of Mineralogy,
- A10 Geochemistry and Petrology, University of Szeged, Egyetem
- A11 utca 2, Szeged 6722, Hungary
- A12 e-mail: geobalazs@gmail.com
- A13 T. Ntaflos
- A14 Department of Lithospheric Research, University of Vienna,
- A15 Althanstrasse 14, 1090 Vienna, Austria
- A16 P. R. D. Mason
- A17 Department of Earth Sciences, Utrecht University, Budapestlaan
- A18 4, 3584 CD Utrecht, The Netherlands

mafic magma indicated by simple-zoned and composite 25 amphiboles. This includes reheating (by ~ 200 °C) and 26 partial remelting of different parts of the crystal mush 27 followed by intensive crystallization of the second mineral 28 population (including pargasites). Breakdown textures of 29 amphiboles imply that they were formed by reheating in 30 case of hornblendes, suggesting that pre-eruptive heating 31 and mixing could take place within days or weeks before 32 the eruption. The decompression rim of pargasites suggests 33 around 12 days of magma ascent in the conduit. Several arc 34 volcanoes produce mixed intermediate magmas with sim-35 ilar bimodal amphibole cargo as the Ciomadul, but in our 36 37 dacite the two amphibole population can be found even in a single crystal (composite amphiboles). Our study indicates 38 that high-Al pargasites form as a second generation in these 39 magmas after the mafic replenishment into a silicic capture 40 zone; thus, they cannot unambiguously indicate a deeper 41 mafic storage zone beneath these volcanoes. The simple-42 zoned and composite amphiboles provide direct evidence 43 that significant compositional variations of amphiboles do 44 not necessarily mean variation in the pressure of crystal-45 lization even if the Al-tschermak substitution can be rec-46 ognized, suggesting that amphibole barometers that 47 consider only amphibole composition may often yield 48 40 unrealistic pressure variation.

KeywordsAmphibole perspective · Intermediate51magmas · Magma mixing · Volcano plumbing system ·52Thermobarometry · Amphibole texture and zoning patterns53

Introduction

Eruptions of intermediate (andesitic to dacitic) arc volcanoes are usually preceded by open-system magmatic 56

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57 processes such as magma mixing, cumulate assimilation, 58 and crustal contamination, producing a petrologically 59 complex mixture of minerals and melts (e.g., Humphreys 60 et al. 2006; Reubi and Blundy 2009; Kent et al. 2010). It is 61 essential to understand these processes because they 62 determine the physical state of the erupted magma and as a 63 consequence the style of the volcanic eruption (Ruprecht 64 and Bachmann 2010; Koleszar et al. 2012).

65 Here, we present the results of a combined textural and 66 chemical (major and trace elements) analyses of amphiboles 67 found in the dacitic rocks of the Ciomadul volcano. The 68 Ciomadul located in the SE Carpathians is the youngest 69 volcano of Carpathian-Pannonian region and referred as a 70 potentially active volcano (Szakács et al. 2002; Harangi 71 2007; Popa et al. 2012; Szakács and Seghedi 2013). Despite 72 this, very little is known about how the erupted magmas were 73 formed or what triggered their eruptions (Vinkler et al. 2007). 74 Using amphiboles, we could imply the volcanic plumbing 75 system and constrain the pre-eruptive magma chamber pro-76 cesses. Amphibole is a powerful tool as they can record 77 parallel pre-eruptive processes and conditions due to its 78 sensitivity for changing magmatic variables such as temper-79 ature, pressure, redox state, H₂O content, melt composition, 80 and co-crystallizing mineral phases (e.g., Johnson and 81 Rutherford 1989a; Rutherford and Hill 1993; Scaillet and 82 Evans 1999; Bachmann and Dungan 2002; Rutherford and 83 Devine 2003; Sato et al. 2005; Humphreys et al. 2006, 2009b; 84 Thornber et al. 2008; Ridolfi et al. 2010; Krawczynski et al. 85 2012). Consequently, the careful investigation of their crystal 86 growth stratigraphy can provide an "amphibolic" perspec-87 tive toward understanding the dynamics and processes of 88 dacitic-andesitic volcanoes before eruptions (Thornber et al. 89 2008) including Ciomadul.

90 We investigate the implications of our study for using 91 amphibole to constrain the subvolcanic plumbing systems 92 of andesitic to dacitic arc volcanoes in general. Finally, we 93 discuss the origin, conditions, and processes lead to com-94 monly observed bimodal amphibole populations in inter-95 mediate mixed magmas erupted at composite arc 96 volcanoes. Our study highlights that different amphibole 97 thermobarometers can produce essentially different results, 98 which may lead to false interpretations on the magma 99 evolution and architecture of the magma storage system 100 without clear textural control and crystal growth stratigraphy. We point out the deficiency of the Ridolfi's thermo-101 102 barometric model that yields always the same p-T 103 evolution path for amphiboles along their stability curve.

Fig. 1 a Geological sketch map of the Carpathian–Pannonian region ► in Eastern-Central Europe. The cross-hatched areas show the Neogene calc-alkaline volcanic rocks (on the surface). b Topographic map of Ciomadul (Csomád) volcano (Karátson et al. 2013). The sample locations are marked by white stars

volcano of this area (Szakács and Seghedi 1995; Szakács 107 et al. 2002; Vinkler et al. 2007; Harangi et al. 2010; Kar-108 átson et al. 2013). It is found at the southern termination of 109 the Călimani-Gurghiu-Harghita (CGH) and esitic-dacitic 110 volcanic chain (Fig. 1) that shows a gradually youngling 111 volcanism from 11.3 Ma (Peltz et al. 1987; Pécskay et al. 112 113 1995). There was a sharp compositional change in the erupted magmas within the Harghita Mountains around 114 3 Ma that spatially coincides with a major tectonic line 115 (Trotus line; Harangi and Lenkey 2007; Seghedi et al. 116 2011). The post-3 Ma magmas at the southern Harghita are 117 more potassic and show different trace element composi-118 tions compared with the older rocks in the north (Seghedi 119 120 et al. 1987; Szakács et al. 1993; Mason et al. 1996). Volcanic eruptions in Ciomadul could have started around 121 200 ka, and the last volcanic event occurred 122 at $31,000 \pm 260$ cal BP. Initially, the volcanism was mostly 123 effusive and a lava dome complex developed called here 124 "old Ciomadul." Later, the volcanic activity became more 125 explosive and as a result of successive phreatomagmatic 126 and subplinian eruptions, the edifice of the "old Ciomadul" 127 was partially destroyed and two deep explosive craters 128 were formed (Szakács and Seghedi 1995; Karátson et al. 129 2013). The erupted magma remained fairly homogeneous 130 through time and shows high-K dacitic composition (Sza-131 kács and Seghedi 1986; Vinkler et al. 1997). The geody-Aq1 32 namic background of the CGH volcanism and particularly 133 the volcanic activity of southern Harghita and Ciomadul is 134 still highly debated. Roll-back subduction and gradual 135 break off of the subducted slab and/or gravitational insta-136 bility and lithospheric delamination was invoked to explain 137 the geodynamics of the area (e.g., Seghedi et al. 2011; 138 Mason et al. 1998; Gîrbacea and Frisch 1998; Chalot-Prat 139 and Gîrbacea 2000; Lorinczi and Houseman 2009; Fillerup 140 et al. 2010; Ren et al. 2012). Whatever is the explanation 141 142 for the origin of the near-vertical slab beneath this area, the active seismicity in the Vrancea zone as well as the gas 143 chemistry and seismicity beneath Ciomadul (Vaselli et al. 144 145 2002; Popa et al. 2012) implies that rejuvenation of the volcanism cannot be unambiguously excluded (Szakács 146 et al. 2002; Harangi 2007; Szakács and Seghedi 2013). 147

Samples and analytical techniques

104 **Geological setting**

105 Ciomadul volcano is located at the southeastern edge of the 106 Carpathian-Pannonian region, and this is the youngest

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149 Samples were collected in Ciomadul during several field campaigns. The fresh lava dome rocks and pumices of the 150 explosive eruptions cover most of eruptive events. During 151



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152 this study, we focused on the lava dome rocks collected at 153 the northern crater rim of the "old Ciomadul." They rep-154 resent effusive products formed about 100-150 ka and 155 contain the entire mineral assemblage as well as felsic and 156 mafic inclusions found in the Ciomadul dacite. Textural 157 characterization of the mineral phases was performed by 158 combined investigation with petrographic microscope and 159 an AMRAY 1830 I/T6 scanning electron microscope at the Department of Petrology and Geochemistry of the Eötvös 160 161 Loránd University. The in situ analyses of the mineral 162 phases were carried out using a CAMECA SX100 electron 163 microprobe equipped with four WDS and one EDS at the 164 University of Vienna, Department of Lithospheric 165 Research (Austria). The operating conditions were as fol-166 lows: 15 kV accelerating voltage, 20 nA beam current, 167 20 s counting time on peak position, and PAP correction 168 procedure for data reduction. Amphibole and plagioclase 169 crystals were measured with defocused beam $(3-5 \ \mu m)$. 170 Calibration was based on the following standards: Amelia 171 albite (Na, Si, Al), San Carlos olivine (Mg) (Jarosewich 172 et al. 1980), almandine 112140 (Fe) (McGuire et al. 1992), microclin (K) (Jarosewich et al. 1980), and the depart-173 174 ment's own standards wollastonite (Ca), rutile (Ti), spes-175 sartine (Mn), synthetic Mg chromite (Cr), and Ni oxide 176 (Ni).

177 Trace elements in amphibole phenocrysts were mea-178 sured by LA-ICP-MS using a 193-nm ArF excimer laser 179 ablation system (MicroLas GeoLas 200Q) in combination 180 with quadrupole ICP-MS (Micromass Platform ICP) at 181 Utrecht University (Mason and Kraan 2002) following the 182 methodology described by Harangi et al. (2005). Ablation 183 was performed at a fixed point on the sample with an irradiance of 0.2 GW cm⁻², a laser pulse repetition rate of 184 10 Hz, and an ablation crater diameter of 40-60 µm. The 185 186 signal recorded by the ICP-MS during ablation was care-187 fully checked for compositional boundaries to ensure that 188 only data for the amphiboles were integrated. Quantitative 189 concentrations were calculated using NIST SRM 612 as a 190 calibration standard (Pearce et al. 1997) with Ca (previ-191 ously determined by electron microprobe analysis) as an 192 internal standard element. The USGS reference glass BCR-193 2 G was continuously measured throughout the analysis of 194 the amphiboles, and the results were within 5-10 % of 195 recommended values. Detection limits were typically in the range 0.01–1 μ g g⁻¹, and internal precision was <5 % 196 RSD (1 σ) for concentrations above 1 µg g⁻¹ and <15 % 197 RSD (1 σ) below 1 µg g⁻¹. 198

199 Petrology of the dacite

200 The studied lava dome rocks are calc-alkaline, high-K 201 dacites (SiO₂ = 62–68 wt%; $K_2O = 3.0-3.6$ wt%)

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Fig. 2 Photomicrographs of amphiboles in the Ciomadul dacite. **a** Hornblende (hbl) and adjacent embayed quartz (q). **b** Hornblende with thick opaque rim and touched pair of K-feldspar (kfp); a small pargasite (parg) microphenocryst with thin opaque rim. c, d Pargasite crystal overgrowth on clinopyroxene (cpx) and olivine (ol) crystals, respectively. e Hornblende with thick opaque reaction rim with biotite inclusion (bt) next to simple-zoned amphibole with thin reaction rim; a plagioclase (pl)-pargasite microphenocryst pair. f Hornblende inclusion in a large plagioclase phenocryst; pargasite microphenocrysts and plagioclase laths in the matrix. g Black opacitized hornblendes in a felsic crystal clot (microdioritic microinclusion) that also contain plagioclase, titanite (tit), zircon (zrn), and intersticial glass; pargasite phenocrysts in the host are also indicated; enlargement (BSE image) of the hornblende in the framed area: close view of the microdioritic inclusion, the hornblende inside is surrounded by clinopyroxene corona set in vesiculated glass gl*. h Simple-zoned amphibole with a "dirty" opacitic core (hornblende), a clear pargasitic rim and a thin opaque reaction rim; apatite (ap) inclusions are present in the core

according to the classification of Gill (1981) and Miyashiro 202 (1974). The erupted dacites are poorly vesicular and por-203 phyritic (phenocrysts content is 30-40 vol%), crystal-rich 204 rocks. They contain abundant glomerocrystic aggregates 205 and crystal clots, and the groundmass is totally crystalline. 206 Thus, the texture of the dacite is glomeroporhyritic holo-207 208 crystalline. The term "phenocryst" is used here for crystals exceeding 300 µm, whereas "microphenocrysts" are in the 209 size range of \sim 300–50 µm. "Microlites" are smaller than 210 50 µm. The phenocrysts are (in the order of relative 211 212 occurrence) plagioclase, amphibole, biotite, clinopyroxene, quartz, K-feldspar, and olivine. The Ciomadul dacite is rich 213 in accessories such as apatite, titanite, and zircon. Allanite 214 is observed occasionally. The groundmass is composed of 215 plagioclase laths, pyroxene microlites, Fe-Ti oxides, and 216 217 SiO₂ patches, probably tridymite.

218 Plagioclase is the most common phenocrystic phase. Two major types were observed: they occur in large 219 glomerocrystic aggregates and as euhedral micro-220 phenocrysts (Fig. 2f). The compositional range of the 221 phenocrysts and microphenocrysts overlaps but the An 222 contents of the phenocrysts' cores are lower (An: 223 $34 \pm 8 \text{ mol}\%$) than those of the cores of microphenocrysts 224 225 (An: $52 \pm 4 \mod \%$). Amphibole phenocrysts are the most 226 common mafic minerals. Detailed descriptions of amphiboles are presented in the following sections. Biotite 227 phenocrysts are euhedral or subhedral, often rounded and 228 always have a reaction rim. Common inclusions in biotite 229 are apatite, zircon, and glass. Quartz phenocrysts (<5 vol% 230 of the phenocrysts) are always rounded and embayed 231 (Fig. 2a), and occasionally, there are a clinopyroxene-rich 232 rim around them. They are found mostly in the samples 233 with higher proportions of mafic crystal clots, clinopyrox-234 ene, and olivine phenocrysts. K-feldspar phenocrysts 235 236 (<5 vol% of the phenocrysts) are also embayed and rounded (Fig. 2b). Some of them are large ($\sim 5 \text{ mm sized}$) 237

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238 crystals enclosing amphibole, biotite, apatite, titanite, pla-239 gioclase, and quartz. Rare olivine crystals are anhedral and 240 usually strongly reacted (Fig. 2d). They are surrounded by 241 fine-grained reaction products (pyroxenes and oxides) and 242 overgrown by amphibole crystals. Tiny euhedral and 243 homogenous Cr-spinel inclusions often occur in the oli-244 vines. Clinopyroxenes usually form crystal clots but they 245 also appears as single crystals. They have euhedral or subhedral habit and show various internal zoning. It is 246 247 often observed that clinopyroxene crystals have rounded 248 edges and they are overgrown by amphibole crystals 249 (Fig. 2c).

250 The dacites contain various crystal clots. Two larger 251 groups were recognized according to their mineralogy: 252 felsic and mafic clots. Felsic clots are \sim 1-cm-sized inclu-253 sions consisting dominantly of felsic minerals such as pla-254 gioclase \pm K-feldspar \pm quartz. The most common type of the felsic clots is microdioritic (Fig. 2g) composed of pla-255 256 gioclase, amphibole, biotite, titanite, zircon, and apatite. 257 They are texturally similar to adcumulate plutonic rocks but 10-15 vol% interstitial vesiculated glass is always present. These felsic inclusions could represent a highly crystalline magmatic body in the magma chamber (i.e., crystal mush). Mafic clots contain clinopyroxene or olivine or both with 262 minor amount of plagioclase laths. They are occasionally 263 overgrown by amphibole crystals. (Fig. 2c, d).

264 Amphibole texture and chemistry

265 All of the studied rocks contain two major amphibole types according to their general optical appearances. Amphiboles 266 267 that are "dirty" and have dark brown-reddish brown pleochroism are termed here "hornblendes" (Fig. 2e). They 268 usually occur as euhedral or subhedral single phenocrysts 269 270 with no optical zoning and often coexist with K-feldspar, 271 glomerocrystic plagioclase, and quartz (Fig. 2a, b, f) where 272 hornblende is found as intergrowth or inclusions. Amphi-273 boles in the microdioritic inclusions are also hornblendes. 274 These amphiboles contain biotite, apatite, rhyolitic glass, 275 and rarely titanite and plagioclase inclusions (Fig. 2e, h). 276 Hornblendes show various breakdown textures. The second 277 type of amphiboles is "clear" with light brown-yellow 278 pleochroism and will be called "pargasite." They occur as 279 single phenocrysts with optical zoning, but are often found 280 also as overgrowth rim on hornblende, clinopyroxene, and 281 olivine crystals (Fig. 2b-d, h). Amphiboles in the mafic 282 clots are also pargasites. They contain small clinopyroxene, 283 apatite, and sulfide inclusions (Fig. 2d). Pargasites also 284 show breakdown texture, which is described in the fol-285 lowing section in more detail. Amphiboles often coexist 286 with plagioclase. The hornblendes typically occur with 287 glomerocrystic plagioclase and pargasites with euhedral 288 plagioclase microphenocrysts (Fig. 2e, f).

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Amphibole breakdown textures

Amphibole phenocrysts show different breakdown tex-290 tures: reaction rims, internal breakdown patches, and 291 clinopyroxene-rich zones or coronas (Fig. 2, 3). Optically, 292 the reaction rims and internal breakdown patches are op-293 acitic. These various textures can be observed in single thin 294 295 sections and even in single crystals (Fig. 2e, h). The two amphibole groups show marked differences in their 296 breakdown textures. Hornblendes are characterized by a 297 higher degree of breakdown. They are surrounded by a 298 299 thick, coarse-grained rim consisting of pyroxene, Fe-Ti oxide, feldspar, and glass. Additionally, similar breakdown 300 products also appear within the crystals as patches or along 301 cleavages (internal breakdown patches; Fig. 3). The reac-302 tion rim is often present also where the hornblende is in 303 solid-solid contact with other phenocryst (Fig. 3d). The 304 average rim thickness is $\sim 15 \,\mu\text{m}$ (Fig. 3e) but in some 305 samples, the hornblendes are almost totally replaced by the 306 reaction products (Fig. 2g). Clinopyroxene-rich corona was 307 also observed around some hornblendes, partly in crystal 308 clots (Fig. 2g). Around the pargasite crystals, the reaction 309 rim is finer-grained compared to the breakdown products of 310 hornblendes (Fig. 3c). These rims consist of pyroxenes, 311 Fe-Ti oxides, and some feldspar. The reaction rim can be 312 seen only where the pargasite is in direct contact to the 313 matrix. 314

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315 Some amphiboles show multiple breakdown textures. 316 The core of these crystals is characterized by internal breakdown patches similar as it was observed in the 317 hornblendes, but their outer reaction rim is thin and fine-318 grained as it was observed around pargasites (Fig. 3a, b). 319 Additionally, thin clinopyroxene-rich zone is occasionally 320 321 observed at the core-rim boundary.

Amphibole major and trace element chemistry, 322 classification 323

Amphiboles show large intercrystalline compositional 324 variation in major (e.g., $Al_2O_3 = 6.4-15$ wt%, MgO = 325 9.3–17.6 wt%) and trace elements (e.g., Ba = 20-500 ppm, 326 Sr = 100-800 ppm) (Table 1, 2, 3). In spite of the diversity AQ2 S27 of their textural appearance and variable zoning patterns, the 328 amphiboles can be divided into two separate compositional 329 groups, which correspond to the petrographically determined 330 groups. Hornblendes show low-Al and Ba, Sr, Zr contents, 331 and they are characterized by high SiO₂ and MnO, and low 332 TiO_2 and Na_2O contents (Fig. 4, 5) compared with the 333 pargasites. Their low Al/Si (0.19 \pm 0.02) ratio, low Ba, Sr, Zr 334 concentrations, and the negative Eu anomaly in the normal-335 ized trace element patterns (Fig. 5) suggest that they origi-336 nated from an evolved, fractionated silicic magma. Pargasite 337 338 crystals are rich in Al as well as Ba, Sr, Zr, and they have high



Fig. 3 Pictures of amphibole breakdown textures and histograms of reaction rim thicknesses. **a** Photomicrograph of an amphibole phenocryst with dirty opacitic core and clear rim (plane-polarized light). **b** Close-up (BSE image) of the area shown by the *rectangle* in picture **a** showing that the internal breakdown patches of the core

contain fine-grained reaction products. c Close view of the finegrained thin reaction rim of pargasite and the coarser-grained thick reaction rim of hornblende. d Hornblende inclusion in plagioclase: the hornblende is surrounded by a reaction rim. e Histogram of the reaction rim thicknesses of the studied amphiboles

 Table 1 Representative major element compositions of Ciomadul amphibole crystals

sample	NCS2-30a	Tc-30a	kcs17-10	00	1	kcs17-10)		kcs17-30	Mo2-30b
	Hornblende:					Pargasite	:			
crystal	NCS2 am7 felsic clot	Tc30a_am4 felsic clot	am_2 phcr	am_19_1 phcr	am_19_2 phcr	am_15 phcr	am_17 phcr	am_13 mafic clot	38_am4 mafic clot	mo2 am4 mphcr
SiO2	46.02	44.98	46.75	45.88	46.43	43.07	41.33	42.65	43.08	43.23
TiO2	1.02	1.32	0.97	1.21	1.13	2.18	2.14	2.18	2.35	2.35
Al2O3	7.66	8.57	7.35	8.46	7.70	11.92	13.73	13.30	11.91	11.2
FeO	14.54	15.33	14.57	14.38	14.05	9.15	9.68	7.30	7.67	10.2
MnO	0.39	0.42	0.47	0.39	0.38	0.12	0.09	0.11	0.09	0.14
MgO	12.98	12.19	13.50	13.20	13.40	15.92	14.46	16.56	16.55	15.11
CaO	11.79	11.86	11.81	11.74	11.88	11.69	11.60	11.80	11.96	11.57
Na2O	1.46	1.48	1.27	1.53	1.32	2.30	2.40	2.44	2.48	2.34
K2O	0.80	0.94	0.68	0.84	0.77	0.93	0.98	1.16	0.97	0.75
BaO	n.a.	n.a.	0.07	0.07	0.09	0.11	0.09	0.09	n.a.	n.a.
Total	96.67	97.09	97.43	97.69	97.14	97.38	96.50	97.57	97.06	96.94

FeO, total amount of iron

phcr phenocryst, mphcr microphenocryst, n.a. not analyzed

TiO₂ and Na₂O, and low SiO₂ and MnO contents (Fig. 4, 5). The lack of negative Eu anomaly, along with the high Al/Si ratio (0.33 ± 0.03), suggests that these amphiboles crystallized from a more primitive, presumably mafic magma. The CaO content is similar in both groups and show minor variations. The MgO concentration of the two groups overlaps but343ations. The MgO concentration of the two groups overlaps but344the low-Al amphiboles cluster at lower MgO than the high-Al345amphiboles. The compositional variation can be also large346

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shaped bright and dark patches and oscillatory-zoned rims.

In the dark patches, melt inclusion with euhedral apatite and

mineral inclusions of biotite and apatite are present (Fig. 6).

Where it is appreciable, the bright patch looks as a spongy

framework with strongly dissolved margin. Brighter por-

tions are richer in Al, Ti, and $(Na + K)^A$ and have lower Mg

number than the dark regions (Fig. 8). Oscillatory-zoned

rims are characterized by Al spikes along with increase in Fe,

Ti, and $(Na + K)^A$, and a decrease in Mg and Si that follows

the dissolution surface (Fig. 7). The other type (type A2) of

patchy-zoned amphiboles has rounded patchy-spongy cel-

lular core and dark rim (Fig. 6) and can be observed less

frequently. The patches have sharper boundary, and an

abrupt shift was observed in their composition. Brighter

patches have lower Al, Ti, $(Na + K)^A$, and Mg contents

compared to the dark ones (Fig. 8). The dark patches have

similar compositions as the rim that contains pyroxene

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Table 2 Representative trace element compositions of Ciomadul amphibole crystals

sample	kcs17-10	0			MO2	kcs17-10	0		Mo2	Ncs2
	Hornblen	nde:				Pargasite	2:			
crystal	am_2 phcr	am_19_1 phcr	am_19_2 phcr	am_6 S-Z core	MO2-7 phcr	am_15 phcr	am_17 phcr	am_13 mafic clot	MO2-1 phcr	NCS2-4 phcr
v	332.17	323.88	296.97	319.30	315.27	352.49	421.31	479.17	426.33	444.71
Cr	84.56	179.22	114.45	135.51	114.98	79.13	18.61	405.56	33.27	90.2
Ni	50.84	45.31	37.65	39.04	56.18	47.23	53.61	56.48	28.91	22.76
Rb	5.53	2.66	2.45	15.21	3.49	3.83	8.08	5.53	4.58	3.94
Sr	73.92	132.21	109.02	168.27	84.44	749.01	822.30	585.64	652.25	745.93
Y	45.84	31.06	27.87	33.50	36.22	20.85	22.76	19.04	25.64	20.53
Zr	28.93	35.53	31.27	34.14	25.75	60.20	68.39	58.71	77.56	56.27
Nb	18.29	18.51	16.27	16.17	19.09	15.74	10.53	19.25	14.29	14.57
Ba	74.77	119.66	104.02	115.83	65.24	347.92	401.52	294.31	331.53	366.85
La	15.53	23.51	16.70	18.29	12.98	14.15	15.21	11.81	15.71	11.38
Ce	54.25	77.01	61.27	60.10	48.76	46.91	50.42	38.50	47.35	37.01
Pr	9.04	12.34	9.15	9.47	8.18	8.08	8.51	6.59	7.75	6.38
Nd	40.52	48.18	39.35	41.16	36.65	35.74	39.78	28.08	37.64	30.53
Sm	10.74	9.89	7.66	9.31	8.35	9.04	10.00	7.71	8.84	7.45
Eu	2.55	2.45	2.02	2.34	1.85	3.08	3.30	2.23	2.73	2.55
Gd	9.79	9.57	8.40	8.83	7.42	7.34	7.76	6.07	7.59	6.6
Dy	10.90	8.35	6.86	8.62	8.18	6.49	7.55	5.90	6.82	6.12
Er	6.38	4.41	3.88	5.21	4.53	3.08	3.83	3.19	3.44	3.19
Yb	5.37	3.88	3.51	4.41	4.42	2.71	2.98	2.50	2.78	2.13
Lu	0.74	0.53	0.43	0.64	0.65	0.32	0.43	0.32	0.44	0.32
Hf	2.45	2.13	2.02	2.34	2.29	3.19	3.19	2.87	3.27	2.87
Та	0.74	0.43	0.53	0.53	0.55	0.96	0.64	1.17	0.98	0.74
Pb	3.08	1.70	2.23	2.02	2.07	2.45	4.68	1.49	1.42	1.28
Th	0.53	0.19	0.19	0.43	0.1	0.43	0.53	0.24	0.44	0.32
U	0.14	0.04	0.05	0.12	0.04	0.09	0.22	0.04	0.15	0.05

Phcr phenocryst, S-Z core core of simple-zoned amphibole

within a single crystal; it can be equal to the whole variation ofthe phenocrysts (Fig. 4).

According to the classification of IMA (Leake et al. 1997), both amphibole groups belong to the Ca amphiboles. The low-Al amphiboles are dominantly Mg hornblendes and some of them are edenites. The high-Al amphiboles are dominantly pargasites but Mg hastingsites also occur.

355 Zoning types and amphibole profiles

356 Patchy zoning

Two types of patchy zoning were distinguished in the dacite
(based on BSE images). Type A1 patchy zoning was
observed in the backscattered electron images of hornblende
phenocrysts. These crystals consist of a core of irregular-

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Zoning-type	Simple-z (Type C)	oned	Patchy-zone	d (Type A1)	Patchy- A2)	zoned (T	ype	Oscillatory-zo B1)	ned(Type	Oscillate	ory-zoned(Ty	pe B2)	Composit	te crystal	(Type D)				
crystal	core	rim	dark patch	light patch	core		rim	bright band	dark band	core	bright zone	nim	core		oscillator	y rim			
	kcs17 an	ıf sz3	am5		kcs_17_	_amf_1_4		am8		am12			kcs17 am	f x					
SiO2	47.19	43.91	46.44	44.07	45.86	43.14	43.56	44.10	48.67	44.03	38.80	42.91	48.02	44.76	43.05	38.89	44.64	41.08	43.75
TiO2	0.83	1.99	1.20	1.43	0.83	2.07	2.17	1.50	0.85	2.05	2.89	2.30	0.84	1.32	2.18	2.06	2.18	2.37	2.20
A12O3	7.19	11.74	7.84	9.41	7.78	11.52	11.28	9.14	6.09	11.76	14.43	12.18	6.46	8.72	12.17	14.48	10.87	13.33	11.86
FeO	14.27	8.76	14.42	15.48	14.58	8.23	6.86	15.42	13.29	7.14	14.19	8.58	13.63	15.31	9.70	16.87	7.01	12.71	8.67
MnO	0.46	0.12	0.43	0.38	0.43	0.09	0.09	0.41	0.44	0.11	0.17	0.13	0.40	0.41	0.12	0.20	0.05	0.15	0.15
MgO	13.39	15.82	13.18	11.92	13.44	16.75	17.56	11.90	14.27	17.48	11.19	15.62	14.20	12.51	15.02	9.31	17.26	12.53	15.62
CaO	11.91	11.88	11.68	11.74	11.75	11.87	11.81	11.67	11.95	11.23	11.73	12.17	11.72	11.82	11.80	11.69	11.91	11.83	11.94
Na2O	1.23	2.12	1.26	1.58	1.52	2.35	2.24	1.49	1.05	2.48	2.40	2.28	1.51	1.70	2.29	2.20	2.16	2.22	2.26
K20	0.74	0.88	0.74	0.97	0.84	0.97	0.97	1.00	0.51	0.96	1.10	1.03	0.64	0.93	0.94	1.10	1.00	1.04	0.98
BaO	n.a.	n.a.	0.02	0.07	0.75	0.25	0.41	n.a.	n.a.	n.a.	n.a.	n.a.	n.a.	n.a.	n.a.	n.a.	n.a.	n.a.	n.a.
Total	97.22	97.23	97.25	97.04	77.79	97.28	96.96	96.63	97.17	97.23	96.90	97.26	97.436	97.488	97.251	96.786	97.064	97.265	97.443
FeO, total am	nount of ire	u																	

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inclusion. The bright patches fall into the compositional field 378 379 of hornblendes, while the compositions of the dark patches 380 and the rim belong to the pargasite field.

Cyclic zoning

Cyclic zoning (type B) is a commonly observed zoning 382 pattern in Ciomadul amphiboles, and it occurs in two 383 subtypes, i.e., type B1 and type B2 (Fig. 6). The subtypes 384 can be distinguished by their Altot content and the com-385 positional pattern of the growth zones. Type B1 zoning is 386 typical in the hornblende crystals, and these crystals show 387 compositional variation at the low-Al level. Type B1 cyclic 388 zoning is characterized by internal asymmetric growth 389 zones and Al spikes at their rim. Each growth zone con-390 391 tains a dark Mg- and Si-rich band and a brighter Al-, Fe-, Ti-, $(Na + K)^{A}$ -rich band. The growth zones are bounded 392 by resorption surfaces (Fig. 7, line 2). Each zone begins 393 with Al-rich and Mg-poor composition but Al gradually 394 decreases and Mg increases toward the end of the zone. Al 395 spikes are thin, symmetrical Al-, Fe-, Ti-, and $(Na + K)^{A}$ -396 rich zones. 397

Type B2 zoning was observed in pargasite crystals. This 398 399 zoning type is present in single phenocrysts or in overgrowth rim on hornblende and olivine crystals. Composi-400 tion of Type B2 crystals varies at the high-Al range. This 401 402 zoning type is built up by asymmetric growth zones that have resorbed boundaries (Fig. 6). In these crystals, the 403 bright zones show often rounded edges suggesting disso-404 lution. The bright zones enriched in Al, Fe, and $(Na + K)^A$, 405 whereas the dark zones are Mg- and Si-rich (Fig. 7, line 3, 406 4). Each growth zone begins with Mg-rich and Al-poor 407 composition, and they show gradual change (increasing Al, 408 decreasing Mg) toward the end of the growth zone. 409

Simple zoning

Simple-zoned (Type C) amphiboles are frequent in the 411 412 dacite and show revers geochemical variation toward the 413 crystal rims (Fig. 6). The most important features are the presence of major resorption surfaces with an abrupt shift 414 in Al, Ti, $(Na + K)^A$, and Mg at the core-rim boundary 415 (example shown in Fig. 7, line 5). The bright cores are 416 typically rounded and often spongy cellular, and the cavi-417 418 ties are filled by fine-grained reaction products of pyrox-419 ene, magnetite, and glass. In extreme cases, the core is totally replaced by these reaction products. The resorbed 420 cores are overgrown by dark amphibole rims enriched in 421 Al, Ti, $(Na + K)^A$, and Mg. The composition of the cores 422 overlaps the hornblende field, while the rims have com-423 position akin to pargasites (Fig. 4). 424

Composite crystals (Type D) are special types of simple-425 zoned amphiboles. They (Fig. 6) consist of the same 426

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not analyzed

n.a.

Table 3 Representative major element compositions of zoned amphibole crystals



Fig. 4 Major element variation diagrams for the studied amphiboles of the Ciomadul dacite: Al_2O_3 versus SiO_2 (a), MgO (b), TiO_2 (c), CaO (d), FeO (e), Na_2O (f), MnO (g), and K_2O (h) contents. All oxides are in wt%

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Fig. 5 Chondrite-normalized (Nakamura 1974) REE variation diagrams (\mathbf{a}, \mathbf{b}) and primitive mantle-normalized (Sun and McDonough 1989) trace element variation diagrams (\mathbf{c}, \mathbf{d}) of the studied amphiboles from the Ciomadul dacite

427 compositions that are seen in the two main amphibole 428 populations with the full range of compositional variation 429 seen even in a single crystal. These crystals are 2-3 mm 430 sized, usually much larger than the other phenocrysts. The 431 crystals can be divided into three parts: a rounded, reacted 432 core, which is surrounded by a thin zone of clinopyroxene 433 and an oscillatory-zoned amphibole rim (Fig. 7, line 6). 434 The crystal core contains mineral inclusion (biotite, apatite, titanite) and compositionally is indistinguishable from 435 436 hornblende phenocrysts. The rim shows the same type of 437 oscillatory zoning as Type B2 pargasites. The composi-438 tional variation in these crystals is particularly remarkable, 439 since it overlaps the whole chemical variability of the 440 amphibole phenocrysts (Fig. 4). Despite the relatively rare 441 occurrence of the composite crystals, they have a great 442 significance concerning the condition of the magma reservoir. 443

444 Al^{IV}–Mg# systematics and substitutions in the zoned 445 amphibole phenocrysts

Three intracrystalline compositional trends can be distinguished in the Al^{IV}–Mg# diagram (Fig. 8a). Trend 1 is a continuous low-slope negative trend that is observed in the

449 hornblende group including type A1 patchy and type B1 oscillatory-zoned crystals. Trend 2 is also a continuous 450 negative trend, but its slope is much steeper and it is typical 451 for the pargasite group (type B2 and rim of type D). The 452 composition of these crystals is controlled by the Ti-453 tschermak $[Si^{IV} + Mg^{VI} = Al^{IV} + Ti^{VI}]$ and edenite sub-454 stitution $[Si^{IV} + (\Box)^{A} = Al^{IV} + (Na + K)^{A}]$. Some role 455 of the Al-tschermak substitution $[Si^{IV} + Mg^{VI} = -$ 456 $AI^{IV} + AI^{VI}$] can be inferred in the Type B2 crystals. The 457 third intracrystalline compositional variation of amphiboles 458 is the covariation of Al^{IV} and Mg# characterizing core-to-459 rim variation of Type C and Type A2 crystals. It is rep-460 resented by the positive, non-continuous trend 3 that 461 associates with increasing Ti, Al^{VI} , and $(Na + K)^A$ from 462 core to rim suggests the role of edenite and tschermak 463 substitutions. 464

Amphibole thermobarometry

465

Empirical and experimental studies indicate that amphibole466composition can be effectively used to quantify the pre-
eruptive p–T conditions during crystallization (e.g., Ham-
marstrom and Zen 1986; Johnson and Rutherford 1989a;469



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470 Blundy and Holland 1990: Anderson and Smith 1995: 471 Ernst and Liu 1998; Bachmann and Dungan 2002; Ridolfi 472 et al. 2010). For the estimation of the pre-eruptive condi-473 tions, the following strategy was applied: first, we com-474 pared the composition of the Ciomadul amphiboles with 475 experimental results that was followed by the application 476 of different thermobarometric techniques. The experimen-477 tal data set was filtered as suggested by Ridolfi et al. (2010) 478 to avoid inconsistent amphiboles. This comparison could 479 be used as a first approximation of the pre-eruptive 480 parameters.

481 The composition of the Ciomadul amphiboles and the 482 coexisting minerals is such that several different thermo-483 barometric methods can be applied. The Ridolfi et al. 484 (2010) amphibole thermobarometer called here R2010, and 485 its later extension (Ridolfi and Renzulli 2012, called here 486 as R2012) is calibrated for a wide range of conditions and 487 can thus be applied for both amphibole groups in the 488 Ciomadul dacite. The other advantage of this thermoba-489 rometer is that it does not require other minerals to be in 490 equilibrium, and therefore, the intensive parameters can be 491 calculated even along zoning profiles in single amphiboles. 492 Petrographic observations and trace element data indicate 493 that low-Al hornblendes could crystallized along with the 494 mineral assemblage that is required for the amphibole 495 thermobarometric calculation developed by Anderson et al. 496 (2008) (called here as A2008) (RiM69_Ch04_hbld_-497 plag_thermo-jla.xls spreadsheet). This thermobarometry comprises the Al^{tot} in amphibole barometry (P_{Al-in-am}) 498 499 (Schmidt 1992; Anderson and Smith 1995) and the 500 amphibole-plagioclase thermometry (Blundy and Holland 501 1990; Holland and Blundy 1994). Hornblendes often 502 coexist with glomerocrystic plagioclases (Fig. 2f), and they 503 can be considered as cocrystallized phases. On the other 504 hand, pargasites were observed with plagioclase micro-505 phenocrysts as touched pairs (Fig. 2e) and were used in the 506 Anderson's calculation in spite of the lack of the necessary 507 mineral phases. The hygro-barometer of Krawczynski et al. 508 (2012) (K2012) was used to estimate the crystallization 509 pressure of pargasites. The results of thermobarometric 510 calculations are summarized in Fig. 9.

511 Pre-eruptive temperature estimation

512 Comparison of the Ciomadul amphiboles with experimental 513 amphibole compositions indicates that the two amphibole 514 groups could crystallize over two distinct temperature ranges 515 (Fig. 10), and this is corroborated by the thermometric cal-516 culations. Low-Al hornblendes coexist and probably co-517 crystallized with a mineral assemblage of biotite, K-feldspar, 518 quartz, plagioclase, titanite, apatite, zircon, allanite, and 519 rhyolithic melt. Experimental studies on similar dacitic rock 520 composition and mineral assemblage (e.g., Fish Canyon tuff

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Fig. 6 BSE images showing the zoning patterns and textures of ► amphiboles in the Ciomadul dacite. Smallest arrows show dissolution surfaces. Compositional profiles of amphiboles are indicated by large thick arrows in Fig. 7. a Type A1 patchy-zoned amphibole; the points indicate the places of chemical measurements from the patchy core. **b** Type A2 patchy zoning with spongy core in amphibole. Irregular dotted lines indicate the outlines of the patchy cores. c Type B1 cycliczoned amphiboles (hornblende). d Type B2 cyclic zoning in amphiboles (pargasite). e Type B2 cyclic-zoned amphibole (pargasite). f Type C simple zoning in amphibole. g Type D composite amphibole crystal with internal breakdown patches in its core. h Enlargement of the area indicated by the rectangle in picture g displaying the dissolution surfaces as well as the clinopyroxene zone. Bt biotite, ap apatite, smi silicate melt inclusion, cpx clinopyroxene

dacite, Johnson and Rutherford 1989b) reproduced amphi-521 boles with similar composition at near-solidus temperatures 522 523 (<800 °C). The amphibole–plagioclase thermometry yields $T_{\text{am-plag}} = 732 \pm 27 \text{ °C}$ for mineral pairs, which are in 524 525 contact with each other and $T_{\text{am-plag}} = 730 \pm 15$ °C for the hornblende composition combined with mean plagioclase 526 glomerocrystal core composition. For the same amphibole 527 crystals, the R2010 thermometer gives an average tempera-528 ture $T_{R2010} = 821 \pm 20$ °C, which is almost 100 °C higher 529 then the previous values. The R2012 calibration provides a 530 little bit lower temperature ($T_{\rm R2012} = 776 \pm 21$ °C) than the 531 R2010, and this is close to the $T_{\text{amp-plag}}$ within error. 532

The high-Al pargasites fall into the experimental com-533 position field of amphiboles that was produced at much 534 higher temperatures, i.e., >900 °C. Thermometric calcu-535 lations give high crystallization temperature values 536 $(T_{\rm R2010} = 972 \pm 20 \text{ °C}, T_{\rm R2012} = 944 \pm 20)$. The amphi-537 bole-plagioclase thermometry indicates lower temperature Tam-plag = 837 ± 20 °C for the analyzed pargasite-pla-539 gioclase pairs and Tam-plag = 840 ± 11 °C for the 540 541 pargasite composition combined with mean the core composition of plagioclase microphenocrysts.

Estimated depth of the magma storage

544 Pressure of the amphibole crystallization, i.e., the depth of 545 the magma storage, was also calculated using the same strategy outlined in the previous section. Comparison of 546 our data and the experimental data set suggests amphibole 547 crystallization in the mid- to lower-crust beneath the 548 Ciomadul volcano (P < 400 MPa, D < 15 km). Hornb-549 550 lendes crystallized between 200 and 300 MPa and parga-551 sites between 50 and 400 MPa according to the experimental data set. The so-called Ridolfi barometers 552 (R2010, R2012) indicate that low-T hornblendes crystal-553 lized in a shallow magma chamber ($P_{\rm R2010} = 130 \pm 24$ -554 MPa, $D = 5 \pm 1$ km depth) while the high-T pargasites 555 could evolve in a significantly deeper level 556 $(P_{\rm R2010} = 395 \pm 104 \text{ MPa}, D = 15 \pm 4 \text{ km})$ (Fig. 9). 557 The two barometers give fairly similar pressure range for 558

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On the other hand, the Al^{tot} in amphibole barometry (P_{Al} 562 in-am) indicates that the two amphibole group crystallized at 563 564 the same depth interval. For the low-temperature hornblen-565 des, this calculation indicates deeper storage (262 \pm 40 MPa, 566 $D 10 \pm 2$ km) than the R(2010) and R2012 barometry. In the case of the high-T pargasites, the calculated pressure is lower 567 568 $(290 \pm 48 \text{ MPa}, D \ 11 \pm 2 \text{ km})$ than the values that were 569 calculated by the R2010 and R2012 equations. The Kra-570 wczynski et al. (2012) barometry yields similar pressure for pargasites (306 \pm 124 MPa)—using the oxygen fugacity 571 572 values calculated by the R2012-that is partially overlapping with the results of Altot in amphibole barometry while 573 574 yielding larger pressure range.

575 Amphibole $p-T-H_2Om-fO_2$ profiles

576 The R2010 and R2012 equations require only the composition 577 of amphibole and thus enable $p-T-H_2Om-fO_2$ profiles to be 578 calculated in zoned amphiboles. We selected representative 579 crystals having the typical zoning patters of the Ciomadul amphiboles to check the intracrystalline variation of the 580 581 intensive parameters. The R2010 profiles show significant 582 variations in temperature and pressure. In the oscillatory-zoned 583 amphiboles, the temperature variation is about 60-80 °C, 584 while the pressure varies by 100-400 MPa in single amphi-585 boles. The H_2Om and fO_2 profiles also show fluctuations 586 principally in the amphiboles having type B2 oscillatory zon-587 ing. The fluctuations in the intensive parameters clearly mirror the Al^{tot} and Mg-number variations along the profiles (Fig. 11, 588 589 lines 2 and 3). In the simple-zoned amphiboles, a strong and 590 sharp increase in temperature (by ~ 150 °C) and pressure (by 591 more than 200 MPa) was observed in the rim, but H2Om and 592 fO₂ remain constant. The R2012 equations also show fluctu-593 ations in the estimated parameters but the interpretation of 594 these changes is often controversial (e.g., high fO_2 in the low 595 Mg-number zone). Hence, these profiles were not used during 596 the interpretation. The composite crystals show the largest 597 intracrystalline variation in the calculated parameters. The p-T 598 values of a single composite crystal cover the whole range that 599 was calculated for the Ciomadul amphibole phenocrysts 600 (Fig. 11, line 6).

601 Discussion

602 The significance of Al^{IV} versus Mg# interrelation

The large compositional variation in the Ciomadul
amphiboles covers the range typical for intermediate
magmas. However, unusually in our case, this large

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Fig. 7 Compositional profiles— $(Na + K)^A$, Al^{VI} , Ti, Mg, and Al^{IV} —of the studied amphiboles, which are shown in Fig. 6. Lines 2, 3, 5, and 6 are core-to-rim profiles, and line 4 is a rim-to-rim traverse. *Line1* the compositions of the patches in the core (Fig. 6a) are plotted as separate points, and the compositions of the Al^{IV} -richest and Al^{IV} -poorest points are connected with dotted vertical lines. *Dashed vertical lines* indicate dissolution surfaces. *Line5* the zone boundary of the simple-zoned amphibole (Fig. 6f) is affected by diffusion (d.z.). The typical error is indicated by the *symbol size*

variability can be detected even in a single crystal (com-606 posite amphiboles; Fig. 12). Large chemical variation of 607 amphiboles is often difficult to interpret (De Angelis et al. 608 2013), because the amphibole composition responds sen-609 sitively to changes in many thermodynamic parameters, 610 such as melt composition, pressure, temperature, redox 611 state, and volatile content of the magma and the co-crys-612 tallizing mineral phases (e.g., Johnson and Rutherford 613 1989a; Schmidt 1992; Anderson and Smith 1995; Ernst and 614 Liu 1998; Scaillet and Evans 1999; Bachmann and Dungan 615 2002; Pichavant et al. 2002; Rutherford and Devine 2003; 616 Sato et al. 2005; Krawczynski et al. 2012). Consequently, 617 the measured compositional variation is a net result of the 618 combination of these parameters. Thus, it is important to 619 determine which parameters are responsible for the 620 observed inter- and intracrystalline compositional varia-621 tions to avoid misinterpretations of the thermobarometric 622 results (Shane and Smith 2013). The substitution analysis AQ3 23 (i.e., studying the correlation of Al^{IV} and Ti, $A^{(Na+K)}$, Al^{VI} . 624 Ca, etc.) is the most common way to determine the main 625 (intensive) parameters that could control the amphibole 626 crystallization (e.g., Bachmann et al. 2002, Humphreys 627 et al. 2006, De Angelis 2013). Our findings indicate that 628 although this step should not be ignored, a better inter-629 pretation can be achieved, when it is completed by the 630 analyses of Al^{IV} versus Mg# interrelation and the varia-631 tions of trace element contents. 632

The changes in the conditions during the magma evolution 633 are reflected principally in the variation in the tetrahedral 634 alumina (Al^{IV}) value and the Mg number of amphiboles (e.g., 635 Bachmann and Dungan 2002; Rutherford and Devine 2008; 636 Krawczynski et al. 2012), suggesting that these two variables 637 can be useful to recognize and distinguish the main processes 638 that control the growing amphibole composition. The 639 amphiboles of the Ciomadul dacite define two main contin-640 uous compositional trends in the Al^{IV} versus Mg-number 641 diagram (Fig. 8a). The inverse variation of Al^{IV} and Mg 642 number can be interpreted as coupled substitutions according 643 to the substitution equations that were presented for amphi-644 boles by Vyhnal et al. (1991) and Almeev et al. (2002). These 645 negative compositional trends as well as the positive corre-646 lation between Al^{IV} and Ti, $(Na + K)^A$, Al^{VI} indicate the role 647 of edenite $[Si^{IV} + ()^A = Al^{IV} + (Na + K)^A]$, Ti-tschermak $[Si^{IV} + Mg^{VI} = Al^{IV} + Ti^{VI}]$, and Al-tschermak $[Si^{IV} + Mg^{VI} = Al^{IV} + Ti^{VI}]$ 648 649

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Fig. 8 Plots of Al^{IV} versus Mg# (Mg/(Mg + Fe²⁺)) (**a**), (Na + K)^A (**b**), Ti (**c**), and Al^{VI} (**d**) illustrating the substitution mechanisms in the Ciomadul amphiboles. The *exchange vectors* are indicated in the **b**, **c**, and **d** diagrams. **a** Trends 1, 2, 3 show intracrystalliny Al^{IV} -Mg# systematics of zoned amphiboles. All plotted values are in apfu



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 $Mg^{VI} = Al^{IV} + Al^{VI}$ substitutions (Fig. 8) as it has been 650 described also in other localities (e.g., Bachmann and Dungan 651 2002; Rutherford and Devine 2003; Sato et al. 2005; Ruth-652 653 erford and Devine 2008; Humphreys et al. 2009a, b). These 654 substitutions are sensitive primarily to the crystallization 655 temperature and pressure (e.g., Johnson and Rutherford 656 1989a; Blundy and Holland 1990; Rutherford and Devine 657 2003). The clear separation of the two trends (indicated by "trend 3") suggests that the crystallization of amphiboles 658 659 may have occurred in two different magmas. According to the 660 experimental data, the Al/Si ratio of amphibole depends on 661 the Al/Si ratio of the melt (Sisson and Grove 1993; Pichavant 662 et al. 2002). Increasing silica activity in the melt increases the silica content of the coexisting amphibole following the 663 equation pargasite $+ 4SiO_2 =$ hornblende + albite (Blun-664 dy and Holland 1990). As a consequence, the alumina content 665 in the tetrahedral site decreases and therefore the Al^{IV} content 666 of the amphiboles could be viewed as a relative 667

differentiation index too (Ridolfi et al. 2010). This can be 668 refined by plotting experimentally produced amphibole 669 compositions on the Al^{IV} versus Mg diagram. Figure 10 670 shows that low-T amphiboles crystallized from dacitic 671 magmas have low-Al^{IV} and a low-Mg number (blue field). In 672 contrast, amphiboles formed from more mafic magmas at 673 high temperature tend to have higher Al^{IV} and higher Mg 674 numbers. As the silica content of the parent magma decreases. 675 the Al^{IV} values of the amphiboles increase (yellow field in 676 Fig. 10). Mafic magma replenishment into a silicic reservoir 677 could cause abrupt decrease in silica activity in the mixed 678 magma. Moreover, mafic input will also increase the Mg/ 679 (Mg + Fe) ratio in the mixed magma; thus, concomitant 680 increase in Al^{IV} and Mg# of amphibole will be expected due 681 to magma replenishment. The trace element composition of 682 the studied amphiboles appears to support the mixing model 683 since the low-Al hornblendes have much lower Sr, Ba, Eu, 684 and Zr contents than the high-Al pargasites. As the trace 685

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Fig. 10 Plot of Al^{IV} versus Mg# showing the Al^{IV}-Mg# systematics of Ciomadul and experimental amphiboles. Symbols (for Ciomadul amphiboles) as in Fig. 4. The black arrows at the upper left edge indicate the interpretation of Al^{IV}-Mg# systematics of amphiboles (for details, see text). Experimental results were filtered by the method of Ridolfi et al. (2010). High- and low-temperature experimental amphiboles can be separated in the plot. The starting materials used in the experiments are also indicated (D: dacite, A: andesite, BA: basaltic andesite, B: basalt). Results of the following experiments are plotted: Rutherford and Devine (1988): Johnson and Rutherford (1989a, b); Schmidt (1992); Martel et al. (1999); Sato et al. (1999); Scaillet and Evans (1999); Pichavant et al. (2002); Grove et al. (2003); Rutherford and Devine (2003); Barclay and Carmichael (2004); Costa et al. (2004); Holtz et al. (2005); Sato et al. (2005); Larsen (2006); Adam et al. (2007); Rutherford and Devine (2008); Simakin et al. (2009); Krawczynski et al. (2012); Simakin et al. (2012)

686 element content is primarily influenced by the composition of the coexisting melt, the observed differences indicate that the 687 688 two amphibole populations in the Ciomadul dacites were formed in different magmas. Similarly, positive Al^{IV}-Mg# 689 trends in amphiboles were reported at other volcanoes (e.g., 690 691 Soufriere Hills, Humphreys et al. 2009a; Unzen, Sato et al. 692 2005) where compositional mixing was detected (Nakamura 693 1995; Murphy et al. 2000).

694 Interpretation of amphibole thermobarometry

695 Application of various amphibole thermometers indicates 696 that the hornblendes in the Ciomadul dacite crystallized at 697 much lower temperature than the pargasites, but they yield 698 different temperature ranges for each population. According 699 to the amphibole-plagioclase thermometry, the hornblendes 700 crystallized at almost 100 °C lower temperature than it is 701 indicated by the R2010 thermometer. This difference is 702 notable even if we consider the relatively large error of the thermometers ($T_{\text{am-plag}}$ error = 40 °C according to Holland 703 704 and Blundy (1994), and T_{R2010} error = 22 °C according to Ridolfi et al. (2010)). Experimental works on the dacitic rocks 705

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of the Fish Canvon Tuff resulted in amphiboles and coexis-706 707 ting mineral assemblage similar to the Ciomadul rocks at low temperatures (\sim 740 °C; Johnson and Rutherford 1989b). 708 Thus, the R2010 thermometry appears to overestimate the 709 crystallization temperature of the Ciomadul hornblendes but 710 the R2012 yields more appropriate results. Consequently, the 711 712 hornblendes crystallized below 800 °C may be at \sim 730 °C as indicated by amphibole-plagioclase thermometry. For the 713 high-Al amphiboles (pargasites), the amphibole-plagioclase 714 thermometer gives consistently lower crystallization tem-715 716 perature than the R2010 and R2012. This is corroborated by the comparison with the experimental data set (Fig. 10) and is 717 in agreement with the suggestion of Blundy and Cashman 718 (2008) who claimed that this thermometry is not so accurate 719 for amphiboles with such high-Mg number as shown by the 720 Ciomadul pargasites. On the other hand, the results of the 721 R2010 and R2012 thermometers are in agreement with the 722 723 experimental data and most likely indicate the crystallization temperature (940-980 °C) of the pargasites. Thus, careful 724 evaluation of the thermometric results indicates that the 725 crystallization of the hornblendes and pargasites in the 726 Ciomadul magma could occur at significantly different 727 (about $\Delta T = 200$ °C) temperatures. Additionally, the Ridolfi 728 et al. (2010) thermometer likely overestimates the crystalli-729 zation temperature of low-Al (cold) amphiboles. 730

The barometric calculations can be used to infer the pre-731 eruptive magma chamber architecture beneath the Ciom-732 733 adul volcano. Both the R2010 and R2012 barometry imply a vertically extended magma plumbing system in which the 734 two amphibole populations could crystallize at two sepa-735 rated magma storage levels. According to these barome-736 ters, the hornblendes were crystallized at shallow depth 737 $(\sim 5 \text{ km})$, whereas the pargasites were formed at much 738 greater depth, in the middle crust (~ 15 km). The Al^{IV} 739 versus Al^{VI} correlation in the two amphibole populations 740 indicates the role of the pressure-sensitive Al-tschermak 741 substitution during their crystallization and seemingly 742 support these barometric results. However, the crystal 743 growth stratigraphy and the inferred growth history indi-744 745 cate that the crystallization of the high-Al pargasites could not occur deeper than the low-Al hornblendes. Namely, the 746 calculated pressure profiles of simple-zoned and composite 747 amphiboles using the R2010 and R2012 barometer indicate 748 an abrupt increase in the crystallization pressure at the 749 750 core-rim boundary. This would mean that the rim of these 751 amphiboles would crystallize about 8-10 km deeper than the cores, and therefore, the low-Al hornblendes should 752 sink many kms in the magma reservoir not long before the 753 754 eruption. A similar model (i.e., circulating crystals) was suggested for the origin of the cyclically zoned amphiboles 755 in the Mt. St. Helens dacites, where a vertically elongated, 756 large magma chamber is hypothesized (Pallister et al. 757 758 2008). However, no such simple-zoned and composite

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Fig. 11 Temperature (T), pressure (p), melt H_2O content (H_2O_m), oxygen fugacity (fO_2), and compositional (Al^{tot} , Mg# [apfu]) profiles of three selected amphiboles shown in Fig. 6c, d and h, respectively.

Vertical lines indicate dissolution surfaces. The intensive parameters were estimated based on the Ridolfi et al. (2010) formulation

759 amphibole crystals were described in the Mt. St. Helens 760 rocks as we found here in the Ciomadul dacite. Moreover, in our system, the R2010 and R2012 barometry would 761 762 indicate two separated magma chamber for the bimodal 763 amphiboles, at least 5 km vertical transportations of crys-764 tals seems to be highly unrealistic scenario and rather 765 implies that these barometers significantly overestimate the 766 crystallization pressure of the high-Al pargasites and/or 767 underestimate the crystallization pressure of low-Al 768 hornblendes. Interestingly, the Al-in-amphibole barometry 769 (Schmidt 1992; Anderson and Smith 1995) yields appro-770 priate results (i.e., overlapping pressure values for both 771 hornblendes and pargasites) in spite of the lack of the 772 required coexisting mineral assemblage in the case of the 773 pargasites. Additionally, Grove et al. (2003) and Kra-774 wczynski et al. (2012) proposed that the Mg# of the primitive (i.e., near liquidus) amphiboles is related to the 775 P_{H2O}, and therefore, their hygro-barometer can be used for 776 777 the pargasites. Although we do not have direct evidence of H₂O saturation in the Ciomadul magmas, the lack of neg-778 ative Eu anomaly in the pargasites probably suggests that 779 they are originated from a melt that was too H2O-rich to 780 crystallize plagioclase. Nevertheless, the overlapping result 781 of the PAl-in-am calculations and the Krawczynski's 782 barometry is notable and likely informative and is in 783 agreement with the crystal growth stratigraphy. All of these 784 suggests that the depth of the main (eruption feeder) 785 magma storage is indicated by the range of the overlapping 786 barometric results of hornblendes and pargasites that is 787 approximately 200-300 MPa, 8-12 km. An important 788 point is that the simple-zoned and composite crystals pro-789 790 vide direct evidence that the significant variation in Al

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Fig. 12 Compositional variation of amphiboles from selected volcanoes. It is important to note that a single composite crystal of the Ciomadul dacite overlaps almost the entire Al_2O_3 range of the other volcanoes. Amphibole compositions are from the following references: Mt. St. Helens, Thornber et al. (2008); Mt. Pinatubo, Pallister et al. (1996); Soufriére Hills (SHV), Murphy et al. (2000); Redoubt, Coombs et al. (2012); Unzen, Sato et al. (2005); Mt Pelée, Pichavant et al. (2002) Ciomadul, this study

791 content in amphiboles cannot be always explained by 792 crystallization at different pressures/depths even if the Al-793 tschermak substitution can be recognized. Furthermore, the 794 low-Al hornblendes-crystallized well inside their stability 795 field-of the Ciomadul dacite also demonstrate that 796 amphiboles do not necessarily crystallize along their sta-797 bility curve in a narrow stability field as indicated by the 798 thermobarometric formula of Ridolfi et al. (2010). Thus, it 799 can yield unrealistic estimations for the crystallization 800 conditions of crystal mush-derived amphiboles and high-Al 801 amphiboles derived from mafic hybrid magmas.

802 The origin of amphiboles in the Ciomadul dacites

Bespite the observed diversity in zoning types and the petrographic occurrences of the amphiboles, they can be divided
into two coherent compositional groups (Fig. 10). The origin
of the two amphibole populations is constrained based on
their major and trace element compositions, the coexisting
mineral phases, and the thermobarometric results.

Hornblendes in the Ciomadul dacites are characterized 809 typically by low-Al^{IV} and low-crystallization temperature, 810 811 suggesting that they are derived from a cold silicic magma. 812 This is supported by the petrographic observations, because 813 hornblendes are coexistent with a mineral assemblage 814 involving quartz, biotite, K-feldspar, plagioclase, titanite, 815 apatite, and zircon common to silicic magmas. Inter-816 growths with these minerals (Fig. 2a) unambiguously 817 indicate their common origin. The trace element content of 818 hornblendes shows depletion in Ba, Sr, Zr, and Eu (Fig. 5).

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These trace elements behave nearly compatible in amphi-819 boles (Viccaro et al. 2007 and references therein); thus, 820 their relative depletion can be explained by co-crystalli-821 zation of minerals such as plagioclase, K-feldspar, titanite, 822 and zircon. The observed mineral assemblage is often form 823 824 felsic crystal clots in the studied dacites with up to a few cm size. The texture of these felsic clots resembles plutonic 825 rock textures, suggesting that they could be derived from a 826 highly crystalline silicic magma body, where the consti-827 tuting minerals are usually joined by solid-solid contacts. 828 829 Nevertheless, occurrence of thin interstitial rhyolitic glass implies that this magma body still contained some melt 830 portions, i.e., it could have been a hornblende-bearing 831 mush/sponge residing at shallow 832 crystal depth $(\sim 8-12 \text{ km})$ as shown by the geobarometric calculations. 833

The high-Al pargasites is characterized by high-Al^{IV} 834 content and high crystallization temperature (>900 °C), 835 suggesting that they could crystallize in a hotter, more 836 mafic magma. This is corroborated by the petrographic 837 observations, since pargasites often coexist with Fo-rich 838 olivine and Mg-rich clinopyroxenes (Fig. 2c, d). Further-839 more, they have a distinct trace element content with no 840 depletion in the components (Ba, Sr, Zr, Eu) seen in the 841 842 hornblendes and resemble the trace element composition of the amphiboles found in the mafic volcanic products of 843 Etna (Viccaro et al. 2007). The origin of such high-Al 844 amphiboles is often interpreted as forming in a deep mafic 845 source (e.g., Grove et al. 2003, 2005; Ridolfi et al. 2010; 846 Krawczinsky et al. 2012) and is then transferred and 847 incorporated into a more differentiated magma via mafic 848 replenishment, partly by disaggregation of the mafic 849 enclaves (Humphreys et al. 2009a). This commonly cited 850 scenario, however, does not work for the Ciomadul dacite 851 852 as it was discussed in the former section. In fact, there are a few studies, which suggest that high-Al amphiboles can 853 crystallize also at low pressures in upper crustal magma 854 storage or even in the conduit (Sato et al. 1999; Coombs 855 et al. 2013) and this condition could be valid for the origin 856 of the Ciomadul amphiboles, as well. The low-pressure 857 858 formation of the Ciomadul pargasite is supported by further petrographic observation, such as their occurrence around 859 reacted Mg-rich olivine crystals. Coombs and Gardner 860 (2004) showed that a reaction rim is formed around mag-861 nesian olivine in contact with silicic magmas. Pargasites 862 are grown on this reaction rim implying that the olivines 863 were already in the silicic magma and pargasites crystal-864 lized from a hybrid melt after the mixing of mafic and 865 silicic magmas. Although experimental studies produced 866 high-Al amphiboles such as our pargasites at low pressure 867 (e.g., Sato et al. 1999, Browne 2005, Barclay and Carmi-868 chael 2004), direct evidence for low-pressure crystalliza-869 tion of high-Al amphiboles is rarely presented in natural 870 rock samples (Sato et al. 2005). 871

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872 Magma chamber processes

873 Formation of simple zoning and the composite amphiboles

874 Thermometric calculation of the simple-zoned and com-875 posite amphiboles from the Ciomadul dacite implies a rim ward temperature rise at least with 150 °C, but it could be 876 as high as 200 °C if the results of the Holland and Blundy 877 878 (1994) amphibole-plagioclase thermometry are used for 879 the low-Al hornblendes (Fig. 11, line 6). This is consistent 880 with the experimental data shown in Fig. 10. Such a 881 major increase in temperature can be explained only by intrusion of basaltic magma into the shallow, cold, silicic 882 883 magma chamber. The strong reheating is reflected also in 884 the complex zoned, composite amphiboles (Fig. 6e, f). At 885 the core-rim boundary, a narrow zone composed of 886 clinopyroxenes is observed that is interpreted as the 887 remnant of the reaction rim around the low-Al amphibole 888 formed due to thermal breakdown. This could mean that 889 the low-Al amphiboles came out of the stability field due 890 to the reheating of the silicic crystal mush body. Experi-891 mental data on amphibole stability suggest that amphi-892 boles in dacitic magmas are stable up to about 910 °C 893 (e.g., Rutherford and Devine 2008) that is approximately 894 150 °C higher than the calculated crystallization temper-895 ature of the hornblendes. This abrupt temperature increase as well as change in the Mg number toward the rim is 896 897 consistent with the injection of fresh mafic magma before 898 eruption. Isobaric reheating experiments of Browne 899 (2005) on dacitic magma produced a breakdown rim around amphiboles with minor pargasite overgrowth. 900 901 However, the extensive crystallization of the high-Al 902 pargasites in the dacitic magma of Ciomadul needs 903 additional mafic components during reheating. This sug-904 gests that mafic magma not only reheated the silicic res-905 ident magma but hybridization could have also occurred 906 (e.g., Costa and Singer 2002; Rutherford and Devine 907 2003). A hybridization zone at the boundary of silicic and 908 mafic magmas is also likely due to the observation that 909 pargasites are present around hornblendes from the silicic 910 magma and also around olivine crystals transported by the 911 intruded mafic magma. Another explanation of this com-912 positional shift could be a concordant increase in fO_2 and 913 temperature, since experiments on dacitic magmas indi-914 cate that this change in redox state positively influences 915 the amphibole Mg number and temperature also positively influences the Al^{tot} content (e.g., Scaillet and Evans 1999; 916 917 Costa et al. 2004). According to Scaillet and Evans (1999), a temperature increase from 780 to 900 °C indu-918 919 ces an increase of ~ 0.4 apfu in Al^{tot} at isobaric conditions. However, in the Ciomadul amphiboles, this 920 921 variation in Al^{tot} is almost the double (~ 0.78) from core 922 to rim, thus this cannot be produced only by a temperature increase. Furthermore, Costa et al. (2004) reported that 923 high T and low fO₂ promote high-Al content in the 924 experimentally produced amphiboles instead of high T 925 and high fO_2 that needs to increase both Al^{tot} and the Mg 926 number. Thus, it is not likely that the observed core-to-927 928 rim compositional variation can be produced without the 929 involvement of mafic magma replenishment and magma mixing. According to the experimental results of Kra-930 wczynski et al. (2012), the high-Al^{tot} content in pargasites 931 suggests that the mafic magma could be hot and also 932 water-rich. The mafic replenishment is also supported by 933 934 the presence of magnesian olivine and clinopyroxene crystals in the erupted dacitic magma. 935

Origin of the cyclic zoning in amphiboles 936

The Ciomadul dacite contains two types of oscillatory-937 zoned amphiboles: type B1 and type B2. Both types show a 938 negative trend on the Al^{IV}-Mg number diagram (Figs. 8a, 939 10) indicating that their crystallization was influenced by 940 the fluctuation in intensive thermodynamic parameters. The 941 variation in the Al content in both zoning types is primarily 942 controlled by the temperature-sensitive edenite and Ti-943 944 tschermak substitutions. Temperature profiles of the 945 oscillatory-zoned amphiboles suggest 60-80 °C variation (Fig. 11, line 2, 3). This temperature change could be 946 947 explained by convective motion of the magma in the magma chamber due to the temperature contrast (Couch 948 et al. 2001). Oscillatory-zoned amphiboles, which experi-949 ence convection, may also show Al-tschermak substitution 950 951 as was demonstrated by Rutherford and Devine (2008). However, in our case, only type B2 crystals show Al-952 tschermak substitution, suggesting that type B1 crystals 953 954 require another explanation. Input/degassing of volatiles 955 could be an alternative model for the formation of oscil-956 latory zoning as was indicated by Humphreys et al. (2009b) and Sato et al. (2005). Experimental data on the Pinatubo 957 dacite show that the H₂O content of the melt can affect the 958 Al^{tot} content of the amphibole (Scaillet and Evans 1999). 959 Furthermore, fluid input/degassing may also affect the 960 961 redox state of the magma, which influences the amphibole composition as well. Hence, changes in the volatile budget 962 can indeed be reflected in the amphibole composition. 963 Calculation of H_2 Omelt and fO_2 values along profiles in 964 cyclic-zoned amphibole crystals was performed using Ri-965 dolfi et al.'s (2010) equations. These variables show minor 966 fluctuations and change at the dissolution surfaces, sug-967 gesting that volatiles might play a role in addition to the 968 temperature change in the formation of type B1 oscillatory 969 970 zoning (Fig. 11, line 2). Heat can be transferred by volatiles; thus, input of hot fluids into the magma reservoir may 971 be responsible for the temperature variation and dissolution 972 973 of the growth zones.



974 Origin of patchy zoning in amphibole

975 Patchy zoning can form due to open-system processes or 976 diffusional chemical re-equilibration of a zoned crystal 977 (Streck 2008). In the former case, disequilibrium condi-978 tions initiate a spongy-like dissolution of the mineral and parallel or after this, crystallization can occur with a 979 980 composition reflecting changing conditions. This type of 981 patchy zoning is commonly observed in plagioclase, 982 clinopyroxene, or amphibole (Streck 2008). The most 983 important feature that could help us to identify this texture 984 is the sharpness of the compositional transition between 985 patches and the enclosed mineral or melt inclusions at the 986 boundary of the patches due to rapid re-growth. According 987 to these criteria, both types of the patchy-zoned amphiboles 988 in the Ciomadul dacite could be formed during open-sys-989 tem processes. Type A1 crystals show similar chemical 990 features as the type B1 oscillatory-zoned amphiboles. 991 Additionally, this type of patchy-zoned amphibole often 992 has a type B1 oscillatory rim. Hence, these two zoning 993 types might have been formed due to the same petrogenetic 994 processes. Type A2 patchy-zoned amphiboles show similar 995 chemical variability as the simple-zoned amphiboles sug-996 gesting their common origin. Based on these similarities, 997 we can conclude that similar processes, i.e., mafic magma 998 replenishment, could lead to the formation of the (Type 999 A2) patchy zoning.

1000 Processes and rates of amphibole breakdown

1001 The amphiboles in the Ciomadul dacite show diverse 1002 breakdown textures, suggesting that they were caused by a 1003 combination of processes. Similar diversity of amphibole 1004 breakdown textures was also found in the Soufriére Hills 1005 andesite, where decompression, heating, and late stage oxidation were all referred to as potential mechanisms 1006 1007 leading to amphibole breakdown (Rutherford and Devine 2003). 1008

1009 Hornblendes with opacitic breakdown textures in the 1010 Ciomadul dacite are often surrounded by high-temperature pargasitic rims (Fig. 2h). Additionally, coarse-grained 1011 1012 clinopyroxene rims around hornblendes were also observed 1013 (Fig. 2g) providing important evidence of reheating (Ruth-1014 erford and Devine 2003). The breakdown rim of high-Al 1015 pargasites is only developed where these crystals are in direct 1016 contact with the matrix, i.e., it should have formed by crystal-1017 melt reaction (Fig. 2e), suggesting that they were formed due 1018 to decompression and degassing during magma ascent to the 1019 surface (Rutherford and Hill 1993).

Experimental works indicate that the extent of amphibole breakdown is a function of the time that the amphibole
spends outside its stability field (Rutherford and Hill
1993),although it depends also on the ambient conditions

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(Rutherford and Devine 2003: Browne and Gardner 2006). 1024 1025 We used the rim thickness data to estimate the time of heating and ascent. According to the different calibrations 1026 of the rim formation rate, the pargasites spent a maximum 1027 of approximately 12 days outside their stability field. As 1028 1029 the rim of pargasites most probably formed via decompression, this estimate may correspond to the magma 1030 ascent rate from the magma chamber. The dacites con-1031 taining hornblendes with the thinnest rims ($\sim 15 \ \mu m$) may 1032 1033 suggest the timescale of pre-eruption heating and magma 1034 mixing. Unfortunately, very little experimental data exist about the rate of thermal breakdown of amphiboles. The 1035 available amphibole breakdown rates were calibrated for 1036 decompression and suggesting $\sim 6-22$ days for hornblende 1037 breakdown. As the thermal breakdown rate could be much 1038 1039 faster Browne (2005), the reaction rim of the Ciomadul hornblendes might have been formed only within a few 1040 days, suggesting that the magma chamber was reheated just 1041 before the eruption. 1042

Implications for the volcanic plumbing systems beneath1043intermediate arc volcanoes1044

Petrological mapping of the magmatic plumbing system 1045 beneath active and dormant volcanoes provides constraints 1046 on eruption style and aids interpretation of geophysical 1047 signals. This information is an essential component of 1048 1049 volcanic hazard assessment. The composition, especially the alumina content of amphibole phenocrysts, often shows 1050 large variation in the erupted intermediate magmas even in 1051 1052 the thin-section scale, and this is commonly interpreted as the result of amphibole crystallization at different pressures 1053 and temperatures (e.g., Pichavant et al. 2002; Thornber 1054 et al. 2008; Ridolfi et al. 2010; Chambefort et al. 2013; 1055 Walker et al. 2013; Costa et al. 2013; and Turner et al. 1056 2013). Results of such studies imply that these volcanoes 1057 are underlined by a vertically extended magma storage 1058 zone, where the different amphibole populations crystallize 1059 at different depths. This zone may consist of separated 1060 1061 magma chambers (e.g., Soufriere Hills and Bezymianny, Turner et al. 2013) or a larger, vertically extended magma 1062 reservoir (e.g., Mt. Pelée, Pichavant et al. 2002; Santiag-1063 uito, Scott et al. 2012; Mt. St. Helens, Pallister et al. 2008). 1064 However, in the Ciomadul, dacite single composite crystals 1065 1066 show almost the same variation in the Al content as the amphibole crystals in the entirety of other volcanic systems 1067 (Fig. 12). Such crystals have a major importance as they 1068 verify that if the composition of amphiboles is controlled 1069 by temperature and magma composition at constant pres-1070 sure (as in our rock samples), the variation in the crystal-1071 lization pressure/depth detected with amphiboles can be 1072 only apparent. This finding highlights that amphibole 1073 barometers that consider only amphibole composition such 1074

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Fig. 13 a Comparison of amphibole populations of mixed intermediate magmas from different volcanic arcs defines two separate groups: 1. cold amphiboles (cold am) derive from a silicic crystal mush, and hot amphiboles (hot am) originate from a hybrid melt that was formed due to fresh mafic intrusion into the mush and mixing of the mafic and silicic magmas, thus they are mafic recharge-related amphiboles. **b** Crystallization conditions of cold and hot amphiboles according to the Ridolfi et al. (2010) thermobarometer. Data of Mt. Pinatubo (PTB): Pallister et al. (1996); Soufriére Hills (SHV): Murphy et al. (2000); Redoubt (RDT): Coombs et al. (2012); Unzen (UNZ): Sato et al. (2005); Mt Pelée (MPL): Pichavant et al. (2002); Ciomadul (CIOM): this study





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Magma chamber architecture and petrogenesis of mixed intermediate magmas:

1. based on the Ridolfi thermobarometry

2. based on our study



Fig. 14 Schematic cartoon that shows two different models for the inferred magma storage system and petrogenesis of mixed intermediate magmas based on bimodal amphibole populations. As it is

as Ridolfi et al. (2010) may often yield unrealistic pressure 107(AQ4 variation and magma chamber architecture (Fig. 13).

1077 Processes and conditions of magma mixing beneath1078 andesite-dacite composite volcanoes—the amphibole

1079 perspective

1080 Intermediate mixed magmas of volcanic arcs often host 1081 almost the same bimodal amphibole cargo as it was 1082 observed in the Ciomadul dacites. Bimodal amphiboles 1083 characterize the erupted products of, e.g., Mt. Pinatubo 1084 (Pallister et al. 1996), Unzen (Sato et al. 1999, 2005), 1085 Soufriére Hills (Humphreys et al. 2009a), Redoubt (Wolf 1086 and Eichelberger 1997; Coombs et al. 2012), Mt. Pelée 1087 (Pichavant et al. 2002), and the Central Volcanic Zone 1088 dacites in Chile (Nakada 1991). Although the similar 1089 amphibole cargo may suggests similar processes and con-1090 ditions beneath these volcanoes, several models have been 1091 established to explain the origin of the bimodal amphiboles 1092 (e.g., Pichavant et al. 2002; Coombs et al. 2012; Ridolfi 1093 et al. 2010; Sato et al. 1999; Koleszar and Kent 2011). Our 1094 amphibole perspective investigation indicates that bimodal 1095 amphibole populations cannot unambiguously mean crys-1096 tallization at different depths as it is commonly indicated

shown, only the second model provides realistic interpretation for the formation of simple-zoned and composite amphiboles. For further details, see the discussion. The figures are not to scale

1097 (Pichavant et al. 2002; Ridolfi et al. 2010; Koleszar and Kent 2011; Chambefort et al. 2013), but an alternative 1098 model could be the remobilization of a long-lived, near-1099 solidus crystal mush by hot mafic magma intrusion (Na-1100 kamura 1995; Pallister et al. 1996; Murphy et al. 2000). 1101 This suggests that these volcanoes can be underlain by a 1102 shallow (100-300 MPa) storage zone where cold, silicic 1103 magmas are captured and low-Al amphiboles are formed. 1104 High-Al amphiboles can crystallize at shallow depth during 1105 reactivation and remobilization of this crystal mush due to 1106 mafic magma intrusion followed by hybridization of the 1107 different magmas (Fig. 14). The amphibole thermobarom-1108 1109 etry of Ridolfi et al. (2010) and Ridolfi and Renzulli (2012) seems to be useful in the case of mixed magmas in which 1110 equilibrium mineral phases are difficult to find as well as 1111 experiments are limited due to their mixed character. 1112 However, these equations lead to the same p-T evolution 1113 path for all amphiboles along their stability curve; at the 1114 same time, bimodal amphiboles in mixed intermediate 1115 magmas usually not follow this trend, as it is demonstrated 1116 in this study. Thus, these formulations are unable to 1117 reproduce the conditions and p-T path of mafic magma 1118 replenishment into shallow eruption feeder magma storage 1119 filled with cold silicic crystal mush, that is, however, a 1120

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1121 commonly observed process in arc volcanoes (Nakamura

- 1122 1995; Pallister et al. 1996; Murphy et al. 2000; Eichel-
- 1123 berger et al. 2000).

1124 Conclusion

1125 The amphibole perspective of the studied Ciomadul's da-1126 cites suggesting that eruptions of the volcano were pre-1127 ceded by complex magma chamber processes. An 1128 important implication is that the erupted dacitic magma 1129 was formed in an upper crustal magma storage zone where 1130 reheating and partial remelting of silicic crystal mush 1131 occurred due to mafic magma replenishment days or weeks 1132 before the eruption. Our model is in contrast to the former 1133 view that suggested lower crustal conditions for the dacite 1134 genesis including continuous melting of the lower crust by 1135 mafic under platting and mixing-hybridization of mantle-1136 and crustal-derived magmas (Vinkler et al. 2007).

1137 Our study highlights that without the knowledge of the 1138 processes that is responsible for the compositional varia-1139 tion of amphiboles, thermobarometers based on solely their 1140 chemistry can lead misleading conclusions on magma 1141 chamber architecture and conditions of the magma 1142 evolution.

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AQ4	Please check and confirm the inserted citation of Fig. 13 is correct. If not, please suggest an alternative citation. Please note that figures and tables should be cited in sequential order in the text.	
AQ5	Please check and confirm the author names and initials are correct in the references Hammarstrom and Zen (1986).	
AQ6	Please provide page range for reference Harangi et al. (2005).	
AQ7	Please provide complete details for reference Pallister et al. (1996, 2008), Pécskay et al. (1995) and Thornber et al. (2008).	
AQ8	Please provide publisher name and location for reference Rutherford and Devine (2008).	
AQ9	Please provide publisher location for reference Streck (2008).	
AQ10	Please provide editor names for reference Szakacs et al. (2002).	